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# CHARACTERIZATION OF DEFECTS IN METAL SHEETS VIA INFRARED THERMOGRAPHY

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#### **ABSTRACT OF THESIS**

#### CHARACTERIZATION OF DEFECTS IN METAL SHEETS VIA INFRARED THERMOGRAPHY

Defects in Aluminum, Stainless steel and galvanized steel sheets are studied in reflection mode infrared thermography. The effect of material properties, surface finish, heating intensity, heater emission spectra, pixel size and defect size are studied. Contrast is governed by heat quality, emissivity and defect geometry—which follows a logarithmic trend. The diameter detected via infrared thermography is found to be at least 30% larger than the measured diameter and sub-pixel defects can be detected. The use of gradient and Laplacian of temperature is introduced as a means of increasing defect contrast and mitigating heater variation.

KEYWORDS: Infrared Thermography, Non-destructive Testing, Non-destructive Examination, Inspection, Manufacturing

	Marc Harik
_	December 7, 2010

CHARACTERIZATION OF	DEFECTS IN	METAL SHEET	'S VIA INFRARED	) THERMOGRAPHY

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#### **THESIS**

Marc Anthony Harik

The Graduate School
University of Kentucky
2010

#### CHARACTERIZATION OF DEFECTS IN METAL SHEETS VIA INFRARED THERMOGRAPHY

#### **THESIS**

A thesis submitted in partial fulfillment of the requirements for the degree of Master of Science in Mechanical Engineering in the College of Engineering at the University of Kentucky

Ву

Marc Anthony Harik Lexington, Kentucky

Director: Dr. Kozo Saito, Professor of Mechanical Engineering

Lexington, Kentucky

2010

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## TABLE OF CONTENTS

ACKNOWLEDGEMENTS	iii
LIST OF TABLES	vi
LIST OF FIGURES	vii
Chapter One: Introduction	1
i. Review of Nondestructive Technology	1
ii. Background	3
Chapter Two: Experimental Methods	5
i. Mathematical Formulation	5
ii. Apparatus	12
Detector	12
Heaters	13
iii. Sample Preparation	16
iv. Estimation of Uncertainty	20
v. Experimental Procedure	21
Chapter Three: Results and Discussion	22
i. Sample Images	23
ii. Effect of Diameter	28
iii. Detected Diameter	31
iv. Effect of Aspect Ratio	34
v. Apparent Emissivity	35

vi. Effect of Heating and Heaters	37
vii. Derivatives as Filters	39
Chapter Four: Conclusions and Future Work	43
Appendix A: Defect Sizes	45
References	49
Vita	51

# LIST OF TABLES

Table 2.1 Heater settings for the black body and ceramic heater; controlled parameters are in	1
bold	15
Table 2.2 Thermal properties of materials studied	16
Table 2.3 The nominal depth and diameter for each defect plate	17
Table A.1 Defect Measurements for Aluminum Mirror-Like Finish Samples	45
Table A.2 Defect Measurements for Aluminum Satin Finish Samples.	46
Table A.3 Defect Measurements for Steel Mirror-Like Finish Samples.	47
Table A.4 Defect Measurements for Steel Unpolished Finish Samples	47
Table A.5 Defect Measurements for Galvanized Steel Samples	48

## LIST OF FIGURES

Figure 1-1 Reflection and transmission mode schematics
Figure 2-1 Energy balance diagram for a grey body5
Figure 2-2 Schematic of Experimental Setup
Figure 2-3 Infared Salamander Half Trough Emitter Schematic [34]14
Figure 2-4 The defect plate arrangement
Figure 2-5 Circle Fitted to points chosen along the edge of the defect
Figure 2-6 a) Microscopic image of defect in mirror finish steel sheet b) Processed black and
white image c) Processed image with surroundings removed and holes filled in d) Errant pixels
removed and morphological operations performed19
Figure 3-1 Array 3 of mirror-like finish Aluminum with a) black body heater at 373 K b) ceramic
heater at 1093 W/m <sup>2</sup> 23
Figure 3-2 Array 3 of satin finish Aluminum with a) black body heater at 473 K b) ceramic heater
at 1093 W/m <sup>2</sup> 24
Figure 3-3 Array 2 of mirror-like finish stainless steel with a) black body heater at 303 K b)
ceramic heater at 478 W/m²25
Figure 3-4 Array 3 of unpolished stainless steel with a) black body heater at 323 K b) ceramic
heater at 615 W/m <sup>2</sup> 26
Figure 3-5 Temperature maps of Array 3 of galvanized steel with a) black body heater at 573 K b)
ceramic heater at 6831 W/m <sup>2</sup>
Figure 3-6 Diameter divided by pixel size plotted against contrast for array 3 with the black body
heater at 373 K
Figure 3-7 Effect of diameter for the array 3 with the black body heater at 373 K a) Low
emissivity samples b) High emissivity samples29

Figure 3-8 Contrast of a defect with D/px of 6	0
Figure 3-9 Detected diameter divided by the measured diameter plotted against the measured	
diameter over pixel size. Ceramic heater set to 478 W/m23	1
Figure 3-10 Detected diameter divided by the measured diameter plotted against the measured	
diameter over pixel size. Black body heater set to 573 K	2
Figure 3-11 Measured diameter divided by the detected diameter and plotted against the	
measured diameter over pixel size for all array 3 tests	3
Figure 3-12 The effect of aspect ratio on contrast using data from array 1 with the black body	
heater at 373 K	4
Figure 3-13 Apparent emissivity of a cavity opening for a cylindrical cavity of finite length with	
diffuse reflecting walls at a constant temperature. From Robert Siegel, John R. Howell; Thermal	
Radiation Heat transfer, 4th Edition [37]3!	5
Figure 3-14 Apparent emissivity for the array 1 defects in the aluminum samples30	6
Figure 3-15 The effect of heating temperature on contrast by heater type on Aluminum samples	;
with D ≈ 2mm	7
Figure 3-16 Effect of heat quality on contrast with respect to heater type38	8
Figure 3-17 Aluminum Mirror Array 3 with ceramic heater at 2732 W/m <sup>2</sup> a) thermogram b)	
magnitude of the gradient of the thermogram c) Laplacian of the thermogram40	0
Figure 3-18 Contrast maps for a) gradient of temperature b) Laplacian of temperature4	1
Figure 3-19 Average contrast by material for all defects, heating conditions and arrays4	2

**Chapter One: Introduction** 

#### i. Review of Nondestructive Technology

Nondestructive testing is often an essential step in ensuring product quality economically as safety factors are reduced and the demand on machines increases. Testing nondestructively transitioned from a laboratory operation to a valuable production tool in the 1920's [1]. The man-power shortage caused by the First World War, demonstrated the need for automated and fool-proof methods of inspection that did not rely on operator judgment [2]. To reduce defects, an inspection must provide feedback to processes upstream when a defect is discovered; increasing the number of off-line post-mortem judgment inspections can only increase the reliability of inspection [3].

Manual visual inspection is one of the easiest forms of inspection to implement and is the most common [4], but it is not necessarily the most reliable or cost-effective. Machine assisted vision systems may aid inspectors—however, they are subjective, require highly trained personal and relatively slow production speeds [5]. Automated mechanical-optical systems, eddy current detection, ultrasonic testing and infrared thermography all seek to provide more reliable inspection methods and detect a wide range of defects.

Automated optical inspection systems are fast and can inspect as large area, but they cannot detected subsurface defects in metal sheets [6-8]. Eddy current detection is fast and can detect micron scale defects on the surface and in-depth, but the scanning area is relatively small, and is used in the aircraft and space industry [5, 9, 10]. Ultrasonic inspection and acoustic emissions can be used to determine grains size, inspect welds and determine elasticity, but requires contact and in the case of metal sheets, a large sample [11-13]. Infrared thermography

can be implemented on line in both active and passive sensing modes. It has been used to detect defects in the steel making process, marine coatings [14] and painted surfaces in automotive applications [15].

"Quality can be assured reasonable only when it is built in at the process and when inspection provides immediate and accurate feedback at the source of defects." -Shigeo Shingo [3]. In reaching this goal, infrared thermography allows for high speed inspection surface and sub-surface inspection of large areas and may be useful in determining root cause.

#### ii. Background

The infrared spectrum is normally split into three different ranges for inspection; short or near wave (.75–3  $\mu$ m), middle wave (3–6  $\mu$ m) and long or far (8–15  $\mu$ m) [4, 16]. There are two main types of thermal sensors used in focal plane arrays, photonic and micro-bolometer. Micro-bolometers are sensitive over a wide range but have a longer response time than photonic detectors [4]. Photonic detectors have a very fast response time with a limited spectral response and usually require cooling [4, 16].

Infrared sensing is either passive or active. In the active mode, the sample is heated or stimulated for the purposes of inspection. In passive mode, the emissions from the sample and reflections from its surroundings are studied without a stimulus provided for the purpose of inspection [1]. Active sensing can generally be performed in either reflection or transmission or reflection mode as shown in Figure 1-1.

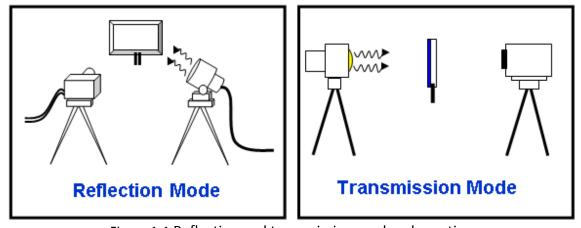


Figure 1-1 Reflection and transmission mode schematics.

Reflection mode thermography may be used to detect surface and subsurface defects, where transmission mode may detect defect throughout the sample. Reflection mode thermography is not ideal for cases with a high rate of transmission—for instance

most plastics at low power intensities [17]. In some cases heating may be applied to the sample volumetrically with eddy currents

Infrared thermography has been used to inspect carbon fiber [18-22], steel [23, 24], ceramics [25, 26], coatings [14], concrete [27, 28] and in medical applications [29]. This thesis seeks to characterize defects in metal sheets via infrared thermography via scaling laws.

#### **Chapter Two: Experimental Methods**

#### i. Mathematical Formulation

A body subjected to incident radiation will reflect, absorb and transmit the incoming energy to various degrees depending on its material properties. For an opaque grey body the incident energy will be absorbed and reflected depending on the properties of the material's surface as described in Equation 2.1.

$$E_{incident} = E_{reflected} + E_{absorbed}$$
 (2.1)

The absorbed energy will be lost to the environment in forms of convection from the surface, radiation to the surroundings and conduction within the body as shown in Figure 2-1.

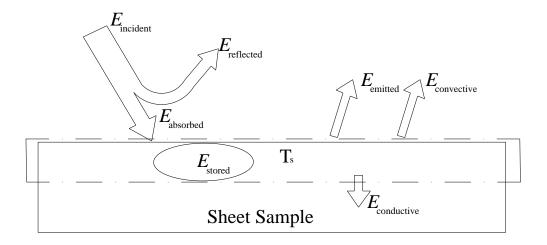


Figure 2-1 Energy balance diagram for a grey body

Performing an energy balance for the grey body, the amount of energy absorbed is equal to the energy stored, conducted through the body, emitted and lost to convection.

$$E_{absorbed} = E_{conductive} + E_{stored} + E_{convective} + E_{emitted,T_{convective}}$$
 (2.2)

Substituting in equations in for each term in Equation 2.2

$$E_{absorbed} = lA_{px}\nabla(k\cdot\nabla T) + lA_{px}\frac{k}{\alpha}\frac{\partial T_{s}}{\partial t} + hA_{px}(T_{s} - T_{\infty})$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \,\sigma A_{px}F_{px-\infty}(T_{s}^{4} - T_{\infty}^{4})\partial\omega\partial\lambda \qquad (2.3)$$

 $A_{px}$ , l and  $F_{px-\infty}$  are the pixel area, thickness of the material and the view factor of one pixel to infinity respectively. The energy imparted unto the infrared detector for each pixel,  $E_d$ , is a combination of the energy reflected and emitted from the grey body as well as radiation from the surroundings as described in Equation 2.4.

$$E_d = F_{px-d} \left( E_{reflected} + E_{emitted,T_d} + C \right) \tag{2.4}$$

C is a constant and includes the radiation from the surroundings and machine error. Solving Equation 2.1 for the energy reflected.

$$E_{reflected} = E_{incident} - E_{absorbed}$$
 (2.5)

Substituting Equation 2.5 into Equation 2.4 to describe the energy reaching the detector in terms of the incident, absorbed and emitted energy:

$$E_d = F_{px-d}(E_{incident} - E_{absorbed} + E_{emitted} + C)$$
 (2.6)

Substituting Equation 2.3 into Equation 2.6:

$$E_{d} = F_{px-d} \left( E_{incident} - lA_{px} \nabla (k \cdot \nabla T) - lA_{px} \frac{k}{\alpha} \frac{\partial T_{s}}{\partial t} - hA_{px} (T_{s} - T_{\infty}) \right)$$

$$- \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma A_{px} F_{px-\infty} (T_{s}^{4} - T_{\infty}^{4}) \partial \omega \partial \lambda$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, A_{px} \sigma (T_{s}^{4} - T_{d}^{4}) \partial \omega \partial \lambda + C$$

$$(2.7)$$

The incident radiation is affected by the view factor of the pixel to the heater ( $F_{px-H}$ ), the area of the pixel ( $A_{px}$ ) and the energy emitted from the heater ( $E_H$ ). Note the energy emitted from the heater is determined a priori. The energy emitted from the heater is assumed to be much greater than that of the surroundings.

$$E_{incident} = A_{px} F_{px-H} E_H \tag{2.8}$$

Substituting the incident energy from Equation 2.8 into Equation 2.7 produces Equation 2.9.

$$E_{d} = F_{px-d} \left( A_{px} F_{px-H} E_{H} - l A_{px} \nabla (k \cdot \nabla T) - l A_{px} \frac{k}{\alpha} \frac{\partial T_{s}}{\partial t} - h A_{px} (T_{s} - T_{\infty}) \right)$$

$$- \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma A_{px} F_{px-\infty} (T_{s}^{4} - T_{\infty}^{4}) \partial \omega \partial \lambda$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, A_{px} \sigma (T_{s}^{4} - T_{d}^{4}) \partial \omega \partial \lambda + C$$

$$(2.9)$$

Factoring out the area of a pixel, assuming it is constant.

$$E_{d} = A_{px} F_{px-d} \left( F_{px-H} E_{H} - l \nabla (k \cdot \nabla T) - l \frac{k}{\alpha} \frac{\partial T_{s}}{\partial t} - h(T_{s} - T_{\infty}) \right)$$

$$- \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma F_{px-\infty} (T_{s}^{4} - T_{\infty}^{4}) \partial \omega \partial \lambda$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma \left( T_{s}^{4} - T_{d}^{4} \right) \partial \omega \partial \lambda + C_{2}$$

$$(2.10)$$

For a steady state condition with  $T_s \approx T_{\infty}$ , Equation 2.10 can be simplified as there is no longer storage, convection or emission to the surroundings.

$$E_d = A_{px} F_{px-d} \left( F_{px-H} E_H - l \nabla (k \cdot \nabla T) + \int_0^\infty \int_0^{4\pi} \varepsilon_S \, \sigma \left( T_S^4 - T_d^4 \right) \partial \omega \partial \lambda + C_2 \right)$$
 (2.11)

Taking the gradient in x, y and z of the detector energy:

$$\nabla E_{d} = A_{px} \nabla F_{px-d} \left( F_{px-H} E_{H} - l \nabla (k \cdot \nabla T) + \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma \left( T_{s}^{4} - T_{d}^{4} \right) \partial \omega \partial \lambda \right)$$

$$+ C \right)$$

$$+ A_{px} F_{px-d} \left[ \nabla \left( F_{px-H} E_{H} \right) - \nabla \left( l \nabla (k \cdot \nabla T) \right) \right]$$

$$+ \nabla \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma \left( T_{s}^{4} - T_{d}^{4} \right) \partial \omega \partial \lambda \right]$$

$$(2.12)$$

Assuming the energy from the heater is uniform.

$$\nabla E_{d} = \frac{\nabla F_{px-d}}{F_{px-d}} E_{d}$$

$$+ A_{px} F_{px-d} \left[ E_{H} \nabla F_{px-H} - \nabla (l \nabla (k \cdot \nabla T)) \right]$$

$$+ \nabla \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma (T_{s}^{4} - T_{d}^{4}) \partial \omega \partial \lambda \right]$$
(2.13)

Taking the Laplacian of the detector energy:

$$\nabla^{2}E_{d} = \frac{F_{px-d}\nabla^{2}F_{px-d} - \nabla F_{px-d}^{2}}{F_{px-d}^{2}}E_{d} + \frac{\nabla F_{px-d}}{F_{px-d}}\nabla E_{d}$$

$$+ A_{px}\nabla F_{px-d}\left[E_{H}\nabla F_{px-H} - \nabla(l\nabla(k\cdot\nabla T))\right]$$

$$+ \nabla\int_{0}^{\infty}\int_{0}^{4\pi}\varepsilon_{s}\,\sigma(T_{s}^{4} - T_{d}^{4})\partial\omega\partial\lambda\right]$$

$$+ A_{px}F_{px-d}\left[E_{H}\nabla^{2}F_{px-H} - \nabla^{2}(l\nabla(k\cdot\nabla T))\right]$$

$$+ \nabla^{2}\int_{0}^{\infty}\int_{0}^{4\pi}\varepsilon_{s}\,\sigma(T_{s}^{4} - T_{d}^{4})\partial\omega\partial\lambda\right]$$

$$(2.14)$$

Substituting in the value for the gradient in Equation 2.13 and simplifying.

$$\nabla^{2}E_{d} = \frac{\nabla^{2}F_{px-d}}{F_{px-d}}E_{d}$$

$$+ 2A_{px}\nabla F_{px-d}\left[E_{H}\nabla F_{px-H} - \nabla(l\nabla(k\cdot\nabla T))\right]$$

$$+ \nabla\int_{0}^{\infty}\int_{0}^{4\pi}\varepsilon_{s}\,\sigma(T_{s}^{4} - T_{d}^{4})\partial\omega\partial\lambda\right]$$

$$+ A_{px}F_{px-d}\left[E_{H}\nabla^{2}F_{px-H} - \nabla^{2}(l\nabla(k\cdot\nabla T))\right]$$

$$+ \nabla^{2}\int_{0}^{\infty}\int_{0}^{4\pi}\varepsilon_{s}\,\sigma(T_{s}^{4} - T_{d}^{4})\partial\omega\partial\lambda\right]$$

$$(2.15)$$

For a defect free isentropic metal sheet of uniform thickness, constant material properties and surface roughness l, k and  $\varepsilon_s$  are constant.

$$\nabla^{2}E_{d} = \frac{\nabla^{2}F_{px-d}}{F_{px-d}}E_{d}$$

$$+ 2A_{px}\nabla F_{px-d}\left[E_{H}\nabla F_{px-H} - lk\nabla^{3}T\right]$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \,\sigma\nabla\left(T_{s}^{4} - T_{d}^{4}\right)\partial\omega\partial\lambda\right]$$

$$+ A_{px}F_{px-d}\left[E_{H}\nabla^{2}F_{px-H} - lk\nabla^{4}T\right]$$

$$+ \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \,\sigma\nabla^{2}\left(T_{s}^{4} - T_{d}^{4}\right)\partial\omega\partial\lambda\right]$$

$$(2.16)$$

Assuming the temperature of the detector is constant with respect to position and that the view factor of each pixel to the heater and to the detector is approximately constant in the neighborhood of a pixel.

$$\nabla^{2} E_{d} = A_{px} F_{px-d} \left( -lk \nabla^{4} T + \int_{0}^{\infty} \int_{0}^{4\pi} \varepsilon_{s} \, \sigma \nabla^{2} T_{s}^{4} \partial \omega \partial \lambda \right)$$
 (2.17)

For metal sheets with a Biot number  $\ll$  .1 the temperature of the sheet can be assumed constant.

$$\nabla E_d \approx 0 \text{ and } \nabla^2 E_d \approx 0$$
 (2.18)

The temperature of each pixel is correlated to the radiation sensed by the detector over its detection band.

$$T_m = \int_{\lambda_{D,l}}^{\lambda_{D,h}} \int_0^{4\pi} E_d \, \partial\omega \partial\lambda \tag{2.19}$$

Where  $\lambda_{D,h}$  and  $\lambda_{D,l}$  are the maximum wavelength and mimum wavelength the detector is sensitive to. Thus  $\nabla T_m \approx 0$  and  $\nabla^2 T_m \approx 0$  if  $\nabla E_d \approx 0$  and  $\nabla^2 E_d \approx 0$  respectively. The

parameters that influence the temperature measurement for cylindrical defects are listed in Equation 2.20.

$$T_m = f(k, \rho, C, \varepsilon_s, T_\infty, T_s, Q_H, T_H, A_{px}, D, l)$$
(2.20)

#### ii. Apparatus

Samples were inspected in reflection mode infrared thermography. Since specular or partial specular reflection is expected with the samples, the infrared detector and heater are positioned at opposite but equal angles to the sample surface's normal as shown in Figure 2-2.

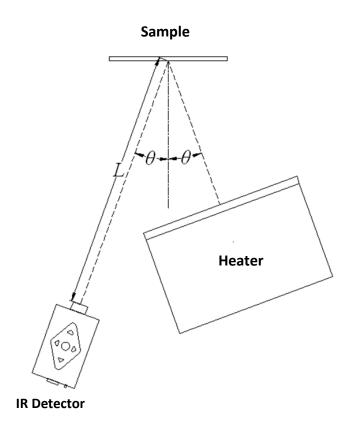


Figure 2-2 Schematic of Experimental Setup

#### Detector

Two FLIR ThermoVision SC4000 photonic—made from Indium Antimonide—mid-wave infrared (MWIR) detectors were used for this experiment. The ThermoVision SC4000 MWIR has a spectral range of 3  $\mu$ m to 5  $\mu$ m with a pixel pitch of 30  $\mu$ m x 30  $\mu$ m and a sensitivity of 0.025 °C [30]. The cameras were equipped with a 25 mm and 50 mm focal length lenses and were

positioned 31 cm and 54 cm away from the sample to produce a spot or pixel (px) size of .375 mm and .325 mm respectively. The SC4000 with the 25 mm lens was used for heater settings less than or equal to  $100\,^{\circ}$ C, while the SC4000 with the 50 mm lens was used for heater temperatures greater than  $100\,^{\circ}$ C.

#### Heaters

The Infrared Systems Development IR-160 black body with an IR-301 temperature controller was selected as it is a nearly uniform heater with black body emission characteristics. The heating portion of the IR-160 is 12" (30.5 cm) square and has a surface temperature that varies less than  $\pm$  0.1 °C spatially and temporally (for a short term) with an emissivity of 0.96  $\pm$  0.02%[31, 32].

Two 500W Infrared Salamander ceramic heaters were chosen as the emission spectrum of ceramics ranges from 2  $\mu$ m to 10  $\mu$ m peaking at 5  $\mu$ m[25, 33]. Each heater draws a maximum of 500W at 120V and is controlled by a potentiometer. Total wattage used was measured by a Kill-A-Watt P4400 which is accurate to  $\pm 0.1$ W was placed before the potentiometer. The dimensions of the heater as provided by the manufacturer are shown in Figure 2-3.

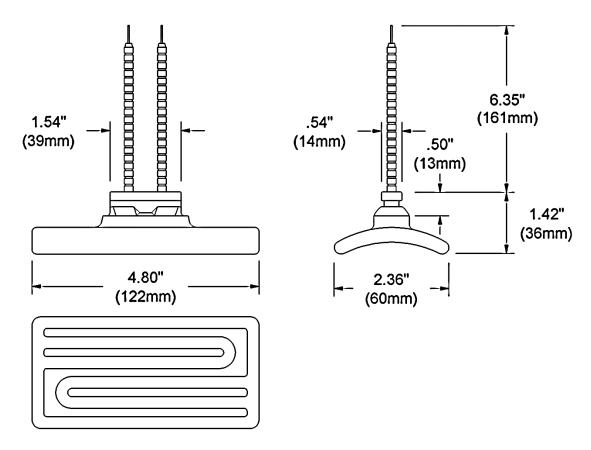


Figure 2-3 Infared Salamander Half Trough Emitter Schematic [34].

The overall view factor to the sample to the blackbody heater is .19 and .035 for the IR ceramic heater. The view factors were calculated using the contour double integral formula (CDIF) while assuming both were planar surfaces[35, 36]. The IR-160 was operated from 303 K to 573 K while the Ceramic Heater was operated with a heat flux between 480 W/m² to 6800 W/m². To compare the two heaters, the heat quality—the amount of exergy destroyed— was considered. For this case as there is no work or chemical exergy to consider.

$$HQ = Q_h \left( 1 - \frac{T_c}{T_h} \right)$$

Where  $T_c$  is the ambient temperature,  $T_h$  is the heater temperature and  $Q_h$  is the heat emitted from the heater. Settings were chosen such that the heat output from both heaters matches as shown Table 2.1 lists the different heater settings used.

Table 2.1 Heater settings for the black body and ceramic heater; controlled parameters are in bold.

В	Black Bod	у	Ceramic Heater			
$Q_h$ [W/m $^2$ ]	<i>T</i> [K]	$HQ$ [W/m $^2$ ]	$Q_h$ [W/m $^2$ ]	<i>T</i> [K]	HQ [W/m <sup>2</sup> ]	
479	303	8	478	313	8	
618	323	48	615	318	47	
1099	373	221	1093	332	218	
2842	473	1051	2732	389	994	
6119	573	2936	6831	488	3374	
			10245	538	4665	

#### iii. Sample Preparation

Sheets of Aluminum alloy 5052, stainless steel type 304 and Galvanized steel ASTM A653 (Carbon Steel type B hot dipped in Zinc finish) where chosen with thicknesses of 1.9mm to 2.0mm. Two finishes were tested for the Aluminum and stainless steel sheets. The Aluminum sheet was selected in a mirror-like No. 8 finish and a Satin No. 4 finish, while the stainless steel was selected in a mirror-like No. 8 finish and an unpolished finish. Thermal properties from literature are listed in Table 2.2 [4, 37, 38].

Table 2.2 Thermal properties of materials studied.

Makadal	Thermal Conductivity	Density	Specific Heat	Emis	sivity
Material	( <i>k</i> )	$(\rho)$	( <i>C</i> )	(8	E)
	W/m.K	kg/m³	J/kg.K	1 <sup>st</sup> finish	2 <sup>nd</sup> finish
Aluminum	167	2700	896	0.01-0.10	0.30-0.60
Stainless Steel	16.2	8000	500	0.01-0.15	0.20-0.60
Galvanized Steel	47.3	7850	475	0.15-0.45	

From the metal sheets, 20cmx20cm plates were cut and 5 cylindrical holes were drilled into each plate with an E-mill Acer EVS-3VKH to simulate defects. The defect pattern is shown in Figure 2-4 and the nominal dimensions for each defect are listed in Table 2.3.

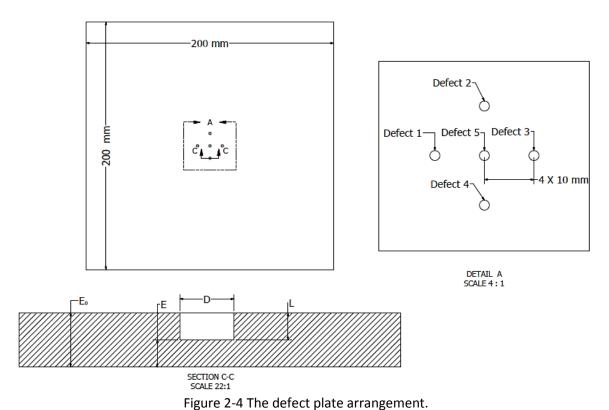


Table 2.3 The nominal depth and diameter for each defect plate.

	Defe	ct 1	Defe	ect 2	Defe	ect 3	Defe	ect 4	Defe	ect 5
Plate	D ( )	L	D ( )	L ()	D ( )	L	D ( )	L ( )	D ( )	L
	(mm)	(mm)	(mm)	(mm)	(mm)	(mm)	(mm)	(mm)	(mm)	(mm)
Array 1	1.00	1.00	1.00	0.50	1.00	0.25	1.00	0.75	1.00	0.10
Array 2	0.50	1.00	0.50	0.50	0.50	0.25	0.50	0.75	0.50	0.10
Array 3	1.50	0.25	1.00	0.25	0.50	0.25	2.00	0.25	0.25	0.25

After milling the defect depths were measured with a Brown and Sharpe 6 inch vernier caliper with an accuracy of ±.0253mm. To measure diameter, microscopic images were taken of each defect with a Scalar DG-2A equipped with a 50x magnification lens to measure the defect diameter with a pixel size of .003mm. Two methods were used in the determination of the defect diameter. In the first method, points along the inner-edge of the defect were manually chosen and then fitted to a circle via the Taubin method[39, 40]. For this method, a minimum of 16 evenly spaced points were chosen, with more points used for larger defects. Figure 2-5 is the image of defect with a nominal size of 1.50mm in the mirror steel sheet with the fitted circle shown.

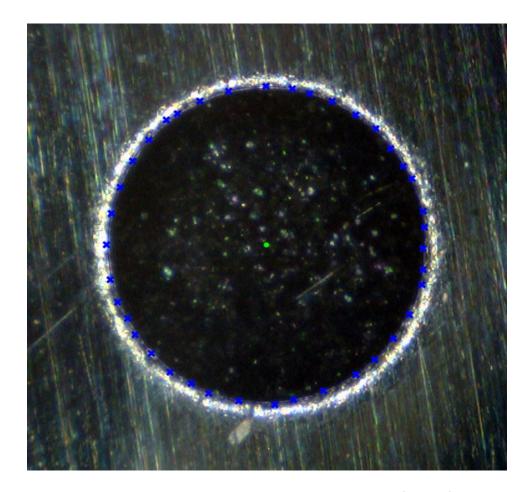


Figure 2-5 Circle Fitted to points chosen along the edge of the defect.

In the second method, the image of the defect is processed into a black and white image and the defect is then isolated from the surroundings and any gaps inside the defect image are filled in manually. This processes is demonstrated in Figure 2-6

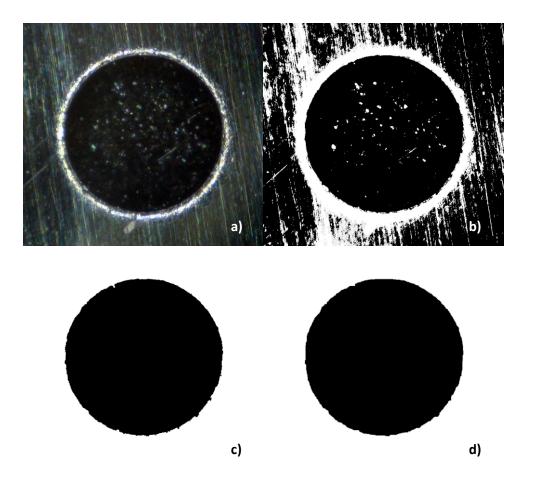


Figure 2-6 a) Microscopic image of defect in mirror finish steel sheet b) Processed black and white image c) Processed image with surroundings removed and holes filled in d) Errant pixels removed and morphological operations performed.

From Figure 2-6 d) the diameter of the defect is calculated in two ways. The pixels constituting the area of the defect are summed to find the area of the defect and from that area the diameter is determined assuming the defect is circular. Secondly, the edge pixels are fitted to a circle by the Taubin method. The results from both methods are averaged and compared to manually fitted circle for verification. The measured defect sizes can be found in Appendix A:.

#### iv. Estimation of Uncertainty

There are two major sources of error in the calculation of contrast, the heaters and the infrared detector. The SC4000 has an uncertainty of ±2% and the actual temperature has a greater error, but the difference (not the absolute value) is what is what is important. The black body heater has a spatial variation of 0.1 K and a temporal variation of 0.1 K, but the temporal variation is unimportant as 1 frame can be used for the calculation and that frame is captured in 2 ms or less. Examining the equation for contrast, S, in terms of temperature:

$$S = \frac{T_{px} - T_b}{T_b}$$

Where,  $T_b$  is the background temperature—for simplicity, the heater temperature—and  $T_{px}$  is the temperature represented by the pixel of interest. Thus the error in  $T_b$  is mainly caused by the heater's variability. The root sum square of these uncertainties is shown below.

$$S_{unc} = \sqrt{\left(\frac{\partial S}{\partial T_{px}} u_{T_{px}}\right)^2 + \left(\frac{\partial S}{\partial T_b} u_b\right)^2} = \sqrt{\left[\frac{1}{T_b} \left(T_{px} * .02\right)\right]^2 + \left[\frac{T_{px}}{T_b^2} \left(.1\right)\right]^2}$$

Considering the worst case scenario with  $T_b=303.15~K$  and  $T_{px}=T_{\infty}=298.15~K$ :

$$S_{unc} \le \sqrt{\left[\frac{1}{303.15}(298.15 * .02)\right]^2 + \left[\frac{298.15}{303.15^2}(.1)\right]^2} = .027$$

There is no data for the variability of the Ceramic heaters and while the error for them should be greater than that of the blackbody heaters it is unknown, but assumed to be of the same order of magnitude as the black body heaters.

#### v. Experimental Procedure

- The sample holder is bolted to an optical table and then the distances for the heater and detector are measured and marked. The heater and camera are then placed in their respective spots.
- 2. The heater is turned on, set to the proper temperature or wattage and left to reach steady state for 1 hour.
- 3. A sample is placed in the holder and the camera is focused.
- 4. Samples to be inspected are then placed in the sample holder and data is recorded at 5 frames per second for one minute
- 5. Sample is removed and the recording is checked
- 6. Repeat steps 2-5 as needed.
- 7. Data is transferred into the MATLAB file format and defect centers are manually located.

#### **Chapter Three: Results and Discussion**

Contrast was used to determine the effect of each parameter tested. Contrast is how distinguishable an object is from its surroundings. However, contrast for an arbitrary image is not uniquely defined in literature [41]. In the case of a large, uniform background with a small test target, Weber contrast is commonly used [41]. The Weber contrast fraction is defined in Equation 3.1.

$$S = \left| \frac{I - I_b}{I_b} \right| \tag{3.1}$$

Where S is contrast, I is the intensity and  $I_b$  is the background intensity. In this case, the background intensity is the average temperature of a 10 px by 10 px sound area which is manually chosen. In the case of non-uniform background heating contrast is not well defined [41], but weber contrast will still be used in those cases. For the cases where contrast is calculated based on temperature the result is multiplied by 100 as dividing differences on the order of .5-10 K by values greater than 300K—this operation is not performed for contrasts of the Laplacian and gradient based thermograms.

# i. Sample Images

Presented in this section are infrared images of each material and finish under different heating conditions.

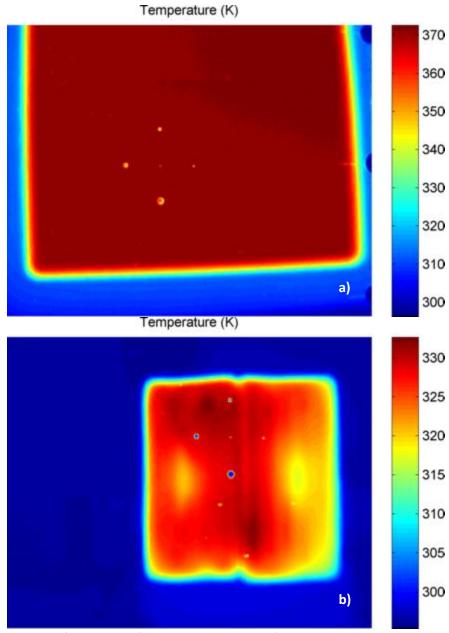


Figure 3-1 Array 3 of mirror-like finish Aluminum with a) black body heater at 373 K b) ceramic heater at 1093 W/m<sup>2</sup>.

As shown in Figure 3-1, the black body heater provides even heating over its heating area, but the unevenness of the ceramic heater is quite pronounced for a low emissivity material such as mirror-like finish Aluminum alloy. In both cases the defects are visible. Processing may be able to mitigate the effect of the temperature variations in the ceramic heater.

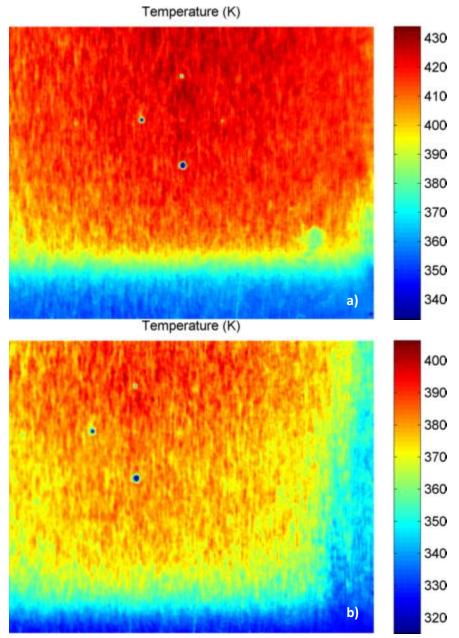


Figure 3-2 Array 3 of satin finish Aluminum with a) black body heater at 473 K b) ceramic heater at 1093  $W/m^2$ .

The surface roughness of the satin finish aluminum samples is noticeable in Figure 3-2. The increased roughness and emissivity minimize the differences between the black body and ceramic heater. Only the larger defects are easily distinguishable; the diameter of the smallest defect, 0.281 mm, appears to be of the same order of magnitude as the width of the surface grooves.

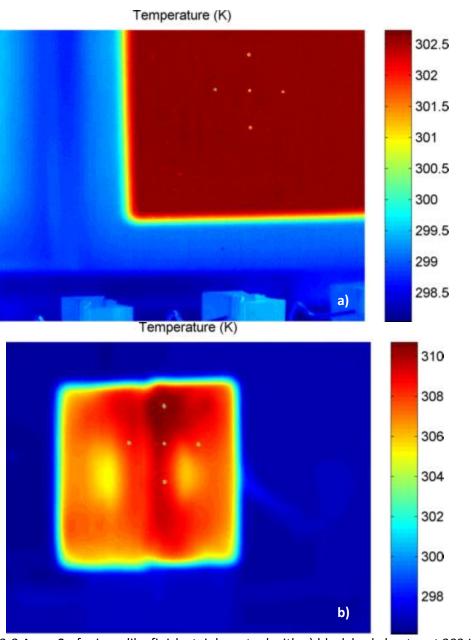


Figure 3-3 Array 2 of mirror-like finish stainless steel with a) black body heater at 303 K b) ceramic heater at 478 W/m<sup>2</sup>.

The mirror-like finish stainless steel shown in Figure 3-3 appears very similar to that of the mirror-like finish Aluminum. Even with defects below the Nyquist frequency are visible on uniform and very low emissivity surfaces.

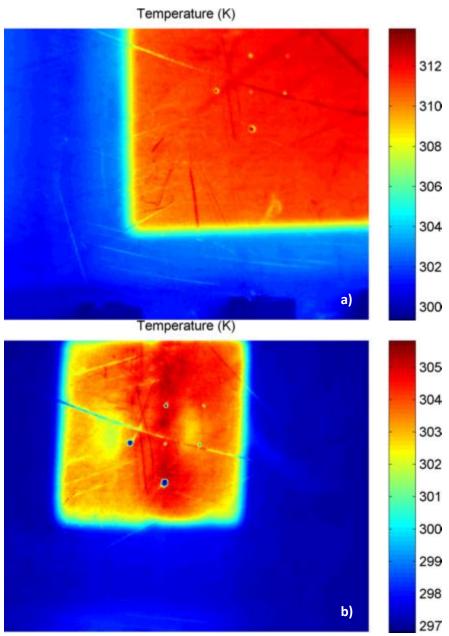


Figure 3-4 Array 3 of unpolished stainless steel with a) black body heater at 323 K b) ceramic heater at 615 W/m<sup>2</sup>.

As shown in Figure 3-4, the unpolished steel is not uniform, with scratches and other markings.

The defects in both are visible, but the 4<sup>th</sup> defect is very close to another discontinuity.

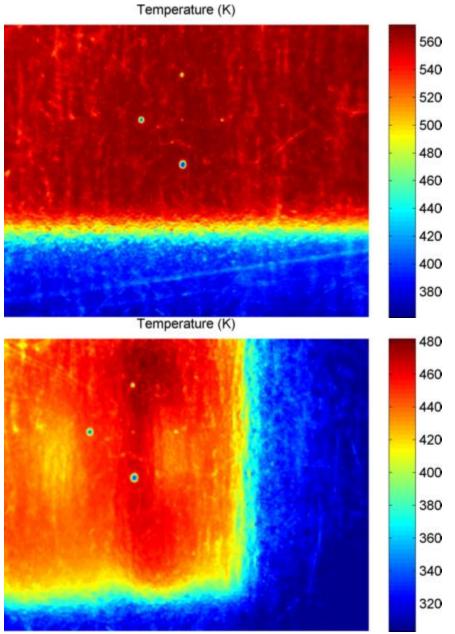


Figure 3-5 Temperature maps of Array 3 of galvanized steel with a) black body heater at 573 K b) ceramic heater at 6831 W/m².

The galvanized steel samples in Figure 3-5 show an uneven surface finish which may be due to variations in emissivity or the thickness of the Zinc coating.

## ii. Effect of Diameter

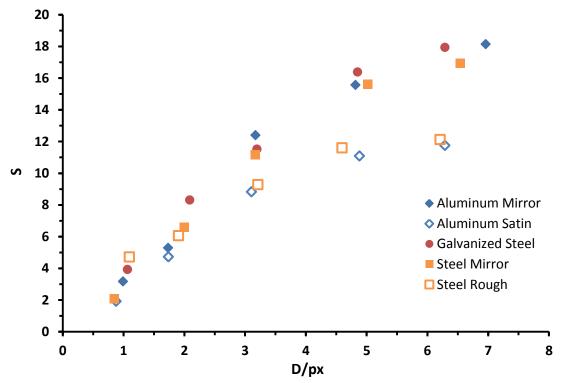


Figure 3-6 Diameter divided by pixel size plotted against contrast for array 3 with the black body heater at 373 K.

The effect of diameter appears to vary logarithmically with emissivity more so than material properties. To illustrate this, the Aluminum and Steel samples were plotted on a logarithmic scale in Figure 3-7.

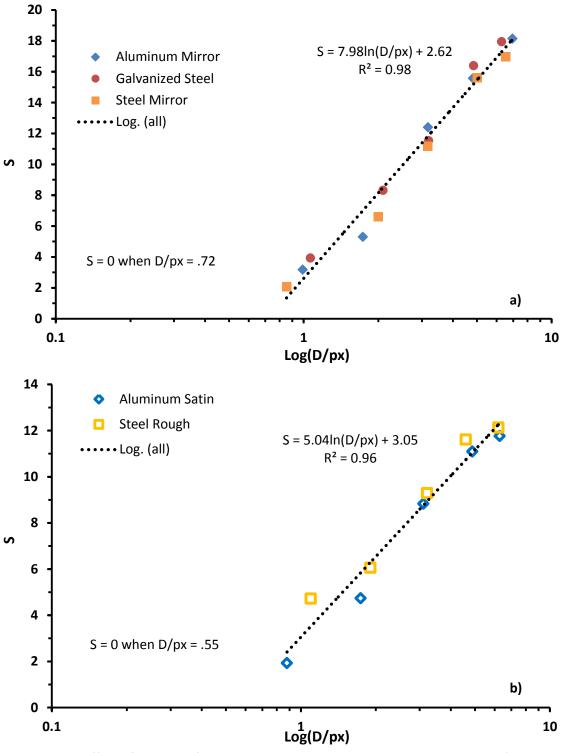


Figure 3-7 Effect of diameter for the array 3 with the black body heater at 373 K a) Low emissivity samples b) High emissivity samples.

The contrast appears to vary in a logarithmic manner and is dependent upon surface finish, not material properties. The change in contrast decreases as the diameter of the defect is represented by more pixels. In both cases, sub-pixel defects can be detected.

For defects that are uniform, the largest change is around the edge of the defect; having more pixels represent the interior of a large, uniform defect does not appear to provide a large increase in contrast. The pixels representing where the discontinuity begins show the largest change, while the uniform portion of the defect does not vary greatly as shown in Figure 3-8.

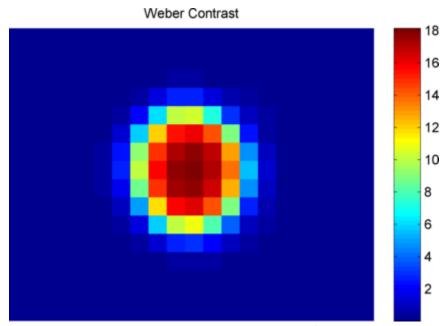


Figure 3-8 Contrast of a defect with D/px of 6.

### iii. Detected Diameter

The diameter of each defect was determined via the same method used in the Sample Preparation section. The defect was isolated and edge detection was performed in MATLAB 2009a via the Sobel method with the default threshold. The area encompassed by the defect was then equated to the defect's diameter by assuming that it was circular.

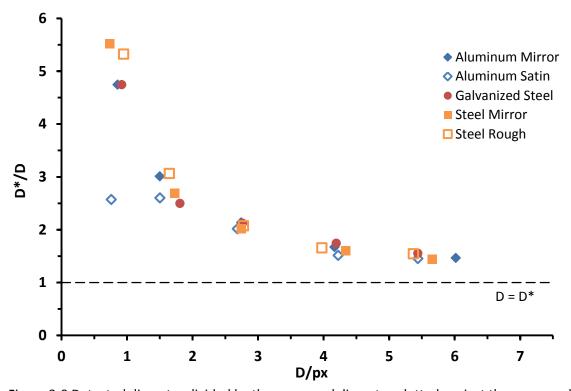


Figure 3-9 Detected diameter divided by the measured diameter plotted against the measured diameter over pixel size. Ceramic heater set to 478 W/m2.

In Figure 3-9, the detected infrared diameter, D\*, is 30% greater than the measured diameter, D, at the highest diameter over pixel size and appears to be approaching one as there are more pixels comprise the defect.

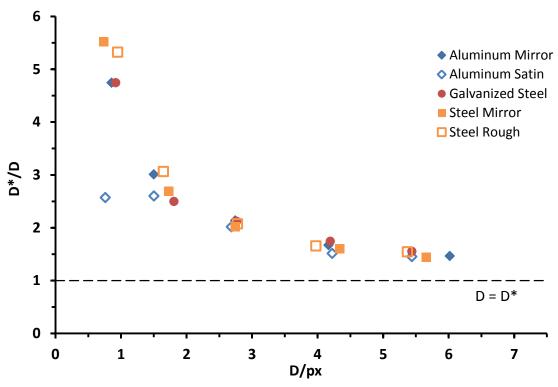


Figure 3-10 Detected diameter divided by the measured diameter plotted against the measured diameter over pixel size. Black body heater set to 573 K.

Figure 3-10 displays a similar trend to that of Figure 3-9, even with the tenfold increase in energy. As the number of pixels per defect increases the detected size approaches the measured size, while consistently staying above the measured size. The detected diameter is approximately 40% greater than the measure diameter at the largest D/px.

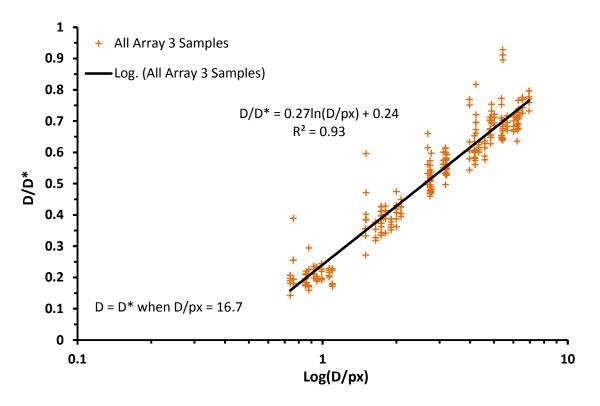


Figure 3-11 Measured diameter divided by the detected diameter and plotted against the measured diameter over pixel size for all array 3 tests.

Considering the data points from all of the tests in Figure 3-11, the detected diameter is 30-40% greater than that of the measured diameter. If the logarithmic trend holds the detected diameter will reach parity with the measured diameter when  $D/px \approx 17$ . This finding is congruent with a finding by Rantala et al. for the detection of cracks in ceramics via photothermal microscopy; the cracks were at least 34% longer than their optical measurement and the measurements were independent of modulation frequency [42]. Furthermore, the emission spectrum of the heater and the intensity does not appear to have a large effect on the detected diameter, nor does the material's thermal properties. Photothermal and infrared techniques may provide reliable crack length information in ceramics [26]. It requires some knowledge of the defect characteristics as it overestimates in a predictable manner. It appears that the infrared detector is capturing the effect of the defect on its surroundings.

### iv. Effect of Aspect Ratio

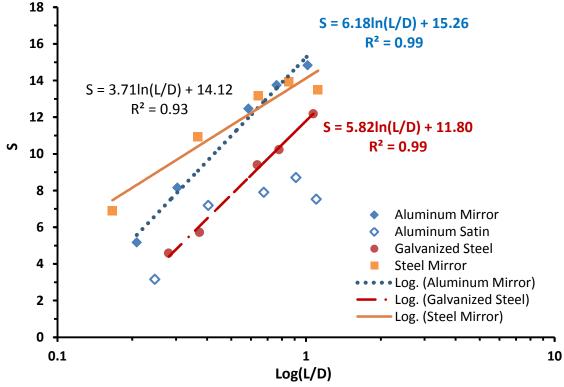


Figure 3-12 The effect of aspect ratio on contrast using data from array 1 with the black body heater at 373 K.

The effect of aspect ratio on contrast appears to be logarithmic for al samples in Figure 3-12 except the satin finish Aluminum which appears to follow no trend. The effect of aspect ratio is dependent upon the surface and material properties of the material and the defect. The detector is measuring the radiance from the defect walls. As the defect aspect ratio increases, less radiation from the bottom and the lower wall portions of the defect reaches the detector. Thus the surface characteristics of the defect are important as well.

### v. Apparent Emissivity

The emissions from a cylindrical cavity may be characterized by the apparent emissivity, which gives the total energy leaving the opening of the cavity divided by that of the energy emitted by a black-walled cavity; which is the same as a black area the size of the cavity opening [37]. The effect of aspect ratio on the hemispherical equivalent emissivity is plotted in Figure 3-13.

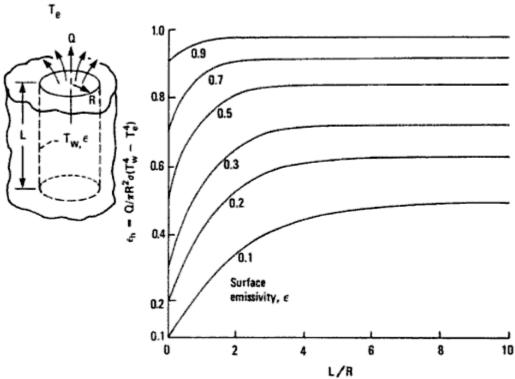


Figure 3-13 Apparent emissivity of a cavity opening for a cylindrical cavity of finite length with diffuse reflecting walls at a constant temperature. From Robert Siegel, John R. Howell; *Thermal Radiation Heat transfer*, 4th Edition [37].

To compare this information to the experimental data, the apparent emissivity,  $\varepsilon_a$ , of each defect and the surface emissivity,  $\varepsilon_s$ , was calculated assuming that the energy absorbed by the sample would be emitted thus  $\varepsilon=\kappa_s$  where  $\kappa_s$  is the surface absorptivity. Since the samples are

opaque  $\varepsilon=1-\beta_s$  where  $\beta_s$  is the reflectivity. Thus  $\varepsilon=1-\left|\frac{T_d^4-T_\infty^4}{T_h^4-T_\infty^4}\right|$ , where  $T_d$  is the defect temperature,  $T_h$  is the heater temperature and  $T_\infty$  is the ambient temperature [37].

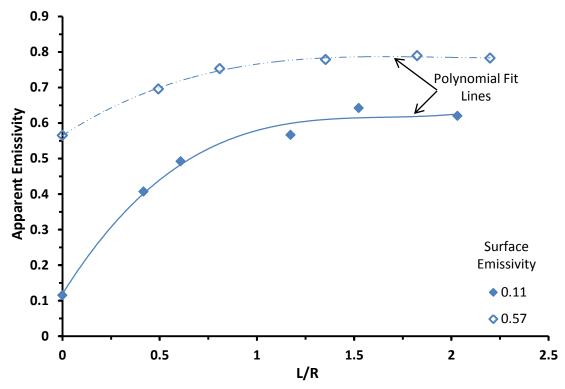


Figure 3-14 Apparent emissivity for the array 1 defects in the aluminum samples.

The trend from Figure 3-13 holds in Figure 3-14 even though the reflection—not the emission—from the samples is what is being measured as the samples are at ambient temperature. After a certain point, the change in aspect ratio has little to no effect on the apparent emissivity.

## vi. Effect of Heating and Heaters

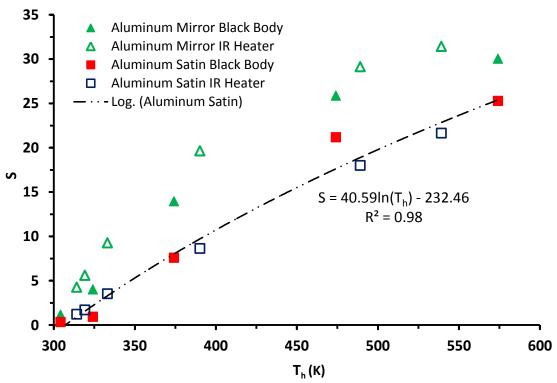


Figure 3-15 The effect of heating temperature on contrast by heater type on Aluminum samples with  $D \approx 2$ mm.

In Figure 3-15 the smoother mirror finish Aluminum the black body heater has a lower contrast in each case. This may be due to the variability of the heater or the difference in emission spectrums. The satin finish Aluminum does not exhibit the same behavior as both heaters are in the same general pattern with a very good curve fit for the combined data set. Instead of the temperature of the heater, consider the heat quality (HQ).

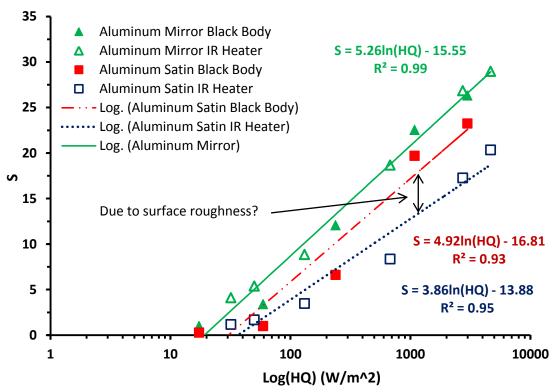
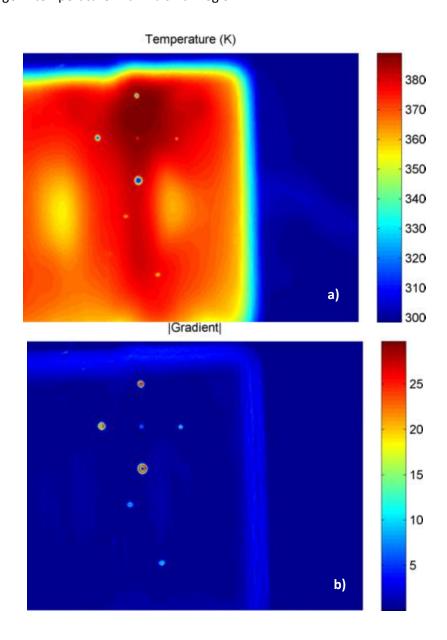


Figure 3-16 Effect of heat quality on contrast with respect to heater type.

In Figure 3-16, when HQ quality is on the abscissa the mirror-like sample converges while the rough sample diverges. The rough sample may interact differently with the longer wavelengths of the ceramic heater affecting the amount of absorption and scattering. It may be possible to use two heaters with different emission spectra to determine surface roughness with only one detection band for the detector.

### vii.Derivatives as Filters

The gradient and Laplacian of the uniform area of a thermogram should be approximately zero as demonstrated in the mathematical formulation. Furthermore, small changes in heater uniformity should also be smoothed out as most heaters do not feature a sharp change in temperature within a small region.



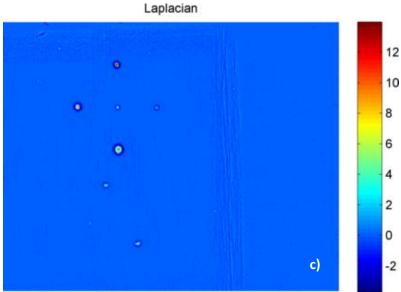


Figure 3-17 Aluminum Mirror Array 3 with ceramic heater at 2732 W/m<sup>2</sup> a) thermogram b) magnitude of the gradient of the thermogram c) Laplacian of the thermogram.

The derivatives images have a much higher signal for the defects than the surroundings as shown in Figure 3-17. The small variations in heater temperature are also mitigated. To demonstrate this contrast maps for the gradient and Laplacian were plotted in Figure 3-18. Note the sound area signal is approximately zero.

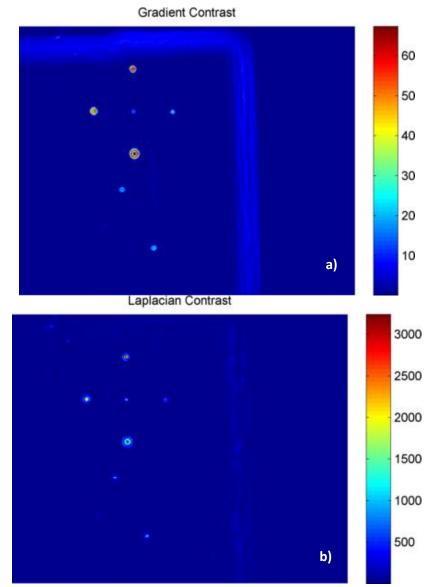


Figure 3-18 Contrast maps for a) gradient of temperature b) Laplacian of temperature.

The contrast for derivatives is much higher than that of the temperature contrast. It appears that the defect on the right is more easily detected by the gradient method, while the smallest defect in the center would be easier to detect by the Laplacian. Looking at the average contrast by material for all arrays and heating conditions in Figure 3-19

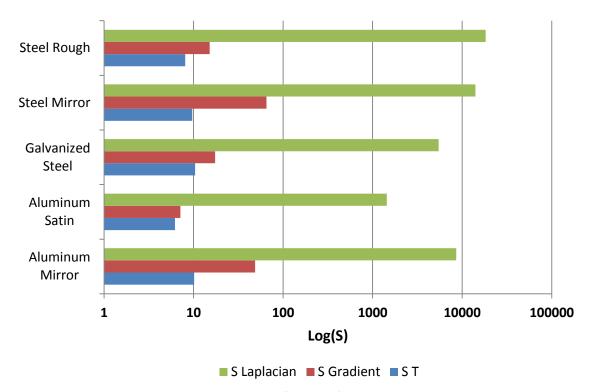


Figure 3-19 Average contrast by material for all defects, heating conditions and arrays.

The Laplacian contrast is always greater than the gradient contrast and both are always greater than the Temperature contrast. The temperature contrast is multiplied by 100 while the gradient and Laplacian contrasts are not.

### **Chapter Four: Conclusions and Future Work**

Reflection mode IRT has been used to characterize cylindrical defects in metal sheets with varying defect sizes, surface finishes and material properties. The contrast of the defects is mainly governed by heat quality, the apparent surface emissivity and defect geometry. The temperature contrast follows a logarithmic trend in the cases shown.

It is possible to detect sub-pixel defects in some of the finishes. The mirror finishes and galvanized steel show better contrast in relation to D/px— even though the galvanized steel does not have a uniform surface finish—indicating that the contrast in relation to the defect diameter depends more on the apparent emissivity. Furthermore, this may also be due to the samples being specular or partially specular reflectors.

Regardless of surface finish, the detected diameter via infrared thermography is larger than the measured diameter. For all cases the measured diameter is at least 30% larger than the measured diameter. The detected diameter will reach parity measured diameter at a D/px  $\approx$  16 if the trend holds. However, thermal effect of the defect on its surroundings should always cause the defect to appear larger than it is.

Derivatives can mitigate uneven heating and increase the contrast for defects even with rough surface finishes. The contrast for the gradient and Laplacian of temperature is always greater than that of the temperature contrast as the background intensity is approximately zero. The derivatives and thresholding can be used to detect defects and areas of interest.

In future work, I would like to attempt to measure surface roughness with two heaters of different wavelengths and one detector. Heat quality shows a large difference between heaters for the rough finish aluminum but not the mirror finish; the emissivity and surface roughness is suspected of being the cause.

Furthermore, study the same samples in transmission mode. Investigate different defect geometry and thermal property ratios; including defects with much larger diameters to verify the logarithmic trends seen and no change in geometry. Study the optical properties of semi-transparent media like plastics. Create a library allowing for the rapid deployment of IRT systems and possibly the automated detection of defects.

# **Appendix A: Defect Sizes**

Compiled is a list of measurements for the depth and diameter of each defect.

Table A.1 Defect Measurements for Aluminum Mirror-Like Finish Samples.

Aluminum Mirror-Like Finish				
	Array 1			
Defect	D (mm)	L (mm)	D/L	
1	1.116	1.14	0.98	
2	1.122	0.64	1.74	
3	1.091	0.36	3.06	
4	1.185	0.85	1.40	
5	1.103	0.23	4.78	
	Array 2			
Defect	D (mm)	L (mm)	D/L	
1	0.726	0.31	2.37	
2	0.593	0.65	0.91	
3	0.631	0.52	1.21	
4	0.637	0.47	1.34	
5	0.568	0.26	2.22	
	Array 3			
Defect	D (mm)	L (mm)	D/L	
1	1.533	0.41	3.76	
2	1.010	0.40	2.53	
3	0.552	0.41	1.36	
4	2.215	0.44	5.04	
5	0.315	0.51	0.61	

Table A.2 Defect Measurements for Aluminum Satin Finish Samples.

Α	Aluminum Satin Finish			
	Array	1		
Defect	D (mm)	L (mm)	D/L	
1	1.082	1.19	0.91	
2	1.064	0.72	1.48	
3	1.064	0.43	2.45	
4	1.064	0.97	1.10	
5	1.014	0.25	4.11	
	Array 2			
Defect	D (mm)	L (mm)	D/L	
1	0.588	1.09	0.54	
2	0.576	0.56	1.03	
3	0.582	0.30	1.93	
4	0.588	0.81	0.73	
5	0.594	0.57	1.04	
-	Array 3			
Defect	D (mm)	L (mm)	D/L	
1	1.563	0.42	3.69	
2	0.994	0.42	2.34	
3	0.556	0.48	1.15	
4	2.013	0.44	4.60	
5	0.281	0.30	0.95	

Table A.3 Defect Measurements for Steel Mirror-Like Finish Samples.

Steel Mirror-Like Finish				
	Array 1			
Defect	D (mm)	L (mm)	D/L	
1	1.084	1.21	0.89	
2	1.058	0.68	1.56	
3	1.184	0.44	2.72	
4	1.071	0.91	1.17	
5	1.052	0.17	6.01	
Array 2				
Defect	D (mm)	L (mm)	D/L	
1	0.650	1.14	0.57	
2	0.644	0.62	1.04	
3	0.625	0.37	1.71	
4	0.606	0.81	0.75	
5	0.613	0.15	4.09	
	Array 3			
Defect	D (mm)	L (mm)	D/L	
1	1.596	1.21	1.32	
2	1.010	0.68	1.48	
3	0.637	0.44	1.46	
4	2.082	0.91	2.28	
5	0.271	0.17	1.55	

Table A.4 Defect Measurements for Steel Unpolished Finish Samples.

Steel Unpolished Finish				
	Array 2			
Defect	D (mm)	L (mm)	D/L	
1	0.654	1.02	0.44	
2	0.635	0.56	1.14	
3	0.559	0.30	1.89	
4	0.616	0.54	1.13	
5	0.578	0.12	4.63	
	Array 3			
Defect	D (mm)	L (mm)	D/L	
1	1.47	0.28	5.32	
2	1.03	0.29	3.52	
3	0.61	0.29	2.09	
4	1.99	0.27	7.46	
5	0.35	0.19	1.88	

Table A.5 Defect Measurements for Galvanized Steel Samples.

Galvanized Steel				
	Array 1			
Defect	D (mm)	L (mm)	D/L	
1	1.164	1.24	0.93	
2	1.113	0.71	1.57	
3	1.069	0.40	2.68	
4	1.176	0.92	1.28	
5	1.063	0.30	3.57	
	Array 2			
Defect	D (mm)	L (mm)	D/L	
1	0.662	1.04	0.64	
2	0.669	0.59	1.13	
3	0.669	0.30	2.21	
4	0.638	0.83	0.77	
5	0.619	0.15	4.14	
	Array 3			
Defect	D (mm)	L (mm)	D/L	
1	1.552	0.43	3.58	
2	1.022	0.42	2.41	
3	0.669	0.42	1.58	
4	2.013	0.44	4.58	
5	0.341	0.37	0.91	

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## **Professional Positions**

May 2009 -Tennessee Valley Authority Fellow, Institute of Research for Technology Development, Lexington, KY May 2006 – May 2009 Bucks for Brains Fellow, University of Kentucky, Lexington, KY

## Scholastic/Professional Honors

UK Engineering Alumni Association; Secretary	2009-2011
Engineering Student Council; President	2009-2010
• American Society of Mechanical Engineering @ UK; President, VP	2007-2009
Nagoya University Summer Intensive Program; Participant	2010
• US Department of Energy Solar Decathlon; Decathlete	2008-2009
AIAA Design-Build-Fly Competition; Participant	2008-2009
International Karakuri Workshop; Participant	2007