# Neutrino and Antineutrino Induced Meson Production 

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# Neutrino and Antineutrino Induced Meson Production 

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## DEDICATION

This dissertation is dedicated to my family.

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#### Abstract

Coherent meson production measurement is very important in physics research. First, the coherent pion production is a potential background to $\nu$ oscillation in next generation of Long Base Line Experiment(ELBNF/DUNE); second, coherent pion and coherent $\rho$ production provide a detailed test of CVC and PCAC hypothesis; third, coherent meson production can be used to monitor the neutrino and anti-neutrino fluxes in the experiment. This dissertation focuses on two parts: coherent $\pi^{-}$production in NOMAD, and coherent $\rho$ simulation using LBNF fluxes. With the NOMAD data, the ratio between cross-sections of coherent $\pi^{-}$and $\bar{\nu}_{\mu}$ charged current interactions was measured and compared with the measurements of coherent $\pi^{+}$. The experience of coherent $\pi$ analysis may be used to evaluate the sensitivity of ELBNF/DUNE project to coherent processes. With the ELBNF process, I wrote a new C++ simulation package and generated 100 k coherent $\rho^{+}$events. It is known that for the neutrino-induced process, the incoming neutrino fluxes could not be measured directly, and the $Q^{2}$ and other variables related to it are unknown in the neutrino-induced neutral current interactions. The photon-induced coherent $\rho^{0}$ provides a way to get additional information to constrain the incoming neutrino fluxes. I calculated the ratios between the cross-sections of neutrino-induced coherent $\rho^{ \pm}, \rho^{0}$ and photon-induced coherent $\rho^{0}$. With these ratios, some kinematic variable distributions are reweighted with this ratio in this thesis.


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## Chapter 1

## Electroweak Theory and Neutrino Interaction

### 1.1 Weak Interaction and Electroweak Theory

Theoretically, all the interactions between particles can be classified into four fundamental interactions and listed in order of decreasing strength as: the strong interaction, electromagnetism, the weak interaction, and gravity. The weak interaction operates between all particles except photons and gravitons. It causes reactions which make particles ultimately decay into the stable leptons and hadrons, such as electrons, neutrinos, protons, and so on. These decays are the natural sources for us to study the weak interaction, however, only in a limited energy region. The advent of neutrino experiments in humans' laboratories ameliorated the situation by enabling us to explore weak interactions in a much wider energy region. The neutrino-hadron scattering is one such experiment.

The neutrino-hadron scattering can be classified with the help of a plane composed of $Q^{2}$ and $\nu$ [16], where $Q^{2}$ and $\nu$ are the square of the momentum transfer and the energy transfer between the initial and final leptons respectively ( $Q^{2}$ and $\nu$ are defined in chapter 2). Figure 1.1 shows the weak interactions with respect to $Q^{2}$ and $\nu$. In this figure, Region I with very small $Q^{2}$ and $\nu$ represents weak decays; Region II (diagonal line) represents (quasi-)elastic scattering; Region III is the resonance region, starting with the line $W=M+m_{\pi}(\mathrm{W}$ is the hadronic mass, M is the proton mass, $m_{\pi}$ is the pion mass); Region IV with high values of $Q^{2}$ and $\nu$ is the domain of deep inelastic scattering; And region V with low $Q^{2}$ and high $\nu$ values is coherent
scattering, which is the focus of this thesis. In this Region, the interactions with very low $Q^{2}$ allow study two basic properties of the weak current: One is the conservation of the vector current (It is also called CVC hypothesis), which was introduced to explain the equality of the vector muon and nuclear beta decay. The other one is the partial conservation of the axial current (It is also called PCAC hypothesis) which can be used to explain the small $(\sim 20 \%)$ renormalization of the nuclear axial decay constant by the strong interactions [36]. When the $Q^{2}$ is very low, the nuclear stays intact in coherent process, then the nucleons inside can not be considered as free. In this case, the perturbative theory of strong interactions can not be used. To study the processes with very low $Q^{2}$, the Hadron Dominance Model has been brought up to describe the hadronic behavior in coherent process, which is going to be introduced in this chapter. Therefore, the study of coherent meson production can also provide a detailed test of Hadron Dominance Model.


Figure 1.1: The $\left(Q^{2}, \nu\right)$ plane in neutrino interactions [36].

## Space-Time Structure of the Weak Charged and Neutral

## Current

This subsection gives basics of electroweak theory (which can be found in many textbooks, such as [33] and [30]). For convenience, we give the convention for the Dirac matrices in this thesis as follows.

Let $\left\{\gamma_{\mu}, \mu=0,1,2,3\right\}$ be an orthonormal set of vectors in space-time. The signature of space-time is expressed by the equations:

$$
\begin{gather*}
\gamma_{0}^{2}=1, \gamma_{1}^{2}=\gamma_{2}^{2}=\gamma_{3}^{2}=-1,  \tag{1.1}\\
\gamma_{\mu}=g_{\mu \nu} \gamma^{\nu} \tag{1.2}
\end{gather*}
$$

and the matrices follow an anti-commutation relation:

$$
\begin{equation*}
\left\{\gamma_{\mu}, \gamma_{\nu}\right\}_{+}=\frac{1}{2}\left(\gamma_{\mu} \gamma_{\nu}+\gamma_{\nu} \gamma_{\mu}\right)=g_{\mu \nu} \tag{1.3}
\end{equation*}
$$

As usual, a special multivector

$$
\begin{equation*}
i \gamma^{0} \gamma^{1} \gamma^{2} \gamma^{3}=\gamma_{5} \tag{1.4}
\end{equation*}
$$

is introduced which can be used to construct peudoscalars. A multivetor is said to be even (odd) if it commutes (anti-commutes) with $\gamma_{5}$. A slashed 4-momentum (or other 4 -vector) represents the product of the 4 -momentum (or other 4 -vector) with $\gamma_{\mu}$.

In the Standard Model the electroweak interaction is described by a gauge field theory. The gauge group is the $S U(2)_{L} \times U(1)_{Y}$ group, where $L$ indicates that this $S U(2)$ group only acts on the left-handed components of fermion fields. The subscript $Y$ for $U(1)$ group is called hypercharge and specifies this $U(1)$ group.

The local $S U(2)_{L} \times U(1)_{Y}$ gauge invariance of the electroweak Lagrangian is guaranteed by introducing the covariant derivative $D_{\mu}$,

$$
\begin{equation*}
D_{\mu}=\partial_{\mu}+i g A_{\mu}^{i} \tau^{i}+i g^{\prime} B_{\mu} \frac{Y}{2} \tag{1.5}
\end{equation*}
$$

where $\tau^{i}=\sigma^{i} / 2(i=1,2,3)$ with $\sigma^{i}$ being the Pauli matrix. $A_{\mu}^{i}$ and $B_{\mu}$ are gauge boson fields. For each generator of the group, there is a gauge field. The covariant derivative transforms as

$$
\begin{equation*}
D_{\mu} \rightarrow D_{\mu}^{\prime}=U\left(\theta^{i}(x), \eta(x)\right) D_{\mu} U^{-1}\left(\theta^{i}(x), \eta(x)\right) \tag{1.6}
\end{equation*}
$$

where

$$
\begin{equation*}
U\left(\theta^{i}(x), \eta(x)\right)=e^{i \theta^{i}(x) \tau^{i}+i \eta(x) Y / 2} \tag{1.7}
\end{equation*}
$$

with $\theta^{i}(x)$ and $\eta(x)$ being the parameters of the transformation. Then the gauge boson fields transform as

$$
\begin{align*}
A_{\mu}^{i} \tau^{i} \rightarrow A_{\mu}^{\prime i} \tau^{i} & =U\left(\theta^{j}(x)\right)\left[A_{\mu}^{i} \tau^{i}-\frac{i}{g} \partial_{\mu}\right] U^{-1}\left(\theta^{i}(x)\right)  \tag{1.8}\\
B_{\mu} & \rightarrow B_{\mu}^{\prime}=B_{\mu}-\frac{1}{g^{\prime}} \partial_{\mu} \eta(x) \tag{1.9}
\end{align*}
$$

The interaction part of the Lagrangian can be written as

$$
\begin{align*}
\mathcal{L}_{I}= & -\frac{1}{2} \overline{L_{L}}\left(g A^{i} \sigma^{i}-g^{\prime} \not B\right) L_{L}-\frac{1}{2} \bar{Q}_{L}\left(g A^{i} \sigma^{i}+\frac{1}{3} g^{\prime} \not B\right) Q_{L} \\
& +g^{\prime}-\overline{e_{R}} B e_{R}-\frac{2}{3} g^{\prime} \overline{u_{R}} \not B u_{R}+\frac{1}{3} g^{\prime} \bar{d}_{R} \not B d_{R}, \tag{1.10}
\end{align*}
$$

where $L_{L}$ and $Q_{L}$ represent the left-handed lepton doublet and the left-handed quark doublet in the fundamental representation of $S U(2)_{L}$ group respectively. Specifically, for the first generation of the Standard Model,

$$
\begin{equation*}
L_{L}=\binom{\nu_{e L}}{e_{L}}, \quad Q_{L}=\binom{u_{L}}{d_{L}} \tag{1.11}
\end{equation*}
$$

In order to see the interaction term for neutrinos explicitly, we take the explicit form of the Pauli matrices and obtain

$$
\begin{align*}
\mathcal{L}_{I, L}= & -\frac{1}{2}\left(\begin{array}{ll}
\nu_{e L} & e_{L}
\end{array}\right)\left(\begin{array}{cc}
g \mathcal{A}_{3}-g^{\prime} \notin B & g\left(\mathcal{A}_{1}-i \mathcal{A}_{2}\right) \\
g\left(\mathbb{A}_{1}+i A_{2}\right) & -g \mathcal{A}_{3}-g^{\prime} \mathbb{B}
\end{array}\right)\binom{\nu_{e L}}{e_{L}} \\
& +g^{\prime}-\overline{e_{R}} \notin e_{R}, \tag{1.12}
\end{align*}
$$

where we have omitted the quark part of the Lagrangian. This equation shows that the interactions are inter-mediated by the mixtures of the gauge fields. The offdiagonal terms of Equation (1.12) are conjugations of each other, so we can introduce a complex field $W_{\mu}$ by

$$
\begin{equation*}
W^{\mu} \equiv \frac{A_{1}^{\mu}-i A_{2}^{\mu}}{\sqrt{2}} \tag{1.13}
\end{equation*}
$$

Replacing $A_{1}^{\mu}$ and $A_{2}^{\mu}$ with $W^{\mu}$, we have, for interactions between neutrinos and electrons,

$$
\begin{align*}
\mathcal{L}^{(C C)} & =-\frac{g}{\sqrt{2}}\left\{\nu_{e L}^{-} W e_{L}+{\overline{e_{L}}}_{L} W^{\dagger} \nu_{e L}\right\} \\
& =-\frac{g}{2 \sqrt{2}} \overline{\nu_{e}} \gamma^{\mu}\left(1-\gamma^{5}\right) e W_{\mu}+H . c . \\
& =-\frac{g}{2 \sqrt{2}} j_{W, L}^{\mu} W_{\mu}+H . c . \tag{1.14}
\end{align*}
$$

where

$$
\begin{equation*}
j_{W, L}^{\mu}=\overline{\nu_{e}} \gamma^{\mu}\left(1-\gamma^{5}\right) e=2 \overline{\nu_{e}} \gamma^{\mu} e_{L} . \tag{1.15}
\end{equation*}
$$

The complex gauge field $W_{\mu}$ carries a charge which should be the electrical charge to guarantee the conservation of electrical charge. So the current $j_{W, L}^{\mu}$ is called the (leptonic) charged current and the corresponding part of Lagrangian is indicated by CC (which represents charged current) as a superscript (see Equation (1.14)).

The diagonal terms of Equation (1.12) are mixtures of $A_{3}^{\mu}$ and $B^{\mu}$, so we can introduce two fields $Z^{\mu}$ and $A^{\mu}$ as the linear combinations of $A_{3}$ and $B^{\mu}$ by

$$
\begin{align*}
& A^{\mu}=\sin \theta_{W} A_{3}^{\mu}+\cos \theta_{W} B^{\mu}  \tag{1.16}\\
& Z^{\mu}=\cos \theta_{W} A_{3}^{\mu}-\sin \theta_{W} B^{\mu} \tag{1.17}
\end{align*}
$$

where $\theta_{W}$ is a parameter to be determined and is usually called the weak mixing angle or Weinberg angle. Recalling that the electromagnetic interaction should come out of the electroweak theory, we require that $A_{\mu}$ is just the photon field. This requirement determines the Weinberg angle and the relation among the couplings $g, g^{\prime}$ and the electric coupling $e$.

To see how this happens, let us pick up the neutral current (NC) Lagrangian from Equation (1.12):

$$
\begin{equation*}
\mathcal{L}^{(N C)}=-\frac{1}{2}\left\{\overline{\nu_{e L}}\left(g A_{3}-g^{\prime} \not B\right) \nu_{e L}-\overline{e_{L}}\left(g \mathcal{A}_{3}+g^{\prime} \not B\right) e_{L}-2 g^{\prime} \overline{e_{R}} \mathbb{B} e_{R}\right\} \tag{1.18}
\end{equation*}
$$

Substituting $A_{3}^{\mu}$ and $B^{\mu}$ with $A^{\mu}$ and $Z^{\mu}$ by using Equations (1.16) and (1.17), we obtain

$$
\begin{align*}
\mathcal{L}^{(N C)}= & -\frac{1}{2}\left\{\nu_{e L}\left[g \cos \theta_{W}+g^{\prime} \sin \theta_{W}\right) \not \subset+\left(g \sin \theta_{W}-g^{\prime} \cos \theta_{W}\right) A\right] \nu_{e L} \\
& \left.-\overline{e_{L}}\left[g \cos \theta_{W}-g^{\prime} \sin \theta_{W}\right) \not \subset+\left(g \sin \theta_{W}+g^{\prime} \cos \theta_{W}\right) \mathcal{A}\right] e_{L} \\
& \left.-2 g^{\prime} \overline{e_{R}}\left[-\sin \theta_{W} \neq+\cos \theta_{W} A\right] e_{R}\right\} . \tag{1.19}
\end{align*}
$$

As neutral particles, neutrinos should not interact with the electromagnetic field, so we have

$$
\begin{equation*}
g \sin \theta_{W}=g^{\prime} \cos \theta_{W} \tag{1.20}
\end{equation*}
$$

Now taking this relation back into Equation (1.19), we obtain

$$
\begin{align*}
\mathcal{L}_{I, L}^{(N C)}= & -\frac{g}{2 \cos \theta_{W}}\left\{\overline{e_{e L}} \not \subset \nu_{e L}-\left(1-2 \sin ^{2} \theta_{W}\right) \overline{e_{L} \not Z e_{L}}+2 \sin ^{2} \theta_{W} \overline{\left.e_{R} \not \subset e_{R}\right\}}\right. \\
& +g \sin \theta_{W} \bar{e} A e . \tag{1.21}
\end{align*}
$$

The last term describes electrons interacting with photons, which is just what we need for electromagnetic interactions, so the coupling $g \sin \theta_{W}$ should be equal to the electrical coupling $e$ :

$$
\begin{equation*}
g \sin \theta_{W}=e \tag{1.22}
\end{equation*}
$$

We have 4 parameters $g, g^{\prime}, e$ and $\theta_{W}$ with 2 equations (1.20) and (1.22), so in principle, we can choose any two of them to describe the theory. From Equations (1.20) and (1.22), we can also deduce

$$
\begin{equation*}
\tan \theta_{W}=\frac{g^{\prime}}{g} \tag{1.23}
\end{equation*}
$$

and

$$
\begin{equation*}
g^{2}+g^{\prime 2}=e^{2} \tag{1.24}
\end{equation*}
$$

Experimentally, the value of $\sin ^{2} \theta_{W}$ can be extracted from neutral current processes, and is different according to different renormalization prescriptions. In the on-shell scheme, it is

$$
\begin{equation*}
\sin ^{2} \theta_{W}=1-\frac{M_{W}^{2}}{M_{Z}^{2}} \tag{1.25}
\end{equation*}
$$

where $M_{W}$ and $M_{Z}$ are the masses of $W$ and $Z$ respectively.
To sum up, the neutral current Lagrangian can be written as

$$
\begin{equation*}
\mathcal{L}^{(N C)}=\mathcal{L}^{Z}+\mathcal{L}^{\gamma}, \tag{1.26}
\end{equation*}
$$

where $\mathcal{L}^{\gamma}$ is the electrodynamic (QED) Lagrangian (interaction part) and is given by

$$
\begin{equation*}
\mathcal{L}^{\gamma}=-e j_{\gamma, L}^{\mu} A_{\mu} \tag{1.27}
\end{equation*}
$$

where the leptonic electromagnetic current $j_{\gamma, L}^{\mu}$ is

$$
\begin{equation*}
j_{\gamma, L}^{\mu}=-\bar{e} \gamma^{\mu} e \tag{1.28}
\end{equation*}
$$

The weak neutral current Lagrangian $\mathcal{L}^{Z}$ is given by

$$
\begin{equation*}
\mathcal{L}^{Z}=-\frac{g}{2 \cos \theta_{W}} j_{Z, L}^{\mu} Z_{\mu} \tag{1.29}
\end{equation*}
$$

where the leptonic weak neutral current is

$$
\begin{equation*}
j_{Z, L}^{\mu}=2 g_{L}^{\nu} \overline{\nu_{e L}} \gamma^{\mu} \nu_{e L}+2 \overline{g_{L}} \overline{e_{L}} \gamma^{\mu} e_{L}+2 g_{R} \overline{e_{R}} \gamma^{\mu} e_{R} . \tag{1.30}
\end{equation*}
$$

The coefficients $g_{L}^{\nu}, g_{L}$, and $g_{R}$ can be read off from Equation (1.21). Finally, the leptonic weak neutral current can be written as

$$
\begin{equation*}
j_{Z, L}^{\mu}=\overline{\nu_{e}} \gamma^{\mu}\left(g_{V}^{\nu}-g_{A}^{\nu} \gamma^{5}\right) \nu_{e}+\bar{e} \gamma^{\mu}\left(g_{V}^{l}-g_{A}^{l} \gamma^{5}\right) e \tag{1.31}
\end{equation*}
$$

where

$$
\begin{align*}
& g_{V}^{\nu, l}=g_{L}^{\nu, l}+g_{R}^{\nu, l}  \tag{1.32}\\
& g_{A}^{\nu, l}=g_{L}^{\nu, l}-g_{R}^{\nu, l} \tag{1.33}
\end{align*}
$$

Following the same procedure, one can deduce the Lagrangian for quarks. The corresponding interaction Lagrangian is

$$
\begin{align*}
\mathcal{L}_{I, Q} & =\mathcal{L}_{Q}^{(N C)}+\mathcal{L}_{Q}^{(C C)} \\
& =-e j_{\gamma, Q}^{\nu} A_{\nu}-\frac{g}{2 \cos \theta_{W}} j_{Z, Q}^{\nu} Z_{\nu}-\frac{g}{2 \sqrt{2}} j_{W, Q}^{\nu} W_{\nu}-\frac{g}{2 \sqrt{2}}\left(j_{W, Q}^{\nu}\right)^{\dagger} W_{\nu}^{\dagger} \tag{1.34}
\end{align*}
$$

The first two terms after the second equality sign constitute the neutral current Lagrangian while the last two terms constitute the charged current Lagrangian. The quark weak charged current $j_{W, Q}^{\nu}$, the quark weak neutral current $j_{Z, Q}^{\nu}$ and the quark electromagnetic current $j_{\gamma, Q}^{\nu}$ read, respectively,

$$
\begin{gather*}
j_{W, Q}^{\nu}=\bar{u} \gamma^{\nu}\left(1-\gamma_{5}\right) d,  \tag{1.35}\\
j_{Z, Q}^{\nu}=\bar{u} \gamma^{\nu}\left(g_{V}^{u}-g_{A}^{u}\right) u+\bar{d} \gamma^{\nu}\left(g_{V}^{d}-g_{A}^{d} \gamma_{5}\right) d,  \tag{1.36}\\
j_{\gamma, Q}^{\nu}=\frac{2}{3} \bar{u} \gamma^{\nu} u-\frac{1}{3} \bar{d} \gamma^{\nu} d . \tag{1.37}
\end{gather*}
$$

For coefficients $g_{V}^{u, d}$ and $g_{A}^{u, d}$, we have, in general,

$$
\begin{gather*}
g_{L}^{f}=I_{3}^{f}-q_{f} \sin ^{2} \theta_{W},  \tag{1.38}\\
g_{R}^{f}=-q_{f} \sin ^{2} \theta_{W},  \tag{1.39}\\
g_{V}^{f}=g_{L}^{f}+g_{R}^{f}=I_{3}^{f}-2 q_{f} \sin ^{2} \theta_{W},  \tag{1.40}\\
g_{A}^{f}=g_{L}^{f}-g_{R}^{f}=I_{3}^{f} . \tag{1.41}
\end{gather*}
$$

where the superscript $f$ denotes a specific fermion (lepton or quark) field. $I_{3}$ is the third component of the weak isospin and $q_{f}$ is the electrical charge of the corresponding fermion.

We have shown that the electroweak interaction includes the weak interaction which is inter-mediated by charged gauge fields $W^{ \pm}$and neutral gauge fields $Z$ (Feynman diagrams of these two interactions are shown in Figure 1.2), and the electromagnetic interaction which is inter-mediated by photon fields.


Figure 1.2: Diagrams of Weak Charged and Neutral Current

The weak interaction is a short range interaction, because the gauge bosons are massive. Roughly, we have $R_{\text {weak }}$ (Range of Weak Interaction)

$$
\begin{equation*}
R_{w e a k}=\frac{\hbar c}{M_{W} c^{2}} \approx 2 \times 10^{-18} \mathrm{~m}=0.002 \mathrm{fm} \ll 0.1 \mathrm{fm} \tag{1.42}
\end{equation*}
$$

## Weak Current of Hadrons and the CVC and PCAC

 hypothesisThe previous subsection gives the weak currents for leptons and quarks. In the real world, quarks never show up as isolated particles, instead only hadrons which are composed with quarks (as well as gluons) appear in experiments. For the coherent pion and rho processes, we encounter weak current for nucleons.

In the quark model proton and neutron are the states of $|u u d\rangle$ and $|u d d\rangle$ respectively. The weak interaction of these nucleons has more complicated structures compared to that of leptons or bare quarks, because it suffers from other interactions (mainly the strong interaction). However, the weak current of nucleons still consists of a vector current $V_{\alpha}$ and an axial-vector current $A_{\alpha}$ according to Lorentz invariance of the theory. So we may write

$$
\begin{equation*}
\left[J_{\text {hadron }}^{\text {weak }}\right]_{\alpha}=V_{\alpha}-A_{\alpha} . \tag{1.43}
\end{equation*}
$$

$V_{\alpha}$ and $A_{\alpha}$ can be expressed as suitable Dirac matrices sandwiched between spinors for nucleons. To specify, the vector current may consist of a $\gamma_{\alpha}$ term (corresponding to a vector form factor or point-like interaction), a $\sigma_{\alpha \beta} q^{\beta}$ term (called weak magnetism;
$q^{\beta}$ is the four-momentum transferred and $\left.\sigma_{\alpha \beta}=\frac{1}{2}\left[\gamma_{\alpha} \gamma_{\beta}-\gamma_{\beta} \gamma_{\alpha}\right]\right)$, and a $q^{\alpha}$ term (corresponding to an induced scalar form factor). The axial current may consist of a $\gamma_{\alpha} \gamma_{5}$ term (corresponding to an axial-vector form factor or point-like interaction), a $\sigma_{\alpha \beta} q^{\beta} \gamma_{5}$ term (pseudotensor form factor), and a $q_{\alpha} \gamma_{5}$ term (induced pseudoscalar form factor). Form factors for each term are dependent on the type of hadrons involved and this is expressed as $h_{1}$ and $h_{2}$ indices.

$$
\begin{align*}
V_{\alpha} & =\bar{\psi}_{h_{1}}\left[\gamma_{\alpha} f_{1}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)-i \sigma_{\alpha \beta} q^{\beta} f_{2}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)+q^{\alpha} f_{3}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)\right] \psi_{h_{2}}, \\
A_{\alpha} & =\bar{\psi}_{h_{1}}\left[\gamma_{\alpha} g_{1}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)-i \sigma_{\alpha \beta} q^{\beta} g_{2}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)+q^{\alpha} g_{3}^{h_{1} h_{2}}\left(Q^{2}, \nu\right)\right] \gamma_{5} \psi_{h_{2}}, \tag{1.44}
\end{align*}
$$

where $\psi_{h_{1}\left(h_{2}\right)}$ represents the spinor of nucleon $h_{1}\left(h_{2}\right)$ [51].
All the form factors $f_{i}$ and $g_{i}$ are real according to the time reversal invariance of the strong interaction [27]. Furthermore, in the exact isospin invariance, one finds

$$
\begin{equation*}
f_{3}\left(Q^{2}\right)=0, \quad g_{2}\left(Q^{2}\right)=0 \tag{1.45}
\end{equation*}
$$

As a consequence, the contraction of vector current $V_{\alpha}$ with $q^{\alpha}$ vanishes, i.e., $q^{\alpha} V_{\alpha}=$ 0 . The $f_{1}$ term vanishes due to the free Dirac equation for the nucleon spinors (remember that in the isospin symmetric-case, the proton and neutron have the same mass). The $f_{2}$ term vanishes due to the anti-symmetric property of $\sigma_{\alpha \beta}$. On the other hand, $q^{\alpha} V_{\alpha}=0$ implies $\partial^{\alpha} V_{\alpha}=0$ in the configuration space, i.e., the vector current is conserved in the case that the strong interaction is isospin invariant. This property was proposed in Ref. [25, 28] and is called the conserved vector current (CVC) hypothesis.

Originally, the CVC hypothesis was proposed to explain why the vector current is not renormalized by the strong interaction in the $\beta$ decay experiment. In the experiment, the nucleons have very small momentum transfer ( $Q^{2} \approx 0$ ). In this region all the kinematic terms (those which include $q_{\alpha}$ ) become negligible and the
weak current takes a form similar to that at the leptonic vertex

$$
\begin{align*}
{\left[J_{\text {hadron }}^{\text {weak }}\right]_{\alpha}\left(Q^{2}=0\right) } & =V_{\alpha}-A_{\alpha} \\
& =\bar{\psi}_{p} \gamma_{\alpha}\left(f_{1}^{p n}\left(Q^{2}=0\right)-g_{1}^{p n}\left(Q^{2}=0\right) \gamma_{5}\right) \psi_{n}, \tag{1.46}
\end{align*}
$$

and by definition

$$
\begin{equation*}
f_{1}(0)=g_{V}, \quad g_{1}(0)=g_{A} . \tag{1.47}
\end{equation*}
$$

Experiments tell us that $g_{V} \simeq 1$, which implies that the coupling of this hadronic current is consistent with that of the corresponding leptonic current. Analogously to the conservation of electrical current leading to the universality of electron/proton charge, this universality of weak vector charge implies a conservation of the weak vector current. It should be mentioned that the isospin is not an exact symmetry, but is violated by the electromagnetic interactions, responsible for the mass difference of $u$ and $d$ quarks (thus proton and neutron), so the vector current is not exactly conserved. However, the violation of isospin symmetry is very small, so the CVC hypothesis is a good approximation.

Now, we consider the axial current. Experiments tell us that $g_{A} \simeq 1.25$, which is different from unity. This implies that the axial-vector current is affected by strong interactions and is not conserved. (In fact, if the axial current were conserved, the commonly observed pion decay $\pi^{ \pm} \rightarrow \mu^{ \pm} \nu$ would be forbidden.) For this non-conserved axial current, there is another hypothesis, called partially conserved axial current (PCAC) hypothesis, that relates the derivative of the axial current to the pion field, which was first proposed in Ref. [26]. We will give more details about PCAC in the next chapter.

## The Goldberger-Treiman Relation

The CVC is a consequence of isospin invariance of the underlying theory, which arises when the quarks' masses are equal. When the quarks masses are zero, another
symmetry arises, that is the chiral $S U(2)$ symmetry. This chiral symmetry would result in the conservation of the axial current, so that $q^{\alpha} A_{\alpha}=0$. Now consider the axial current of neutron-proton transition. From Equation (1.44), we have, for the $g_{1}$ term:

$$
\begin{align*}
q^{\mu} A_{\mu}^{g_{1}} & =g_{1}\left(q^{2}\right) \bar{p}\left(\not p-\not p p^{\prime}\right) \gamma_{5} n \\
& =-g_{1}\left(q^{2}\right) \bar{p} \not p^{\prime \prime} \gamma_{5} n-g_{1}\left(q^{2}\right) \bar{p} \gamma_{5} \not p n \\
& =-\left(m_{p}+m_{n}\right) g_{1}\left(q^{2}\right) \bar{p} \gamma_{5} n, \tag{1.48}
\end{align*}
$$

and for the $g_{3}$ term:

$$
\begin{equation*}
q^{\mu} A_{\mu}^{g_{3}}=\bar{p} q^{2} g_{3} \gamma_{5} n . \tag{1.49}
\end{equation*}
$$

The $g_{2}$ term vanishes automatically. Then we have

$$
\begin{equation*}
\left(m_{p}+m_{n}\right) g_{1}\left(q^{2}\right)=q^{2} g_{3}\left(q^{2}\right) \tag{1.50}
\end{equation*}
$$

From this equation, we see that, either $g_{1}$ goes to zero as $q^{2} \rightarrow 0$ or $g_{3}\left(q^{2}\right)$ has a pole. From experiments, we have known that $g_{1}(0)=g_{A} \simeq 1.25$, so $g_{3}\left(q^{2}\right)$ should have a pole at $q^{2} \rightarrow 0$. This pole is due to the interaction of pions with the weak current of hadrons, which gives a contribution

$$
\begin{equation*}
\frac{q^{\mu}}{q^{2}} f_{\pi} \sqrt{2} g_{\pi N N} \bar{p} \gamma_{5} n \tag{1.51}
\end{equation*}
$$

to the matrix element of the axial vector current.
Comparing Equation (1.50) and Equation (1.51), we obtain

$$
\begin{equation*}
m_{N} g_{A}=\frac{g_{\pi} g_{\pi N N}}{\sqrt{2}} \tag{1.52}
\end{equation*}
$$

where the proton mass and neutron mass are taken as the same and $m_{N}=m_{p}=m_{n}$. This is the Goldberger-Treiman relation [31].

## Vector Meson Dominance Model(VMD) and Hadron Dominance Model(HDM)

The discussion of the previous subsection tells us that the axial vector current of hadrons receives contributions from interactions with pions, and sheds light on how to treat the hadron currents. For example, the electrical current of hadrons may be treated with the so called the Vector Meson Dominance Model (VMD) (details of this model can be found in Refs. $[13,32]$ ) where the amplitude of a photon scattering off a hadron is obtained by summing over the amplitudes of dominant vector mesonhadron scatterings (multiplied by corresponding meson propagators). Specifically one has

$$
\begin{equation*}
\mathcal{M}(\gamma+\alpha \rightarrow \beta)=\sum_{V=\rho^{0}, \omega, \phi} \frac{e}{g_{V}} \frac{m_{V}^{2}}{Q^{2}+m_{V}^{2}} \mathcal{M}(V+\alpha \rightarrow \beta), \tag{1.53}
\end{equation*}
$$

where $m_{V}$ is the mass of corresponding meson and $e \frac{m_{V}^{2}}{g_{V}}$ is the coupling constant of the vector meson to the photon. The foundation of this model is the observation that the photon-induced processes exhibit hadronic properties as shown in Figure 1.3. The total cross-sections of photon-induced processes and pion-nucleon scatterings have similar resonance structures at low energy, but become structureless at higher energy.

On the theoretical side, the VMD model assumes the photon state $\mid \gamma>$ is the superposition of a "bare" photon state $\left(\mid \gamma_{B}>\right)$ and a sum of hadronic states with the same quantum numbers as the photon $\left(J^{P C}=1^{--}, Q=B=S=0\right)$. Then the state vector could be written as

$$
\begin{equation*}
\left|\gamma>\simeq \sqrt{Z_{3}}\right| \gamma_{B}>+\sqrt{\alpha_{e m}} \mid h> \tag{1.54}
\end{equation*}
$$

where $Z_{3}$ is the renormalization constant of photon factor, and $\alpha_{e m}=\frac{1}{137}$ is the fine structure constant. In the simplest case, one picks up the three lightest vector mesons and then obtains Equation (1.53).


Figure 1.3: Total cross-section as a function of energy for interactions of photons and hadrons [50].

The idea of the Vector Meson Dominance Model can be generalized to the weak interacting processes, those via exchanging $W^{ \pm}$. This leads to the Hadron Dominance Model(HDM) [42]. Unlike the VMD model, which only needs contributions from vector mesons, in this model, the vector meson $(\rho)$ contributes to the hadronic vector current and the axial-vector meson $\left(a_{1}\right)$ contributes to the hadronic axial current. In addition, there is a contribution described by PCAC hypothesis. To specify, we have [50]

$$
\begin{align*}
& \sigma\left(\nu_{l}+\alpha \rightarrow l+\beta\right) \propto \sum_{i=\rho, a_{1}, \pi} \sigma(i+\alpha \rightarrow \beta) \\
&+ \text { interference }  \tag{1.55}\\
& \text { terms. }
\end{align*}
$$

## Chapter 2

## Challenges of Precision Oscillation

 ExperimentsIn the past, there have been many detectors designed for neutrino experiments. Neutrino beams are produced at particle accelerators which offer the greatest control over the neutrino being studied.

## Short Baseline Neutrino Experiment:

Experiments at a short baseline (defined as $\lesssim 1 \mathrm{~km}$ ) look at neutrino beams with energy ranging from stopped Muon decay $(<53 \mathrm{MeV})$ to several hundred GeV .

For example, in the LSND experiment [49], the signal $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$ was found [9, 11, 23] as well as $\nu_{\mu} \rightarrow \nu_{e}$ although weaker than the $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}[22,10]$. The MiniBooNE experiment $[19,14]$ is also a short baseline neutrino experiment is designed to check the result of LNSD signal.

## Long Baseline Neutrino Experiment:

The Long Baseline Neutrino Experiment (LBNE), now called the Deep Underground Neutrino Experiment (DUNE), was designed for a high sensitivity measurement to many parameters [1]:
(1) First, it is the measurement of $\nu_{\mu} \rightarrow \nu_{e}$ and $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$ oscillations with $\nu_{e}\left(\bar{\nu}_{e}\right)$ appearance and $\nu_{\mu}\left(\bar{\nu}_{\mu}\right)$ disappearance, including a precision measurement of the third mixing angle $\theta_{13}$, measurement of the CP violation phase $\delta_{C P}$, and determination of the mass hierarchy (MH) (the sign of $\Delta m_{13}^{2}$ ).

The experimental sensitivity, quantified by $\Delta \chi^{2}$ parameters, is calculated by com-


Figure 2.1: The significance with which the mass hierarchy (top) and CP violation ( $\delta_{C P} \neq 0$ or $\pi$, bottom) can be determined by the LBNE experiment with 34 -kt far detector as a function of the value of $\delta_{C P}$. The plots on the left are for normal hierarchy and the plots on the right are for inverted hierarchy. The width of the red band shows the range of the sensitivity that is achieved by LBNE when varying the beam design and the signal and background uncertainties [1]
paring the predicted spectra for various scenarios. As the sensitivity, the definitions of $\Delta \chi^{2}$ for neutrino MH and CP-violation sensitivity are different [1]:

$$
\begin{equation*}
\Delta \chi_{M H}^{2}=\left|\chi_{M H^{\text {test }}=I H}^{2}-\chi_{M H^{\text {test }}=N H}^{2}\right| \tag{2.1}
\end{equation*}
$$

The sensitivity of the mass hierarchy is defined in Equation 2.1. Because the oscillation experiments can only probe the squared difference of the neutrino masses, whether or not $m_{2}$ is heavier than $m_{3}$ remains unknown. If $m_{2}$ is lighter than $m_{3}$, this scenario is called normal mass hierarchy. If $m_{2}$ is heavier than $m_{3}$, this scenario
is called inverted mass hierarchy. Normal and inverted mass hierarchy are shown in Figure 2.2.

Figure 2.1 shows the sensitivities for determining the mass hierarchy (MH) and CP violation as a function of the true value of $\delta_{C P}$ after six years of running in the LBNE 34 kt configuration. Let's consider the sensitivity of neutrino MH first. The sensitivities of normal hierarchy (NH) and inverted hierarchy (IH) are evaluated separately and shown at the top of this figure. The X axis in this figure is the phase factor $\delta / \pi$. In the left top of this figure shows the sensitivity of normal hierarchy, There are three curves, the bottom one (solid line) is the result of prediction without near detector, the middle one is the result of prediction without the high resolution near detector but with beam improvement [1], the top one is the result of prediction with near detector and with beam improvement. We could see, with the beam design and high resolution near detector, the sensitivity is improved. The thicker the band is, the more the sensitivity improved.

The determination of the mass hierarchy (MH, or ordering) is very important in physics because neutrinos are fundamental particles, and knowing their properties is very critical [39]; The mass hierarchy could help people to know the nature of neutrinos: are they Majorana or Dirac particles? The question of CP violation could also be answered through the mass hierarchy, which is a very fundamental parameter of $\nu$ SM (SM +3 right-handed $\nu$ 's), and study mass hierarchy may provide the answers to why there is more matter than anti-matter and the connections between the neutrinos and dark matter. Let's see the relationship between neutrino mass hierarchy and CP violation. The uncertainty of the CP violation reads [1]

$$
\begin{gather*}
\Delta \chi_{C P V}^{2}=\min \left(\Delta \chi_{C P}^{2}\left(\delta_{C P}^{t e s t}=0\right), \Delta \chi_{C P}^{2}\left(\delta_{C P}^{t e s t}=\pi\right)\right)  \tag{2.2}\\
\Delta \chi_{C P}^{2}=\chi_{\delta_{C P}^{\text {test }}}^{2}-\chi_{\delta_{\text {true }}^{t e s t}}^{2} . \tag{2.3}
\end{gather*}
$$

Since the true value of $\delta_{C P}$ is unknown, a scan is performed over all possible values


Figure 2.2: Normal and inverted mass hierarchy [35].
of $\delta_{C P}^{\text {true }}$. The individual $\chi^{2}$ values are calculated using [1]

$$
\begin{equation*}
\chi^{2}\left(\mathbf{n}^{\text {true }}, \mathbf{n}^{\text {test }}, f\right)=2 \sum_{i}^{N_{\text {reco }}}\left(n_{i}^{\text {true }} \ln \frac{n_{i}^{\text {true }}}{n_{i}^{\text {test }}(f)}+n_{i}^{\text {test }}(f)-n_{i}^{\text {true }}\right)+f^{2}, \tag{2.4}
\end{equation*}
$$

where $\mathbf{n}$ are event rate vectors in $N_{\text {reco }}$ bins of reconstructed energy and f represents a nuisance parameter to be profiled. Nuisance parameters include the values of mixing angles, mass splittings and signal and background normalization. The nuisance parameters are constrained by Gaussian priors; in the case of the oscillation parameters, the Gaussian prior has standard deviation determined by taking $1 / 6$ of the $3 \sigma$ range allowed by the global fit [24]. The sensitivity of CP violation in the ELBNF/DUNE project is shown at the bottom of Figure 2.1, and is similar to the mass hierarchy sensitivity. The sensitivity to the CP violation is also improved using a high resolution near detector and with beam improvements. When $\delta$ is around $-0.5 \pi$ or $0.5 \pi$,
the sensitivity is increased from $3 \sigma$ level to $5 \sigma$ level. The inverted mass hierarchy scenario gives similar result as the normal mass hierarchy scenario but with higher sensitivity.


Figure 2.3: The expected reconstructed neutrino energy spectrum of $\nu_{\mu}$ or $\bar{\nu}_{\mu}$ events in a 34 -kt LArTPC for three years of neutrino(left) and anti-neutrino(right) running with a 1.2 -MW beam [1].
(2) The ELBNF/DUNE is also designed for the precision measurement of $\sin ^{2} \theta_{23}$ and $\left|\Delta m_{32}^{2}\right|$ in the $\nu_{\mu} / \bar{\nu}_{\mu}$ disappearance channel. Figure 2.3 and Figure 2.4 show the predicted spectrum of neutrino and anti-neutrino.

In Figure 2.5, the distributions of MH and CP-violation sensitivities as a function of exposure are shown. Similar to the Figure 2.1, there are also three set of lines are shown in these two figures. The top one is the result of $1 \% 5 \%$ which is the goal of the ELBNF/DUNE scientific program, with no systematic; The middle line with the value $2 \% 5 \%$ represents the results with beam improvement, but no near detector; The bottom line represents the results with no beam improvement and no near detector existence. From this figure, we can see an obvious improvement in sensitivity with beam improvements and a near detector. From this figure, it is obvious that the signal and background normalization uncertainties on the MH sensitivity is small, even at high exposures, given the large $\nu / \bar{\nu}$ asymmetry at $1,300 \mathrm{~km}$ compared to CP


Figure 2.4: The expected reconstructed neutrino energy spectrum of $\nu_{e}$ or $\bar{\nu}_{e}$ oscillation events in a 34 -kt LArTPC for three years of neutrino (left) and anti-neutrino (right) running with a $1.2-\mathrm{MW}, 80-\mathrm{GeV}$ beam assuming $\sin ^{2}\left(2\left(2 \theta_{13}\right)=0.09\right.$. The plots on the top are for normal hierarchy and the plots on the bottom are for inverted hierarchy [1].
violation which is even significant at high exposure.
(3) ELBNF/DUNE project is also designed to determine the $\theta_{23}$ octant using combined precision measurements of the $\nu_{e} / \bar{\nu}_{e}$ appearance and $\nu_{\mu} / \bar{\nu}_{\mu}$ disappearance channels.
(4) Another goal of ELBNF/DUNE project is to search for nonstandard physics, which could manifest itself as a difference in the high-precision measurement of $\nu_{\mu}$ and $\bar{\nu}_{\mu}$ oscillations over long baselines.


Figure 2.5: The mass hierarchy (left) and CP violation (right) sensitivities as a function of exposure in kt• year, for true normal hierarchy. The band represents the range of signal and background normalization errors [1].

The massive high-resolution far detector of ELBNF/DUNE project is a 34 kt liquid argon time-projection chamber (LArTPC) deep ground locates at Sanford Laboratory at a 4850 foot depth, 1300 km from Fermilab, and will enable LBNE to significantly expand the search for proton decay as predicted by Grand Unified Theories, as well as study the dynamics of core-collapse supernovae through observation of their neutrino bursts, should any occur in our galaxy during LBNE's operating lifetime [1].

Besides the far detector, a high resolution near neutrino detector is also proposed for high precision measurements of the neutrino mass matrix, neutrino interactions, structure of nucleons/nuclei, and so on. This high resolution detector will be introduced in chapter 3.

## Absolute and Relative Flux Measurements

Relative Flux: Any neutrino experiment needs to predict or measure the neutrino and anti-neutrino flux. In the ELBNF/DUNE project, the primary interest is the relative flux determination [1].

The relative neutrino flux could be measured through the charged current neutrino
scattering process $\left(\nu_{\mu}+e^{-} \rightarrow \mu^{-}+\nu_{e}\right)$ with low hadronic-energy deposition ( $\nu$ ) on the electron. The threshold of this process is $E_{\nu}>10.8 \mathrm{GeV}$. Because of the existence of the neutrino in the final state, the incoming neutrino energy could not be fully reconstructed. The intranuclear effects tell us that not all the hadrons escape from the nucleus, which will also affect the measurement of the visible energy of hadronic system. To minimize the fraction of the total interaction energy carried by the hadronic system, low $\nu_{0}$ is brought up, where $\nu_{0}$ is a given value of visible hadronic energy in the interaction. Then the differential cross-section could be expressed with $\nu_{0}[1]:$

$$
\begin{equation*}
\mathcal{N}\left(\nu<\nu_{0}\right) \simeq C \Phi\left(E_{\nu}\right) \nu_{0}\left[\mathcal{A}+\left(\frac{\nu_{0}}{E_{\nu}}\right) \mathcal{B}+\left(\frac{\nu_{0}}{E_{\nu}}\right)^{2} \mathcal{C}+\mathcal{O}\left(\frac{\nu_{0}}{E_{\nu}}\right)^{3}\right] \tag{2.5}
\end{equation*}
$$

where the coefficients are $\mathcal{A}=\mathcal{F}_{2}, \mathcal{B}=\left(\mathcal{F}_{2} \pm \mathcal{F}_{3}\right) / 2, \mathcal{C}=\left(\mathcal{F}_{2}+\mathcal{F}_{3}\right) / 6$, and $\mathcal{F}_{i}=$ $\int_{0}^{2} \int_{0}^{\nu_{0}} F_{i}(x) d x d \nu$ is the integral of the structure function $F_{i}(x)$. The dynamics of neutrino-nucleon scattering implies that the number of events in a given energy bin with hadronic energy $E_{h a d}<\nu_{0}$ is proportional to the (anti) neutrino flux in that energy bin up to corrections $\mathcal{O}\left(\nu_{0} / E_{\nu}\right)$ and $\mathcal{O}\left(\nu_{0} / E_{\nu}\right)^{0}$. The number of $\mathcal{N}\left(\nu<\nu_{0}\right)$ is therefore proportional to the flux, up to correction factors of the order $\mathcal{O}\left(\nu_{0} / E_{\nu}\right)$ or smaller, which are not significant for small values of $\nu_{0}$ at energies $\geq \nu_{0}$. The coefficients $\mathcal{A}$ and $\mathcal{B}$ and $\mathcal{C}$ are determined for each energy bin and neutrino flavor within the ND data.

## Low Energy Absolute Flux: Neutrino-Electron-NC Scattering

The low energy absolute flux could be determined by the neutral current neutrino scattering on the electron $\left(\nu_{\mu}+e^{-} \rightarrow \nu_{\mu}+e^{-}\right)$[40].

$$
\begin{align*}
& \sigma\left(\nu_{l} e \rightarrow \nu_{l} e\right)=\frac{G_{\mu}^{2} m_{e} E_{\nu}}{2 \pi}\left[1-4 \sin ^{2} \theta_{W}+\frac{16}{3} \sin ^{4} \theta_{W}\right],  \tag{2.6}\\
& \sigma\left(\bar{\nu}_{l} e \rightarrow \bar{\nu}_{l} e\right)=\frac{G_{\mu}^{2} m_{e} E_{\nu}}{2 \pi}\left[\frac{1}{3}-\frac{4}{3} \sin ^{2} \theta_{W}+\frac{16}{3} \sin ^{4} \theta_{W}\right], \tag{2.7}
\end{align*}
$$

From Equation 2.6 and Equation 2.7, we could see the cross-section only depends on the knowledge of $\sin ^{2} \theta_{W}$.

## High Energy Absolute Flux: Neutrino-Electron CC scattering

The $\nu_{\mu^{-}} e^{-}$CC interaction (inverse Muon decay(IMD): $\nu_{\mu}+e^{-} \rightarrow \mu^{-}+\nu_{e}$ ) could be used to determine the high energy absolute flux. Considering the energy threshold for this process, IMD requires $E_{\nu} \geq 10.8 \mathrm{GeV}$. The high-resolution ND in the ELBNF neutrino beam will observe $\geq 2,000$ IMD events within three years. The reconstruction efficiency of the single, energetic forward $\mu^{-}$will be $\geq 98 \%$; the angular resolution of the IMD $\mu$ is $\leq 1 \mathrm{mrad}$ [1]. The background, primarily originated from the $\nu_{\mu}-\mathrm{QE}$ interactions, can be precisely constrained using a control sample [1].

## Low Energy Absolute Flux: QE in Water and Heavy-Water Targets

The quasi-elastic charged current (QE-CC) scattering $\left(\nu_{\mu} n(p) \rightarrow \mu^{-} p(n)\right)$ on deuterium at low $Q^{2}$ could be used to extract the low energy absolute neutrino flux. Since $\left(m_{\mu} / M_{n}\right)^{2}$ at $Q^{2}=0$, could be neglected, the cross-section is independent of neutrino energy for $\left(2 E_{\nu} M_{n}\right)^{1 / 2}>m_{\mu}[1]$ :

$$
\begin{equation*}
\frac{d \sigma}{d Q^{2}}\left|Q^{2}=0\right|=\frac{G_{\mu}^{2} \cos ^{2} \theta_{C}}{2 \pi}\left[F_{1}^{2}(0)+G_{A}^{2}(0)\right]=2.08 \times 10^{-38} \mathrm{~cm}^{2} \mathrm{GeV}^{-2} \tag{2.8}
\end{equation*}
$$

which is determined by neutron $\beta$ decay and has a theoretical uncertainty $<1 \%$. The flux can be extracted experimentally by measuring low $Q^{2}$ QE interactions ( $\leq$ 0.05 GeV ) and extrapolating the result to the limit of $Q^{2}=0$. The measurement requires a deuterium(or hydrogen for anti-neutrino) target to minimize the smearing due to Fermi motion and other nuclear effects. This requirement can only be achieved by using both $\mathrm{H}_{2} \mathrm{O}$ and $\mathrm{D}_{2} \mathrm{O}$ targets embedded in the fine-grained tracker and extracting the events produced in deuterium by statistical subtraction of larger oxygen component [1].


Figure 2.6: $\nu_{\mu}$ and $\bar{\nu}_{\mu}$ flux distribution in ELBNF/DUNE project.


Figure 2.7: $\nu_{e}$ and $\bar{\nu}_{e}$ flux distribution in ELBNF/DUNE project.

## Measurement of $\nu / \bar{\nu}$ Flux Ratio

Figure 2.6 and Figure 2.7 show the distribution of neutrino fluxes at the near detector in predicted ELBNF/DUNE projects. In the ELBNF neutrino beam, there will have $a<10 \%$ contamination of neutrinos of the "wrong sign" in the oscillation energy region ( $\bar{\nu}$ 's in the $\nu$ beam and vice-versa) from the decay of wrong sign hadrons that propagate down the center of the focusing horns (where there is no magnetic field) into the decay volume. $\mathrm{A} \leq 1 \%$ contamination of $\nu_{e}$ and $\bar{\nu}_{e}$ in the $\nu_{e}$ appearance signal region is produced by the decays of tertiary Muons from Pion decays, and decays of Kaons [1].

The measurement of the neutrino fluxes is very important in the neutrino inter-
action analysis for the following reasons:
Similar to the NOMAD Beam, The LBNF fluxes also have two beam modes: neutrino beam mode and anti-neutrino beam mode. The neutrino beam mode is dominated by $\nu$, and the anti-neutrino beam mode is dominated by $\bar{\nu}$. However, there are still a contamination in each beam mode. It is very important to constrain the contamination of each beam.

Second, measuring these fluxes is important to perform procedure measurement of cross sections and electroweak measurement, such weak mixing angle.

Third, since the coherent process is leaving the target nucleus largely unaffected, it is optimum to monitor the neutrino source.

Fourth, Since the ratio of coherent cross section for $\nu$ and $\bar{\nu}$ can be calculated precisely, it is possible to constrain the critical ratio of $\nu / \bar{\nu}$ fluxes.

Finally, the neutrino and anti-neutrino induced coherent Pion production have the same cross-section under PCAC theorem, so we can measure the neutrino vs antineutrino flux precisely which is very important for LBNF/DUNE oscillation measurements.

## Determination of $\nu(\bar{\nu})$ energy scale

In ELBNF/DUNE, the determination of the oscillation parameters depends on the knowledge of the neutrino energy. Thus, reconstructing the neutrino energy becomes very critical in the experiment. It must be reconstructed on an event-by event basis because the neutrino beams are quite broad in energy due to their production in a secondary decay of primarily produced hadrons. The determination of neutrino energy could be measured from the kinematics of the outgoing particles, Since the coherent process is not as sensitive to nuclear effects as other kind of processes, then study of coherent processes is very important in determining the neutrino spectrum.

## Chapter 3

## Coherent Meson Production by Neutrino (Theory)

### 3.1 Kinematics of (Anti)Neutrino Scattering

Before we proceed let's define some notation used throughout this chapter (see Figure 3.1). (The derivation in the section was created in collaboration with C. T. Kullenberg)
$\mathbf{E}$ and $\mathbf{E}^{\prime}$ : The incoming neutrino and outgoing lepton energies.
$\mathbf{k}$ and $\mathbf{k}^{\prime}$ : The four-momenta of the incoming neutrino and outgoing lepton.
$\mathbf{p}$ and $\mathbf{p}{ }^{\prime}$ : The four-momenta of initial and final hadron states.
$\mathbf{q}$ : The four-momentum transfer.
$\boldsymbol{\nu}$ : The energy transfer.
$\mathbf{M}$ : The mass of the target nucleon.
$\mathrm{Q}^{2}$ : The negative of the square of the four-momentum transfer.
The mathematical expressions of $\boldsymbol{\nu}, \mathbf{q}$ and $\mathbf{Q}^{\mathbf{2}}$ are shown in Equation (3.1),

$$
\begin{align*}
\nu & =E-E^{\prime} \\
q & =k-k^{\prime}=\left(E-E^{\prime}, \vec{k}-\vec{k}^{\prime}\right)=(\nu, \vec{q}) \\
Q^{2} & =-q^{2}=|\vec{q}|^{2}-\nu^{2} \tag{3.1}
\end{align*}
$$

Looking more carefully at $Q^{2}$ we have:

$$
\begin{equation*}
Q^{2}=2\left(E E^{\prime}-\vec{k} \cdot \vec{k}^{\prime}\right)-m_{\nu}^{2}-m_{l}^{2} \tag{3.2}
\end{equation*}
$$

$W^{2}$ : Invariant hadronic mass $W^{2}=(q+p)^{2}$.
If we assume that the lepton masses are negligible $\left(m_{\nu}, m_{l} \approx 0\right)$ then $\vec{k} \cdot \vec{k}^{\prime}=$ $|\vec{k}|\left|\overrightarrow{k^{\prime}}\right| \cos \theta \approx E E^{\prime} \cos \theta$ and we have:


Figure 3.1: Kinematics of neutrino scattering.

### 3.2 Neutrino Induced Coherent Pion

According to the Feymann Rules, the amplitude of a neutrino induced hadronic interaction can be obtained from the weak current at the leptonic and hadronic vertexes:

$$
\begin{equation*}
\mathcal{M}=\frac{G_{F}}{\sqrt{2}}\left[J_{\text {lepton }}^{\text {weak }}\right]^{\alpha}\left[J_{\text {hadron }}^{\text {weak }}\right]_{\alpha}, \tag{3.4}
\end{equation*}
$$

where

$$
\begin{equation*}
\frac{G_{F}}{\sqrt{2}}=\frac{g^{2}}{8 M_{W}^{2}} . \tag{3.5}
\end{equation*}
$$

The form of the weak current in purely leptonic interactions is known as:

$$
\begin{equation*}
\left[J_{l e p t o n}^{w e a k}\right]_{\alpha}=\bar{\psi}_{l} \gamma_{\alpha}\left(1-\gamma_{5}\right) \psi_{\nu} . \tag{3.6}
\end{equation*}
$$

The exact form of the weak current is known for free quark interactions; however, it is more complicated in the case of bound quarks in nucleons, which can be written as

$$
\begin{equation*}
\left[J_{\text {hadron }}^{\text {weak }}\right]_{\alpha}=V_{\alpha}-A_{\alpha} . \tag{3.7}
\end{equation*}
$$

While $V_{\alpha}$ transforms like a vector under parity, $A_{\alpha}$ transforms like an axial vector. Combining the leptonic current and hadronic current, we have:

$$
\begin{equation*}
\mathcal{M}=\frac{G_{F}}{\sqrt{2}} \mathcal{L}^{\alpha}\left[V_{\alpha}-A_{\alpha}\right] \tag{3.8}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{L}^{\alpha}=\bar{\psi}_{l} \gamma^{\alpha}\left(1-\gamma^{5}\right) \psi_{\nu} . \tag{3.9}
\end{equation*}
$$

The scattering amplitude can be written as:

$$
\begin{equation*}
\mathcal{M}^{2}=\frac{G_{F}^{2}}{2} \mathcal{L}^{\alpha} \mathcal{L}^{\beta *}\left[V_{\alpha}+A_{\alpha}\right]\left[V_{\beta}^{*}+A_{\beta}^{*}\right] \tag{3.10}
\end{equation*}
$$

and

$$
\begin{equation*}
A_{\alpha \beta}=\left[V_{\alpha}-A_{\alpha}\right]\left[V_{\beta}^{*}-A_{\beta}^{*}\right] . \tag{3.11}
\end{equation*}
$$

Besides the leptonic tensor and the hadronic tensor, in some conditions, the interference between the vector and axial vector terms is also considered. However, this term vanishes in the case of coherent interactions in the limit $Q^{2} \rightarrow 0$. Here the leptonic tensor is defined as

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =\mathcal{L}^{\alpha} \mathcal{L}^{\beta *} \\
& =8\left\{p_{\nu}^{\alpha} p_{\mu}^{\beta}+p_{\nu}^{\beta} p_{\mu}^{\alpha}-p_{\nu} \cdot p_{\mu} g^{\alpha \beta}+i \epsilon^{\alpha \beta \gamma \delta} p_{\nu \gamma} p_{\mu_{\delta}}\right\} \tag{3.12}
\end{align*}
$$

Now let's show some details as follows. Using Equation (3.9), we have

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =\mathcal{L}^{\alpha} \mathcal{L}^{\beta *} \\
& =\left(\bar{\psi}_{l} \gamma^{\alpha}\left(1-\gamma_{5}\right) \psi_{\nu}\right)\left(\bar{\psi}_{l} \gamma^{\beta}\left(1-\gamma_{5}\right) \psi_{\nu}\right)^{\dagger} \\
& =\bar{\psi}_{l} \gamma^{\alpha}\left(1-\gamma_{5}\right) \psi_{\nu} \psi_{\nu}^{\dagger}\left(1-\gamma_{5}\right) \gamma^{\beta \dagger} \bar{\psi}_{l}^{\dagger} \tag{3.13}
\end{align*}
$$

Using the anti-communicative relationship and the formula of the sum over the spin of Dirac field:

$$
\begin{gather*}
\left\{\gamma_{5}, \gamma_{\mu}\right\}=0 .  \tag{3.14}\\
\sum_{s} \psi_{l} \bar{\psi}_{l}=\not p_{l}+m . \tag{3.15}
\end{gather*}
$$

In the present analysis, the neutrino mass can be safely taken to be 0 , so we have:

$$
\begin{equation*}
\sum_{s} \psi_{\nu} \bar{\psi}_{\nu}=\not p_{\nu} \tag{3.16}
\end{equation*}
$$

The leptonic tensor now becomes

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =\operatorname{Tr}\left(\psi_{l} \bar{\psi}_{l} \gamma^{\alpha}\left(1-\gamma_{5}\right) \psi_{\nu} \bar{\psi}_{\nu}\left(1+\gamma_{5}\right) \gamma^{\beta}\right) \\
& =\operatorname{Tr}\left(\left(\not p_{\mu}+m\right) \gamma^{\alpha}\left(1-\gamma_{5}\right) \not p_{\nu}\left(1+\gamma_{5}\right) \gamma^{\beta}\right) . \tag{3.17}
\end{align*}
$$

The subscripts $\nu$ and $\mu$ indicate the neutrino and the muon, not the four vector index.

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =\operatorname{Tr}\left(\psi_{l} \bar{\psi}_{l} \gamma^{\alpha}\left(1-\gamma_{5}\right) \psi_{\nu} \bar{\psi}_{\nu}\left(1+\gamma_{5}\right) \gamma^{\beta}\right) \\
& =\operatorname{Tr}\left(\left(\not p_{\mu}+m\right) \gamma^{\alpha}\left(1-\gamma_{5}\right) \not p_{\nu}\left(1+\gamma_{5}\right) \gamma^{\beta}\right) \\
& =\operatorname{Tr}\left(\left(\not p_{\mu}+m\right) \gamma^{\alpha}\left(\not p_{\nu}-\gamma_{5} \not p_{\nu}\right)\left(1+\gamma_{5}\right) \gamma^{\beta}\right) \\
& =2 \operatorname{Tr}\left(\left(\not p_{\mu}+m\right) \gamma^{\alpha}\left(\not p_{\nu}+\not p_{\nu} \gamma_{5}\right) \gamma^{\beta}\right) \tag{3.18}
\end{align*}
$$

In the last equality, the anti-commutation relation has been used.

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =2 \operatorname{Tr}\left[\not p_{\mu} \gamma^{\alpha} \not p_{\nu} \gamma^{\beta}+m \gamma^{\alpha} \not p_{\nu} \gamma^{\beta}+\not{ }_{\mu} \gamma^{\alpha} \not p_{\nu} \gamma_{5} \gamma^{\beta}-m \gamma^{\alpha} \not p_{\nu} \gamma_{5} \gamma^{\beta}\right] \\
& =2 \operatorname{Tr}\left[\not p_{\mu} \gamma^{\alpha} \not p_{\nu} \gamma^{\beta}+\not p_{\mu} \gamma^{\alpha} \not p_{\nu} \gamma_{5} \gamma^{\beta}\right] . \tag{3.19}
\end{align*}
$$

Using the trace technology:

$$
\begin{gather*}
\operatorname{Tr}\left[\not p_{\mu} \gamma^{\alpha} \not{ }_{\nu} \gamma^{\beta}\right]=4 p_{\mu}^{\alpha} p_{\nu}^{\beta}-4 p_{\mu} \cdot p_{\nu} g^{\alpha \beta}+4 p_{\mu}^{\beta} p_{\nu}^{\alpha}  \tag{3.20}\\
\operatorname{Tr}\left[\not p_{\mu} \gamma^{\alpha} \not p_{\nu} \gamma_{5} \gamma^{\beta}\right]=-i \epsilon^{l \alpha m \beta}\left(p_{\mu}\right)_{l}\left(p_{\nu}\right)_{m} . \tag{3.21}
\end{gather*}
$$

The indices $\alpha \beta \gamma \delta$ are four momentum indices and are summed over. For small lepton masses $\left(m_{\mu} \rightarrow 0\right)$, it can be shown that $p_{\nu} \cdot p_{\mu}$ approaches $\frac{Q^{2}}{2}$.

Taking the limit $Q^{2} \rightarrow 0$ then causes the term $p_{\nu} \cdot p_{\mu}$ to vanish. From the equality $p_{\nu} \cdot p_{\mu}=E_{\nu} E_{\mu}(1-\cos \theta)$, again assuming negligible lepton masses, it is obvious that the limit $Q^{2} \rightarrow 0$ refers to the case where the neutrino and muon direction are parallel. In this limit the four momenta can be written $p_{\nu}=\left|E_{\nu} / \nu\right| q$ and $p_{\mu}=\left|E_{\mu} / \nu\right| q$, and the lepton tensor becomes:

$$
\begin{equation*}
\mathcal{L}^{\alpha \beta}=16 \frac{E_{\nu} E_{\mu}}{\nu^{2}} q^{\alpha} q^{\beta} . \tag{3.22}
\end{equation*}
$$

Using this tensor, the amplitude can be expressed in terms of the derivative of the hadronic weak currents. As the vector current is conserved (CVC), the derivative of the vector current $\partial^{\alpha} V_{\alpha}$ vanishes and only the derivative of the axial vector current remains. Hence the aforementioned conclusion that the axial vector current dominates at very low $Q^{2}$.

$$
\begin{equation*}
\mathcal{M}^{2}=8 G^{2} \frac{E_{\nu} E_{\mu}}{\nu^{2}}\left|\partial^{\alpha} A_{\alpha}\right|^{2} \tag{3.23}
\end{equation*}
$$

In 1960 Goldberger and Treiman used the possibility of almost conserving the axial vector current (in the limit of massless pions), by introducing a pseudoscalar current, to predict the rate of pion decay. This result in an expression for the axial -vector form factor, commonly known as the Goldberger-Treiman relation.

$$
\begin{equation*}
\left(M_{p}+M_{n}\right) g_{1}(0)=\sqrt{2} f_{\pi} g_{\pi N} \tag{3.24}
\end{equation*}
$$

Here $M_{p}$ and $M_{n}$ are the masses of the proton and neutron, $f_{\pi}$ is the pion decay constant, and $g_{\pi N}$ is the pion-nucleon coupling constant.

This leads to another statement of the PCAC (partially conserved axial current) hypothesis, first proposed by Gell-Mann and Levy in 1960:

$$
\begin{equation*}
\partial^{\alpha} A_{\alpha}=f_{\pi} M_{\pi}^{2} \phi_{\pi} . \tag{3.25}
\end{equation*}
$$

where $\phi_{\pi}$ is the pion field. Now let's go into some details of the PCAC hypothesis. First, we have

$$
\begin{align*}
<0\left|A^{\mu}(x)\right| \pi\left(p^{\nu}\right)> & =<0\left|\exp (i \hat{P} \cdot x) A^{\mu}(0) \exp (-i \hat{P} \cdot x)\right| \pi\left(p^{\nu}\right)> \\
<0\left|\partial_{\mu} A^{\mu}(x)\right| \pi\left(p^{\nu}\right)> & =\partial_{\mu}<0\left|A^{\mu}(x)\right| \pi\left(p^{\nu}\right)> \\
& =\partial_{\mu}<0\left|\exp (i \hat{P} \cdot x) A^{\mu}(0) \exp (-i \hat{P} \cdot x)\right| \pi\left(p^{\nu}\right)> \\
& =\partial_{\mu}<0\left|A^{\mu}(0)\right| \pi\left(p^{\nu}\right)>\exp (-i P \cdot x) \\
& =-i P_{\mu}<0\left|A^{\mu}(0)\right| \pi\left(p^{\nu}\right)>\exp (-i P \cdot x) \tag{3.26}
\end{align*}
$$

Since $<0\left|A^{\mu}(0)\right| \pi\left(p^{\nu}\right)>$ is an axial vector function of $p_{\mu}$, it can be written as

$$
\begin{equation*}
<0\left|A^{\mu}(0)\right| \pi\left(p^{\nu}\right)>=i f_{\pi} p^{\mu} \tag{3.27}
\end{equation*}
$$

where $f_{\pi}$ is the pion decay constant. So

$$
\begin{align*}
<0\left|\partial_{\mu} A^{\mu}(x)\right| \pi\left(p^{\nu}\right)> & =p_{\mu} \cdot p^{\mu} f_{\pi} \exp (-i P \cdot x) \\
& =m_{\pi}^{2} f_{\pi}<0\left|\phi_{\pi}(x)\right| \pi\left(p^{\nu}\right)> \tag{3.28}
\end{align*}
$$

This equation shows that PCAC is exact on the pion mass shell, while off the pion mass shell, it is assumed that the operator equation 3.25 still holds.

With CVC and PCAC, and neglecting the mass of the leading lepton (muon), now we can deal with the matrix element, which is $<\mathcal{B}\left|\partial_{\alpha} A^{\alpha}\right| \mathcal{A}>$, where $\mathcal{A}$ and $\mathcal{B}$ represent initial and final nucleon respectively.

Using Equation (3.25), we have

$$
\begin{align*}
<\mathcal{B}\left|\partial_{\mu} A^{\mu}\right| \mathcal{A}> & =<\mathcal{B}\left|f_{\pi} m_{\pi}^{2} \phi_{\pi}\right| \mathcal{A}> \\
& =f_{\pi} m_{\pi}^{2}<\mathcal{B}\left|\phi_{\pi}\right| \mathcal{A}> \tag{3.29}
\end{align*}
$$

The transition matrix element $<\mathcal{B}\left|\phi_{\pi}\right| \mathcal{A}>$ has a pole at the $q^{2}=m_{\pi}^{2}$ and its residue is just the amplitude of the scattering $\langle\mathcal{A}+\pi \rightarrow \mathcal{B}\rangle$. So it is reasonable to assume

$$
\begin{equation*}
<\mathcal{B}\left|\partial_{\mu} A^{\mu}\right| \mathcal{A}>=f_{\pi} \frac{m_{\pi}^{2}}{Q^{2}+m_{\pi}^{2}}<\mathcal{A}+\pi \rightarrow \mathcal{B}> \tag{3.30}
\end{equation*}
$$

at the low energy region $0 \lesssim q^{2} \lesssim m_{\pi}^{2}$. Precisely, there are also other pole contributions from other particles [16], and considering them leads to an extension of PCAC.

In the limit of $Q^{2} \rightarrow 0$, the result given above leads us to (Note that $i \partial^{\alpha} \leftrightarrow q^{\alpha}$ )

$$
\begin{align*}
q_{\alpha} A^{\alpha \beta} q_{\beta} & =q_{\alpha}<A^{\alpha}><A^{\beta}>^{*} q_{\beta} \\
& =f_{\pi}<\mathcal{B}\left|\mathcal{A}+\pi>f_{\pi}<\mathcal{B}\right| \mathcal{A}+\pi> \\
& =f_{\pi}^{2}|\mathcal{M}(\mathcal{A}+\pi \rightarrow \mathcal{B})|^{2} \tag{3.31}
\end{align*}
$$

Inserting Equation (3.31) into Equation (3.23) gives:

$$
\begin{equation*}
|\mathcal{M}|^{2}=8 G_{F}^{2} \frac{E_{\nu} E_{\mu}}{\nu^{2}} f_{\pi}^{2}|\mathcal{M}(\mathcal{A}+\pi \rightarrow \mathcal{B})|^{2} \tag{3.32}
\end{equation*}
$$

Then we have

$$
\begin{align*}
d \sigma(\nu+\mathcal{A} \rightarrow \mu+\mathcal{B})= & 8 G_{F}^{2} \frac{E_{\nu} E_{\mu}}{4 E_{\nu} M_{N} \nu^{2}} f_{\pi}^{2}|<\mathcal{M}(\mathcal{A}+\pi \rightarrow \mathcal{B}>)|^{2} \\
& \times(2 \pi)^{4} \delta^{4}\left(p_{\nu}+p_{Z i}-p_{\mu}-p_{Z f}\right) \\
& \times \frac{1}{(2 \pi)^{3}} \frac{d^{3} p_{\mu}}{2 E_{\mu}} \frac{1}{(2 \pi)^{3}} \frac{d^{3} p_{Z f}}{2 E_{Z f}} . \tag{3.33}
\end{align*}
$$

The integrating measure is

$$
\begin{equation*}
d^{3} p_{\mu}=\left|\overrightarrow{p_{\mu}}\right|^{2} d p_{\mu} d \Omega=\left|\vec{p}_{\mu}\right|^{2} d p_{\mu} 2 \pi d \cos \theta \tag{3.34}
\end{equation*}
$$

For the massless particle $\left|\overrightarrow{p_{\mu}}\right|=E_{\mu}, Q^{2}=2 E_{\nu} E_{\mu}(1-\cos \theta) ; d E_{\mu}=-d \nu ; d Q^{2}=$ $-2 E_{\nu} E_{\mu} d \cos \theta$. So

$$
\begin{align*}
d^{3} p_{\mu} & =E_{\mu}^{2} d E_{\mu} 2 \pi \frac{1}{-2 E_{\nu} E_{\mu}} d Q^{2} \\
& =E_{\mu}^{2}(-d \nu) 2 \pi \frac{1}{-2 E_{\nu} E_{\mu}} d Q^{2} \\
& =E_{\mu}^{2} 2 \pi \frac{1}{-2 E_{\nu} E_{\mu}} d Q^{2} d \nu \\
& =\frac{\pi E_{\mu}}{E_{\nu}} d Q^{2} d \nu \tag{3.35}
\end{align*}
$$

Now, we can write the scattering amplitude of $|\mathcal{M}(\mathcal{A}+\pi \rightarrow \mathcal{B})|^{2}$ into the form of cross-sections,

$$
\begin{align*}
d \sigma(\mathcal{A}+\pi \rightarrow \mathcal{B})= & \frac{|\mathcal{M}(\mathcal{A}+\pi \rightarrow \mathcal{B})|^{2}}{4 M_{N} \nu} \\
& \times(2 \pi)^{4} \delta^{4}\left(p_{\nu}+p_{Z i}-p_{\mu}-p_{Z f}\right) \frac{1}{(2 \pi)^{3}} \frac{d^{3} p_{Z f}}{2 E_{Z f}} . \tag{3.36}
\end{align*}
$$

Then the differential cross-section becomes

$$
\begin{equation*}
\left.d \sigma(\nu+\mathcal{A} \rightarrow \mu+\mathcal{B})\right|_{Q^{2} \rightarrow 0}=\frac{G_{F}^{2} f_{\pi}^{2}(1-y)}{2 \pi^{2} \nu} \sigma(\pi+\mathcal{A} \rightarrow \mathcal{B}) d Q^{2} d \nu \tag{3.37}
\end{equation*}
$$

This expression is known as Adler's theorem, it is valid for the limit $Q^{2} \rightarrow 0$; it relates the weak neutrino nucleus cross-section to that of the strong pion-nucleus cross-section. In the calculation above, the mass of the lepton muon is neglected. If the lepton mass is not neglected, the leptonic tensor becomes

$$
\begin{equation*}
\mathcal{L}_{\mu \nu}=8\left\{k_{\mu} k_{\nu}^{\prime}+k_{\nu} k_{\mu}^{\prime}+k \cdot k^{\prime} g_{\mu \nu}\right\} . \tag{3.38}
\end{equation*}
$$

Because the hadronic tensor is symmetric, the anti-symmetric tensor term in leptonic tensor has no contribution.

It is interesting that the cross-section is not dominated by exchanging the pion meson at $Q^{2}=0$ under PCAC. To show this, let's write the hadronic current into 2 parts, one of which represents the pionic contribution and the other one represents nonpionic contribution.

$$
\begin{align*}
j^{\mu} & =i f_{\pi} \frac{q^{\mu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M}(\alpha+\pi \rightarrow \beta)+j^{\prime \mu} .  \tag{3.39}\\
j^{\nu *} & =-i f_{\pi} \frac{q^{\nu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M}^{*}(\alpha+\pi \rightarrow \beta)+j^{\prime \nu *} . \tag{3.40}
\end{align*}
$$

The hadronic tensor becomes

$$
\begin{align*}
W^{\mu \nu}= & j^{\mu} j^{\nu *} \\
= & j^{\prime \mu} j^{\prime \nu *}+j^{\prime \mu} \times\left(-i f_{\pi} \frac{q^{\nu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M}^{*}\right)+j^{\prime \nu *} \times i f_{\pi} \frac{q^{\mu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M} \\
& +f_{\pi} \frac{q^{\mu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M} \times f_{\pi} \frac{q^{\nu}}{Q^{2}+m_{\pi}^{2}} \mathcal{M}^{*} . \tag{3.41}
\end{align*}
$$

$$
\begin{equation*}
i q_{\mu} j^{\mu}=f_{\pi} \frac{Q^{2}}{Q^{2}+m_{\pi}^{2}} \mathcal{M}+i q_{\mu} j^{\prime \mu} \tag{3.42}
\end{equation*}
$$

Using the result of PCAC,

$$
\begin{align*}
& i q_{\mu} j^{\mu}=f_{\pi} \mathcal{M}(\alpha+\pi \rightarrow \beta)  \tag{3.43}\\
& i q_{\mu} j^{\prime \mu}=f_{\pi} \frac{m_{\pi}^{2}}{Q^{2}+m_{\pi}^{2}} \mathcal{M} \tag{3.44}
\end{align*}
$$

Now let's calculate $\sum j^{\prime \mu} \mathcal{M}^{*}$. Assume

$$
\begin{equation*}
\sum j^{\prime \mu} \mathcal{M}^{*}=a p^{\mu}+b q^{\mu} \tag{3.45}
\end{equation*}
$$

where $a$ and $b$ are constants.

$$
\begin{align*}
i q_{\mu} \sum j^{\prime \mu} \mathcal{M}^{*} & =i q_{\mu}\left(a p^{\mu}+b q^{\mu}\right) \\
& =i a p \cdot q+i b q^{2} \tag{3.46}
\end{align*}
$$

When $Q^{2} \approx 0$,

$$
\begin{align*}
i q_{\mu} \sum j^{\mu} \mathcal{M}^{*} & =i q_{\mu}\left(a p^{\mu}+b q^{\mu}\right) \\
& =i a p \cdot q \tag{3.47}
\end{align*}
$$

Then, we have

$$
\begin{equation*}
a=\frac{f_{\pi}|\mathcal{M}|^{2}}{i p \cdot q} \tag{3.48}
\end{equation*}
$$

Combine with the leptonic tensor, and we get the cross-section:

$$
\begin{align*}
\left.\frac{d^{2} \sigma(\nu+A->\mu+\pi+B)}{d Q^{2} d \nu}\right|_{Q^{2} \rightarrow 0}= & \frac{G^{2} f_{\pi}^{2}(1-y)}{2 \pi^{2} \nu} \sigma(\pi+\mathcal{A} \rightarrow \mathcal{B}) \\
& \times\left[1-\frac{\nu}{E^{\prime}} \frac{m_{\mu}^{2}}{Q^{2}+m_{\pi}^{2}}+\frac{\nu^{2}}{4 E E^{\prime}} \frac{m_{\mu}^{2}\left(Q^{2}+m_{\mu}^{2}\right)}{\left(Q^{2}+m_{\pi}^{2}\right)^{2}}\right] . \tag{3.49}
\end{align*}
$$

For the coherent process $\bar{\nu} \mathcal{A} \rightarrow \mu^{+} \pi^{-} \mathcal{B}$, the only change is the leptonic tensor:

$$
\begin{align*}
\mathcal{L}^{\alpha \beta} & =\mathcal{L}^{\alpha} \mathcal{L}^{\beta *} \\
& =8\left\{p_{\nu}^{\alpha} p_{\mu}^{\beta}+p_{\nu}^{\beta} p_{\mu}^{\alpha}-p_{\nu} \cdot p_{\mu} g^{\alpha \beta}-i \epsilon^{\alpha \beta \gamma \delta} p_{\nu \gamma} p_{\mu_{\delta}}\right\} \tag{3.50}
\end{align*}
$$

However, under the assumption of the small $Q^{2}$ and leptonic mass $\approx 0$, then the leptonic tensor $=0$. Then we come to the conclusion that the cross-section of the process $\bar{\nu} \mathcal{A} \rightarrow \mu^{+} \pi^{-} \mathcal{B}$

$$
\begin{equation*}
\left.d \sigma\left(\bar{\nu}+A->\mu^{+}+\pi^{-}+B\right)\right|_{Q^{2} \rightarrow 0}=\left.d \sigma\left(\nu+A->\mu^{-}+\pi^{+}+B\right)\right|_{Q^{2} \rightarrow 0} \tag{3.51}
\end{equation*}
$$

For the neutral current (NC) interaction, we have

$$
\begin{gather*}
f_{\pi^{0}}=\frac{f_{\pi^{ \pm}}}{\sqrt{2}} .  \tag{3.52}\\
d \sigma\left(\nu+\mathcal{A} \rightarrow \nu^{\prime}+\mathcal{A}^{\prime}+\pi^{0}\right)= \\
\left.G_{F}^{2} \frac{E E^{\prime}}{4 E M \nu^{2}} 2 f_{\pi}^{2} \right\rvert\, \mathcal{M}\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime} \mid\right. \\
 \tag{3.53}\\
\times(2 \pi)^{4} \delta^{4}\left(p+k-p^{\prime}-k^{\prime}\right) \\
\frac{1}{(2 \pi)^{3}} \frac{d^{3} p^{\prime}}{2 E^{\prime}} \frac{1}{(2 \pi)^{3}} \frac{d^{3} k^{\prime}}{2 k_{0}^{\prime}}  \tag{3.54}\\
d \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right)=\begin{array}{l}
\left.\frac{\mid \mathcal{M}(\mathcal{A}+}{4 M_{N} \nu} \rightarrow \mathcal{A}^{\prime}\right)\left.\right|^{2} \\
\end{array} \quad \times(2 \pi)^{4} \delta^{4}\left(p_{\nu}+p_{Z i}-p_{\mu}-p_{Z f}\right) \frac{1}{(2 \pi)^{3}} \frac{d^{3} p_{Z f}}{2 E_{Z f}} .  \tag{3.55}\\
d \sigma\left(\nu+\mathcal{A} \rightarrow \nu^{\prime}+\right. \\
\left.\mathcal{A}^{\prime}+\pi^{0}\right)=\frac{G_{F}^{2} f_{\pi}^{2}}{4 \pi^{3} \nu} \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right) d^{3} p^{\prime} .
\end{gather*}
$$

where $d^{3} p^{\prime}=\pi \frac{E^{\prime}}{E} d Q^{2} d \nu$. Making a variable change with

$$
\begin{align*}
x & =\frac{Q^{2}}{2 M \nu} \\
y & =\frac{\nu}{E} \tag{3.56}
\end{align*}
$$

we have

$$
\begin{align*}
d Q^{2} d \nu & =\left|\begin{array}{cc}
\frac{\partial Q^{2}}{\partial x} & \frac{\partial Q^{2}}{\partial y} \\
\frac{\partial \nu}{\partial x} & \frac{\partial \nu}{\partial y}
\end{array}\right| \times d x d y \\
& =2 M E \nu d x d y \\
& =2 M E^{2} y d x d y \tag{3.57}
\end{align*}
$$

and

$$
\begin{align*}
d^{3} p^{\prime} & =\pi \frac{E^{\prime}}{E} 2 M E^{2} y d x d y \\
& =2 M \pi E^{\prime} E y d x d y \tag{3.58}
\end{align*}
$$

SO

$$
\begin{align*}
d \sigma\left(\nu+\mathcal{A} \rightarrow \nu^{\prime}+\mathcal{A}^{\prime}+\pi^{0}\right)= & \frac{G_{F}^{2} f_{\pi}^{2}}{4 \pi^{3} \nu} d \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right)  \tag{3.59}\\
& 2 M \pi E^{\prime} E y d x d y \\
= & \frac{G_{F}^{2} f_{\pi}^{2} M E}{2 \pi^{2}} \frac{E^{\prime} y}{\nu} d \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right) d x d y \\
= & \frac{G_{F}^{2} M E}{\pi^{2}} \frac{1}{2} f_{\pi}^{2}(1-y)\left[d \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right)\right. \\
& \times d x d y] . \tag{3.60}
\end{align*}
$$

In the last equality, we used the formula that $y=\frac{\nu}{E}$, Then, we have:

$$
\begin{align*}
{\left[\frac{d \sigma\left(\nu+\mathcal{A} \rightarrow \nu^{\prime}+\mathcal{A}^{\prime}+\pi^{0}\right)}{d x d y d t}\right]_{Q^{2}=0}=} & \frac{G_{F}^{2} M E}{\pi^{2}} \frac{1}{2} f_{\pi}^{2}(1-y) \\
& \times\left[\frac{d \sigma\left(\mathcal{A}+\pi \rightarrow \mathcal{A}^{\prime}\right)}{d t}\right] . \tag{3.61}
\end{align*}
$$

which is consistent with the result given by [37].

### 3.3 Neutrino Induced Coherent $\rho$

This section deals with high $\nu$ (energy transfer) and low $Q^{2}$ (negative of the square of the 4 -momentum transfer) production of $\rho$ mesons in neutrino interactions. In the regime of low $Q^{2}$ the distances probed are larger than in DIS (deep inelastic scattering). The target nucleon constituents, therefore, cannot be considered as free particles and the perturbative theoretical approach towards strong interactions cannot be applied. Here the theory of VMD (Vector Meson Dominance) applied to weak interactions is used to produce a cross-section for $\rho$ production in the low $Q^{2}$ regime. The derivation of this section was created in collaboration with C.T.Kullenberg. We
will link cross-section of the production of coherent $\rho$ mesons by neutrinos $(\nu+$ $\mathcal{A} \rightarrow \mu+\rho+\mathcal{A})$ to the transverse and longitudinal cross-sections of coherent meson scattering $(\rho+\mathcal{A} \rightarrow \rho+\mathcal{A})$.

During this calculation we will assume that the target is at rest in the lab frame, and we will assume that the lepton masses can be ignored.

The square of the scattering amplitude is written as:

$$
\begin{equation*}
|\mathcal{M}|^{2}=\frac{G_{F}^{2}}{2} \mathcal{L}_{\mu \nu} W^{\mu \nu} \tag{3.62}
\end{equation*}
$$

where $W^{\mu \nu}$ is the hadronic tensor. Experiments are generally blind to particle polarization, so one must average over the initial particle spins and sum over all final particle spin states and momenta. This must be done for both the hadronic and leptonic tensors. One might re-write the square of the amplitude as:

$$
\begin{equation*}
<|\mathcal{M}|^{2}>=\frac{G_{F}^{2}}{2}<\mathcal{L}_{\mu \nu}><W^{\mu \nu}> \tag{3.63}
\end{equation*}
$$

where the brackets indicate the initial and final state averaging and summing process. Often, however, the distinction is implied and one continues to use the form in Equation (3.62).

## The Hadronic Tensor

Using the hadron dominance assumption we will consider only the contribution of the $\rho$ meson to the hadronic vector current. The calculation of the hadronic tensor will proceed in the normal fashion, but we will ignore the axial vector contribution and we will impose $q_{\mu} W^{\mu \nu}=0$ due to CVC (Conservation of Vector Current).

The most general symmetric form of the hadronic tensor is [20]

$$
\begin{equation*}
W^{\mu \nu}=W_{1} g^{\mu \nu}+W_{2} \frac{p^{\mu} p^{\nu}}{M^{2}}+W_{4} \frac{q^{\mu} q^{\nu}}{M^{2}}+W_{5} \frac{p^{\mu} q^{\nu}+q^{\mu} p^{\nu}}{M^{2}} \tag{3.64}
\end{equation*}
$$

where the coefficients $W_{i}$ are functions of $\nu$ and $q^{2}$. We keep only symmetric terms as asymmetric terms arise from the interference of the vector and axial vector currents.

So, from CVC we have:

$$
\begin{equation*}
q_{\mu} W^{\mu \nu}=W_{1} q^{\nu}+W_{2} \frac{q \cdot p p^{\nu}}{M^{2}}+W_{4} \frac{q^{2} q^{\nu}}{M^{2}}+W_{5} \frac{q \cdot p q^{\nu}+q^{2} p^{\nu}}{M^{2}}=0 \tag{3.65}
\end{equation*}
$$

which implies that the factors $W_{4}$ and $W_{5}$ can be expressed in terms of $W_{1}$ and $W_{2}$. Looking at functions of $p^{\nu}$ and $q^{\nu}$ separately:

$$
\begin{align*}
W_{5} & =W_{2} \frac{q \cdot p}{Q^{2}} \\
W_{4} & =W_{2} \frac{(q \cdot p)^{2}}{Q^{4}}+W_{1} \frac{M^{2}}{Q^{2}} \tag{3.66}
\end{align*}
$$

The hadronic tensor can be written as:

$$
\begin{equation*}
W^{\mu \nu}=W_{1}\left(g^{\mu \nu}+\frac{q^{\mu} q^{\nu}}{Q^{2}}\right)+\frac{W_{2}}{M^{2}}\left(p^{\mu}+\frac{p \cdot q}{Q^{2}} q^{\mu}\right)\left(p^{\nu}+\frac{p \cdot q}{Q^{2}} q^{\nu}\right) \tag{3.67}
\end{equation*}
$$

The $\rho$ dominance hypothesis is implemented by expressing the hadronic current as [42]

$$
\begin{equation*}
J^{\mu}=\frac{f_{\rho}}{Q^{2}+m_{\rho}^{2}}\left[g^{\mu \nu}-\frac{q_{\mu} q^{\nu}}{m_{\rho}^{2}}\right] \mathcal{A}_{\nu}(\rho+\alpha \rightarrow \beta) \tag{3.68}
\end{equation*}
$$

where $f_{\rho}$ is the coupling constant of the $\rho$ meson to the $W$ boson and $\epsilon_{\mu}^{(i)} A^{\mu}$ represents the amplitude of a $\rho$ meson with polarization $\epsilon^{(i)}$. We can write the hadronic tensor as a product of the $J^{\mu}$ currents.

$$
\begin{align*}
W^{\mu \nu}= & J^{\mu} J^{\nu *} \\
= & \left(\frac{f_{\rho}}{Q^{2}+m_{\rho}^{2}}\right)^{2}\left(\left[g^{\mu \sigma}-\frac{q_{\mu} q^{\sigma}}{m_{\rho}^{2}}\right] \mathcal{A}_{\sigma}(\rho+\alpha \rightarrow \beta)\right) \\
& \times\left(\left[g^{\nu \omega}-\frac{q_{\nu} q^{\omega}}{m_{\rho}^{2}}\right] \mathcal{A}_{\omega}(\rho+\alpha \rightarrow \beta)\right)^{*} . \tag{3.69}
\end{align*}
$$

By conservation of isospin current we have $q^{\mu} \mathcal{A}_{\mu}=0$, so the $q$ factors disappear giving us simply:

$$
\begin{equation*}
W^{\mu \nu}=C_{\rho}^{2} M^{\mu \nu} \tag{3.70}
\end{equation*}
$$

where we have defined:

$$
\begin{align*}
M^{\mu \nu} & =\mathcal{A}^{\mu} \mathcal{A}^{\nu *} \\
C_{\rho} & =\left(\frac{f_{\rho}}{Q^{2}+m_{\rho}^{2}}\right) \tag{3.71}
\end{align*}
$$

$M^{\mu \nu}$ must undergo the usual averaging and summing of states.

## Linking $W_{1}$ and $W_{2}$ to $\sigma_{T}$ and $\sigma_{L}$

We have written $W^{\mu \nu}$ in terms of the unknown coefficients $W_{1}$ and $W_{2}$, and we have also shown $W^{\mu \nu}$ to be proportional to $M^{\mu \nu}$. We would like to express $W^{\mu \nu}$ in terms of the transverse $\left(\sigma_{T}\right)$ and longitudinal $\left(\sigma_{L}\right)$ polarized cross-sections of the incident $\rho$ meson. We can accomplish this if we use the following definitions from [20] and [50]:

$$
\begin{align*}
\sigma_{T} & =\sigma_{T}(\rho+\alpha \rightarrow \beta)=\frac{1}{|\vec{q}|} \epsilon_{\mu}^{* T} M^{\mu \nu} \epsilon_{\nu}^{T} \\
\sigma_{L} & =\sigma_{L}(\rho+\alpha \rightarrow \beta)=\frac{1}{|\vec{q}|} \epsilon_{\mu}^{* L} M^{\mu \nu} \epsilon_{\nu}^{L} \tag{3.72}
\end{align*}
$$

It should be noted here that these definitions do not include integration over final momenta, which must be done eventually. We will also need the explicit form of the polarization vectors given here [20].

$$
\epsilon^{T}=\mp \frac{1}{\sqrt{2}}\left(\begin{array}{c}
0  \tag{3.73}\\
1 \\
\pm i \\
0
\end{array}\right) \quad \epsilon^{L}=\frac{1}{\sqrt{Q^{2}}}\left(\begin{array}{c}
|\vec{q}| \\
0 \\
0 \\
\nu
\end{array}\right)
$$

Massive vector bosons have three possible polarization vectors (two transverse and one longitudinal). In our calculations we will make use of $\epsilon_{\mu}^{*} \epsilon^{\mu}$, which has the same value for both transverse polarizations. It, therefore, doesn't matter which one we choose when we use $\epsilon^{T}$.

Now from Equation (3.70) we see that $M^{\mu \nu}=\frac{1}{C_{\rho}^{2}} W^{\mu \nu}$, and looking at Equation (3.67) we can write it very simply:

$$
\begin{equation*}
M^{\mu \nu}=\frac{1}{C_{\rho}^{2}}\left[W_{1} \mathcal{F}_{1}+W_{2} \mathcal{F}_{2}\right] \tag{3.74}
\end{equation*}
$$

where, for economy of space, we have made functions $\mathcal{F}_{1}$ and $\mathcal{F}_{2}$ :

$$
\begin{align*}
& \mathcal{F}_{1}=\left(g^{\mu \nu}+\frac{q^{\mu} q^{\nu}}{Q^{2}}\right) \\
& \mathcal{F}_{2}=\frac{1}{M^{2}}\left(p^{\mu}+\frac{p \cdot q}{Q^{2}} q^{\mu}\right)\left(p^{\nu}+\frac{p \cdot q}{Q^{2}} q^{\nu}\right) . \tag{3.75}
\end{align*}
$$

Now putting our new form of $M^{\mu \nu}$ into $\sigma_{T}$ and $\sigma_{L}$ :

$$
\begin{align*}
\sigma_{T} & =\frac{1}{C_{\rho}^{2}|\vec{q}|} \epsilon_{\mu}^{* T}\left(W_{1} \mathcal{F}_{1}+W_{2} \mathcal{F}_{2}\right) \epsilon_{\nu}^{T} \\
\sigma_{L} & =\frac{1}{C_{\rho}^{2}|\vec{q}|} \epsilon_{\mu}^{* L}\left(W_{1} \mathcal{F}_{1}+W_{2} \mathcal{F}_{2}\right) \epsilon_{\nu}^{L} \tag{3.76}
\end{align*}
$$

Or we might make a few simple definitions to clean it up a bit:

$$
\begin{align*}
\sigma_{T} & =\frac{1}{C_{\rho}^{2}|\vec{q}|}\left(W_{1} A_{T}+W_{2} B_{T}\right) \\
\sigma_{L} & =\frac{1}{C_{\rho}^{2}|\vec{q}|}\left(W_{1} A_{L}+W_{2} B_{L}\right) \tag{3.77}
\end{align*}
$$

At this point one need only calculate $A_{T}, B_{T}, A_{L}$ and $B_{L}$, which are just contractions of the polarization vectors with the functions $\mathcal{F}_{1}$ and $\mathcal{F}_{2}$, in order to express $W_{1}$ and $W_{2}$ in terms of $\sigma_{T}$ and $\sigma_{L}$.

By noting that $q^{\nu} \epsilon_{\nu}=0$ (due to gauge invariance), $p=(M, 0,0,0)$ because the target is at rest and using the explicit form of the polarization vectors in Equation (3.73) it is easily shown that:

$$
\begin{align*}
& A_{T}=\epsilon_{\mu}^{* T} \mathcal{F}_{1} \epsilon_{\nu}^{T}=-1 \\
& B_{T}=\epsilon_{\mu}^{* T} \mathcal{F}_{2} \epsilon_{\nu}^{T}=0 \\
& A_{L}=\epsilon_{\mu}^{* L} \mathcal{F}_{1} \epsilon_{\nu}^{L}=1 \\
& B_{L}=\epsilon_{\mu}^{* L} \mathcal{F}_{2} \epsilon_{\nu}^{L}=\frac{|\vec{q}|^{2}}{Q^{2}} . \tag{3.78}
\end{align*}
$$

Putting these back into Equation (3.77) and solving for $W_{1}$ and $W_{2}$ gives us the relationships that we've been looking for.

$$
\begin{align*}
W_{1} & =-C_{\rho}^{2}|\vec{q}| \sigma_{T} \\
W_{2} & =\frac{C_{\rho}^{2} Q^{2}}{|\vec{q}|}\left(\sigma_{T}+\sigma_{L}\right) \tag{3.79}
\end{align*}
$$

We can put this form of $W_{1}$ and $W_{2}$ into our expression for the hadronic tensor $W^{\mu \nu}$
in Equation Equation (3.67).

$$
\begin{align*}
W^{\mu \nu}= & \left(\frac{f_{\rho}}{Q^{2}+m_{\rho}^{2}}\right)^{2}\left[-|\vec{q}| \sigma_{T}\left(g^{\mu \nu}+\frac{q^{\mu} q^{\nu}}{Q^{2}}\right)\right. \\
& \left.\quad+\frac{Q^{2}\left(\sigma_{T}+\sigma_{L}\right)}{|\vec{q}| M^{2}}\left(p^{\mu}+\frac{p \cdot q}{Q^{2}} q^{\mu}\right)\left(p^{\nu}+\frac{p \cdot q}{Q^{2}} q^{\nu}\right)\right] \tag{3.80}
\end{align*}
$$

## Calculating $|\mathcal{M}|^{2}$

Now that we have the Hadronic and Leptonic tensors in a useable form we can put them into the equation for the square of the scattering amplitude Equation (3.62).

To simplify a bit we will rewrite Equation (3.80) using the definitions we made in Equation (3.71) and Equation (3.75).

$$
\begin{equation*}
W^{\mu \nu}=|\vec{q}| f_{\rho}^{2}\left[-\sigma_{T} \mathcal{F}_{1}+\frac{Q^{2}}{|\vec{q}|^{2}}\left(\sigma_{T}+\sigma_{L}\right) \mathcal{F}_{2}\right] . \tag{3.81}
\end{equation*}
$$

Putting this into $|\mathcal{M}|^{2}$ :

$$
\begin{align*}
|\mathcal{M}|^{2}= & 4|\vec{q}| G_{F}^{2} C_{\rho}^{2}\left[-\sigma_{T}\left(T_{1}+T_{2}-T_{3}\right)\right. \\
& \left.+\frac{Q^{2}}{|\vec{q}|^{2}}\left(\sigma_{T}+\sigma_{L}\right)\left(T_{4}+T_{5}-T_{6}\right)\right] \tag{3.82}
\end{align*}
$$

Were we need to calculate the six $T_{i}$ terms:

$$
\begin{align*}
& T_{1}=\left(k_{\mu}^{\prime} k_{\nu} \mathcal{F}_{1}\right) \quad T_{2}=\left(k_{\mu} k_{\nu}^{\prime} \mathcal{F}_{1}\right) \quad T_{3}=\left(k \cdot k^{\prime} g_{\mu \nu} \mathcal{F}_{1}\right) \\
& T_{4}=\left(k_{\mu}^{\prime} k_{\nu} \mathcal{F}_{2}\right) \quad T_{5}=\left(k_{\mu} k_{\nu}^{\prime} \mathcal{F}_{2}\right) \quad T_{6}=\left(k \cdot k^{\prime} g_{\mu \nu} \mathcal{F}_{2}\right) \tag{3.83}
\end{align*}
$$

To do this we will use:

$$
\begin{equation*}
k \cdot k^{\prime} \approx \frac{Q^{2}}{2} \quad\left(m_{l} \rightarrow 0\right) \tag{3.84}
\end{equation*}
$$

which will give us the following for the first three terms:

$$
\begin{equation*}
T_{1}+T_{2}-T_{3}=\left(2 \frac{(k \cdot q)\left(k^{\prime} \cdot q\right)}{Q^{2}}-k \cdot k^{\prime}\right) \approx-Q^{2} . \tag{3.85}
\end{equation*}
$$

For the second set of three terms we will simplify with the following relations that are valid when the target is at rest:

$$
\begin{array}{cl}
p^{2}=M^{2} & p \cdot q=M \nu \\
p \cdot k=M E & p \cdot k^{\prime}=M E^{\prime} . \tag{3.86}
\end{array}
$$

This will give us:

$$
\begin{equation*}
\left(T_{4}+T_{5}-T_{6}\right)=\frac{1}{2}\left(4 E E^{\prime}-Q^{2}\right) . \tag{3.87}
\end{equation*}
$$

Putting Equation (3.85) and Equation (3.87) into Equation (3.82) we can write the square of the scattering amplitude.

$$
\begin{equation*}
|\mathcal{M}|^{2}=4 G_{F}^{2}|\vec{q}| f_{\rho}^{2} \frac{Q^{2}}{\left(Q^{2}+m_{\rho}^{2}\right)^{2}}\left[\sigma_{T}+\frac{\left(\sigma_{T}+\sigma_{L}\right)}{2|\vec{q}|^{2}}\left(4 E E^{\prime}-Q^{2}\right)\right] . \tag{3.88}
\end{equation*}
$$

## Calculating the Cross-Section

The differential cross-section is given by [41].

$$
\begin{equation*}
d \sigma=\frac{|\mathcal{M}|^{2}}{2 E_{A} 2 E_{B} \mathcal{U}_{A B}}\left(\prod_{f} \frac{d^{3} p_{f}}{(2 \pi)^{3}} \frac{1}{2 E_{f}}\right)(2 \pi)^{4} \delta^{4}\left(p_{A}+p_{B}-\Sigma p_{f}\right) \tag{3.89}
\end{equation*}
$$

where $A$ and $B$ are the initial particles, $p_{f}{ }^{\mathrm{s}}$ are final state particle four-momenta, and $\mathcal{U}_{A B}$ is the relative velocity between $A$ and $B$ in the laboratory frame, which in this case is unity because the neutrino effectively moves at the speed of light and the target is at rest.

Here we are considering two final state "particles", the outgoing lepton and the final hadronic state. So the differential cross-section can be written as:

$$
\begin{equation*}
d \sigma=\frac{|\mathcal{M}|^{2}}{2 E 2 M}\left(\frac{d^{3} k^{\prime}}{(2 \pi)^{3}} \frac{1}{2 E^{\prime}}\right)\left(\frac{d^{3} p^{\prime}}{(2 \pi)^{3}} \frac{1}{2 p_{0}^{\prime}}\right)(2 \pi)^{4} \delta^{4}\left(k+p-k^{\prime}-p^{\prime}\right) \tag{3.90}
\end{equation*}
$$

## Integrating Over $p^{\prime}$ for $\sigma_{T}$ and $\sigma_{L}$

At this point we need to expose some slightly ambiguous notation that we have been using for the $\rho$ transverse and longitudinal cross-sections. We will deal here with $\sigma_{T}$, but these arguments apply equally well to $\sigma_{L}$.

We have so far defined the transverse cross-section as:

$$
\sigma_{T}^{0}=\frac{1}{|\vec{q}|} \epsilon_{\mu}^{* T} M^{\mu \nu} \epsilon_{\nu}^{T}=\frac{1}{|\vec{q}|}\left|\epsilon_{\mu}^{T} A^{\mu}\right|^{2}
$$

We can think of this as the amplitude squared for the process $(\rho+\alpha \rightarrow \beta)$, normalized by the inverse of the $\rho$ momentum. For now we have renamed the original definition as $\sigma_{T}^{0}$. To transform it into an actual cross-section we must integrate over the final momenta.

We can use Equation (3.89) with the amplitude we have and note that the velocity of the $\rho$ in the lab frame is $\frac{|\vec{q}|}{\nu}$ (in general $p=\gamma m v=E v \Rightarrow v=\frac{p}{E}$ ), and because the target is at rest we have:

$$
\begin{equation*}
\mathcal{U}_{A B}=\frac{|\vec{q}|}{\nu} . \tag{3.91}
\end{equation*}
$$

So we can write the true cross-section $\sigma_{T}$ as:

$$
\begin{align*}
\sigma_{T} & =\int \frac{\left|\epsilon_{\mu}^{T} A^{\mu}\right|^{2}}{2 \nu 2 M\left(\frac{\mid \vec{q}}{\nu}\right)}\left(\frac{d^{3} p^{\prime}}{(2 \pi)^{3}} \frac{1}{2 p_{0}^{\prime}}\right)(2 \pi)^{4} \delta^{4}\left(q+p-p^{\prime}\right) \\
& =\int \frac{\left|\epsilon_{\mu}^{T} A^{\mu}\right|^{2}}{2|\vec{q}| 2 M}\left(\frac{d^{3} p^{\prime}}{2 p_{0}^{\prime}}\right)(2 \pi) \delta\left(\nu+M-p_{0}^{\prime}\right) \delta^{3}\left(\vec{q}+\vec{p}-\vec{p}^{\prime}\right) \\
& =\left(\frac{\left|\epsilon_{\mu}^{T} A^{\mu}\right|^{2}}{|\vec{q}|}\right) \frac{\pi}{2 p_{0}^{\prime} 2 M} \delta\left(\nu+M-p_{0}^{\prime}\right) \\
& =\left(\frac{\pi}{4 M p_{0}^{\prime}} \delta\left(\nu+M-p_{0}^{\prime}\right)\right) \sigma_{T}^{0} . \tag{3.92}
\end{align*}
$$

We can see that if we multiply the original definition of $\sigma_{T}^{0}$ by the factor $\left(\frac{\pi}{4 M p_{0}^{\prime}} \delta(\nu+\right.$ $\left.\left.m-p_{0}^{\prime}\right)\right)$ then we obtain the full cross-section $\sigma_{T}$, which has been integrated over the final momenta. This is identically true for the longitudinal $\rho$ cross-section $\sigma_{L}$. And because $|\mathcal{M}|^{2}$ is linearly proportional to $\sigma_{T}^{0}$ and $\sigma_{L}^{0}$ we might modify $|\mathcal{M}|^{2}$ itself by this same factor, absorbing it and simply stating that we have performed the momentum integration for $\sigma_{T}$ and $\sigma_{L}$. So we can write:

$$
\begin{equation*}
|\mathcal{M}|^{2}\left(\frac{\pi}{4 M p_{0}^{\prime}} \delta\left(\nu+M-p_{0}^{\prime}\right)\right) \xrightarrow[\sigma_{L} \text { over } p^{\prime}]{\text { Integrate } \sigma_{T} \text { and }}|\mathcal{M}|^{2} . \tag{3.93}
\end{equation*}
$$

If we are presented with this factor then we will absorb it into $|\mathcal{M}|^{2}$ and make reference to this rule.

## Integrating Over $p^{\prime}$ for $d \sigma$

We will now return, after our digression, to Equation Equation (3.90). We must integrate over the final momenta.

$$
\begin{align*}
d \sigma & =\int_{p^{\prime}} \frac{|\mathcal{M}|^{2}}{2 E 2 M}\left(\frac{d^{3} k^{\prime}}{(2 \pi)^{3}} \frac{1}{2 E^{\prime}}\right)\left(\frac{d^{3} p^{\prime}}{(2 \pi)^{3}} \frac{1}{2 p_{0}^{\prime}}\right)(2 \pi)^{4} \delta^{4}\left(k+p-k^{\prime}-p^{\prime}\right) \\
& =\frac{|\mathcal{M}|^{2}}{2 E 2 M}\left(\frac{d^{3} k^{\prime}}{(2 \pi)^{3}} \frac{1}{2 E^{\prime}}\right)\left(\frac{1}{(2 \pi)^{3}} \frac{1}{2 p_{0}^{\prime}}\right)(2 \pi)^{4} \delta\left(E+M-E^{\prime}-p_{0}^{\prime}\right) \\
& =|\mathcal{M}|^{2}\left(\frac{\pi}{4 M p_{0}^{\prime}} \delta\left(\nu+M-p_{0}^{\prime}\right)\right)\left(\frac{d^{3} k^{\prime}}{(2 \pi)^{3}} \frac{1}{2 E E^{\prime}}\right) . \tag{3.94}
\end{align*}
$$

Here we will invoke our rule Equation (3.93) and absorb the middle term into $|\mathcal{M}|^{2}$ giving us:

$$
\begin{equation*}
d \sigma=|\mathcal{M}|^{2}\left(\frac{d^{3} k^{\prime}}{(2 \pi)^{3}} \frac{1}{2 E E^{\prime}}\right) \tag{3.95}
\end{equation*}
$$

## $d \sigma$ in Terms of $d Q^{2}$ and $d \nu$

We would now like to express $d \sigma$ in terms of $d Q^{2}$ and $d \nu$ rather than $d^{3} k^{\prime}$. Let's first look at $d Q^{2}$. When ignoring the lepton masses we have from Equation (3.3) $Q^{2} \approx 2 E E^{\prime}(1-\cos \theta)$. For it's derivative we have:

$$
\begin{equation*}
d Q^{2}=\left(\frac{\partial Q^{2}}{\partial \theta}\right) d \theta+\left(\frac{\partial Q^{2}}{\partial E^{\prime}}\right) d E^{\prime}+\left(\frac{\partial Q^{2}}{\partial E}\right) d E \tag{3.96}
\end{equation*}
$$

We will not integrate over the initial momentum (or energy for $m=0$ ), therefore $d E \rightarrow 0$ leaving us with:

$$
\begin{equation*}
d Q^{2}=\left(\frac{\partial Q^{2}}{\partial \theta}\right) d \theta+\left(\frac{\partial Q^{2}}{\partial E^{\prime}}\right) d E^{\prime} \tag{3.97}
\end{equation*}
$$

Now for $d \nu$ we simply have $d \nu=d\left(E-E^{\prime}\right)=d E-d E^{\prime} \rightarrow d \nu=-d E^{\prime}$. The minus sign seems troublesome, but it is a simple matter of changing the integration limits.

$$
\begin{equation*}
\int_{0}^{E} d E^{\prime}=-\int_{E}^{0} d \nu=\int_{0}^{E} d \nu \tag{3.98}
\end{equation*}
$$

When the outgoing lepton has zero energy, the full energy E has been transferred to the hadronic system. We will absorb the minus sign into $d \nu$ and integrate from
zero momentum transfer to full energy transfer (which one might normally expect). So $d E^{\prime} \rightarrow d \nu$, and if we multiply $d Q^{2}$ by this we have:

$$
\begin{equation*}
d Q^{2} d \nu=\left(\frac{\partial Q^{2}}{\partial \theta}\right) d \theta d E^{\prime}+\left(\frac{\partial Q^{2}}{\partial E^{\prime}}\right)\left(d E^{\prime}\right)^{2} \tag{3.99}
\end{equation*}
$$

We will not integrate twice over the final momenta, so $\left(d E^{\prime}\right)^{2} \rightarrow 0$ and we finally have:

$$
\begin{align*}
d Q^{2} d \nu & =\left(\frac{\partial Q^{2}}{\partial \theta}\right) d \theta d E^{\prime} \\
& =\left(2 E E^{\prime} \sin \theta\right) d \theta d E^{\prime} \tag{3.100}
\end{align*}
$$

For $d^{3} k^{\prime}$ we have:

$$
\begin{align*}
d^{3} k^{\prime} & =\left|\vec{k}^{\prime}\right|^{2} d\left|\overrightarrow{k^{\prime}}\right| d \Omega \approx E^{\prime 2} d E^{\prime} d \Omega \\
& =2 \pi E^{\prime 2} d E^{\prime} \sin \theta d \theta \tag{3.101}
\end{align*}
$$

Putting the above two equations together we can relate $d^{3} k^{\prime}$ to $d Q^{2} d \nu$.

$$
\begin{equation*}
d^{3} k^{\prime}=\pi \frac{E^{\prime}}{E} d Q^{2} d \nu \tag{3.102}
\end{equation*}
$$

After inserting this into Equation (3.95) we have a new form for our cross-section.

$$
\begin{equation*}
d \sigma=|\mathcal{M}|^{2}\left(\frac{d Q^{2} d \nu}{16 \pi^{2} E^{2}}\right) \tag{3.103}
\end{equation*}
$$

## Final Cross-Section

Simply inserting our $|\mathcal{M}|^{2}$ from Equation (3.88) into Equation (3.103) gives us the following:

$$
\begin{align*}
\frac{d^{2} \sigma(\nu \alpha \rightarrow \mu \beta)}{d Q^{2} d \nu}= & \frac{G_{F}^{2}|\vec{q}|}{4 \pi^{2} E^{2}} f_{\rho}^{2} \frac{Q^{2}}{\left(Q^{2}+m_{\rho}^{2}\right)^{2}} \\
& \times\left[\sigma_{T}+\frac{\left(\sigma_{T}+\sigma_{L}\right)}{2|\vec{q}|^{2}}\left(4 E E^{\prime}-Q^{2}\right)\right] \tag{3.104}
\end{align*}
$$

We are basically done here, but one might also make a couple of definitions to write this in another way.

$$
\begin{equation*}
\epsilon=\frac{4 E E^{\prime}-Q^{2}}{4 E E^{\prime}+Q^{2}+2 \nu^{2}} \tag{3.105}
\end{equation*}
$$

$$
\begin{equation*}
R=\frac{d \sigma_{L} / d t}{d \sigma_{T} / d t} \tag{3.106}
\end{equation*}
$$

Then we can modify the term in brackets and write the cross-section in its final form.

$$
\begin{align*}
\frac{d^{3} \sigma\left(\nu_{\mu} \mathcal{A} \rightarrow \mu^{-} \rho^{+} \mathcal{A}\right)}{d Q^{2} d \nu d t}= & \frac{G_{F}^{2}}{4 \pi^{2}} f_{\rho}^{2} \frac{|\vec{q}|}{E^{2}}\left(\frac{Q}{Q^{2}+m_{\rho}^{2}}\right)^{2}\left(\frac{1+\epsilon R}{1-\epsilon}\right) \\
& \times\left[\frac{d \sigma_{T}\left(\rho^{+} \mathcal{A} \rightarrow \rho^{+} \mathcal{A}\right)}{d t}\right] \tag{3.107}
\end{align*}
$$

For the anti-neutrino process $\bar{\nu} \mathcal{A} \rightarrow \mu^{+} \rho^{-} \mathcal{A}$ we have the same coupling constant between the $\rho^{-}$and the $W$ boson.

$$
\begin{equation*}
f_{\rho}^{+}=f_{\rho}^{-} . \tag{3.108}
\end{equation*}
$$

For the anti-particle process the leptonic tensor is different only in the sign of the anti-symmetric term [50].

$$
\begin{equation*}
<L^{\mu \nu}>_{\bar{\nu}}=8\left(k^{\mu} k^{\prime \nu}+k^{\nu} k^{\prime \mu}-k \cdot k^{\prime} g^{\mu \nu}-i \epsilon^{\mu \nu \sigma \lambda} k_{\lambda} k_{\sigma}^{\prime}\right) \tag{3.109}
\end{equation*}
$$

Because the leptonic tensor contracts with the hadronic tensor, and the hadronic tensor is symmetric, the anti-symmetric term must vanish. Therefore the leptonic tensor is the same, the hadronic tensor is unchanged, and we might then conclude that the cross-sections of the neutrino and anti-neutrino processes are equal.

$$
\begin{equation*}
\sigma\left(\bar{\nu} \mathcal{A} \rightarrow \mu^{+} \rho^{-} \mathcal{A}\right)=\sigma\left(\nu \mathcal{A} \rightarrow \mu^{-} \rho^{+} \mathcal{A}\right) \tag{3.110}
\end{equation*}
$$

### 3.4 Cross-section of Coherent $\rho^{0}$

In section 10.6 of Griffiths' book [34] one can find information regarding the neutral weak interactions. I will list some basic information from that resource here.

$$
\begin{align*}
-\frac{i g_{W}}{2 \sqrt{2}} \gamma^{\mu}\left(1-\gamma^{5}\right) & W^{ \pm} \text {vertex } \\
-\frac{i g_{Z}}{2} \gamma^{\mu}\left(C_{V}^{f}-C_{A}^{f} \gamma^{5}\right) & Z^{0} \text { vertex } \tag{3.111}
\end{align*}
$$

Table 3.1: Table of vector and axial vector couplings.

| f | $C_{V}$ | $C_{A}$ |
| :---: | :---: | :---: |
| $\nu_{e}, \nu_{\mu}, \nu_{\tau}$ | $\frac{1}{2}$ | $\frac{1}{2}$ |
| $e^{-}, \mu^{-}, \tau^{-}$ | $-\frac{1}{2}+2 \sin \theta_{W}$ | $-\frac{1}{2}$ |
| $\mathrm{u}, \mathrm{c}, \mathrm{t}$ | $\frac{1}{2}-\frac{4}{3} \sin ^{2} \theta_{W}$ | $\frac{1}{2}$ |
| $\mathrm{~d}, \mathrm{~s}, \mathrm{~b}$ | $-\frac{1}{2}+\frac{2}{3} \sin ^{2} \theta_{W}$ | $-\frac{1}{2}$ |

where, in the GWS model, all parameters are determined by the weak mixing angle $\theta_{W}$, which must itself be measured from experiment as we currently have no method to calculate it. A reasonable measure for the angle is $\theta_{W}=28.7^{\circ}$, or $\sin ^{2} \theta_{W}=0.23$. $g_{W}$ and $g_{Z}$ are related to the electromagnetic coupling constant $g_{e}$.

$$
\begin{equation*}
g_{W}=\frac{g_{e}}{\sin \theta_{W}} \quad g_{Z}=\frac{g_{e}}{\sin \theta_{W} \cos \theta_{W}} . \tag{3.112}
\end{equation*}
$$

The GWS vector and axial vector couplings can be gotten from Table 10.1 in [34], shown here in Table 3.1.

Additionally, the $W$ and $Z$ masses are related in a simple way.

$$
\begin{equation*}
M_{W}=M_{Z} \cos \theta_{W} \tag{3.113}
\end{equation*}
$$

For the charged current amplitude, we have:

$$
\begin{align*}
\mathcal{M}_{W}= & <\rho^{+}\left|\bar{u}\left[-i \frac{g}{\sqrt{2}} \gamma^{\mu} \frac{1}{2}\left(1-\gamma^{5}\right) V_{u d}\right] \bar{d}\right| 0>\frac{-1}{M_{W}^{2}} \bar{u}_{\mu}\left[-i \frac{g}{\sqrt{2}} \gamma_{\alpha} \frac{1}{2}\left(1-\gamma^{5}\right)\right] u_{\nu_{\mu}} \times \\
& \times A\left(\rho^{+} \mathcal{A} \rightarrow \rho^{+} \mathcal{A}\right) \\
= & V_{u d} \frac{g^{2}}{8 M_{W}^{2}}<\rho^{+}\left|\bar{u} \gamma^{\alpha} \bar{d}\right| 0>\bar{u}_{\mu} \gamma_{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A \\
= & \frac{G_{F}}{\sqrt{2}} V_{u d} M_{\rho^{+}} f_{\rho^{+}} \epsilon_{\rho^{+}}^{\alpha} \bar{u}_{\mu} \gamma_{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A . \tag{3.114}
\end{align*}
$$

While for the neutral current we have (because $\rho^{0}=\frac{1}{\sqrt{2}}(u \bar{u}-d \bar{d})$ ):

$$
\begin{align*}
\mathcal{M}_{Z}= & \frac{1}{\sqrt{2}}\left\{\frac{i g}{\cos \theta_{W}}<\rho^{0}\left|\bar{u} \gamma^{\alpha}\left[\frac{1}{4}-\frac{2}{3} \sin ^{2} \theta_{W}-\frac{1}{4} \gamma^{5}\right] u\right| 0>+\right. \\
& \left.+\frac{i g}{\cos \theta_{W}}<\rho^{0}\left|\bar{d} \gamma^{\alpha}\left[-\frac{1}{4}+\frac{1}{3} \sin ^{2} \theta_{W}+\frac{1}{4} \gamma^{5}\right] d\right| 0>\right\} \times \\
& \times \frac{1}{M_{Z}^{2}} \bar{u}_{\nu_{\mu}}\left[\frac{-i g}{\cos \theta_{W}} \gamma_{\alpha}\left(\frac{1}{4}-\frac{1}{4} \gamma^{5}\right)\right] u_{\nu_{\mu}} A\left(\rho^{0} \mathcal{A} \rightarrow \rho^{0} \mathcal{A}\right) \\
= & \frac{1}{\sqrt{2}}\left\{\frac{-i g}{\cos \theta_{W}}\left(\frac{1}{4}-\frac{2}{3} \sin ^{2} \theta_{W}\right)<\rho^{0}\left|\bar{u} \gamma^{\alpha} u\right| 0>+\right. \\
& \left.+\frac{i g}{\cos \theta_{W}}\left(-\frac{1}{4}+\frac{1}{3} \sin ^{2} \theta_{W}\right)<\rho^{0}\left|\bar{d} \gamma^{\alpha} d\right| 0>\right\} \times \\
& \times \frac{-1}{M_{Z}^{2}} \bar{u}_{\nu_{\mu}}\left[\frac{-i g}{\cos \theta_{W}} \gamma^{\alpha}\left(\frac{1}{4}-\frac{1}{4} \gamma^{5}\right)\right] u_{\nu_{\mu}} A \\
= & \frac{1}{\sqrt{2}} \frac{-i g}{\cos \theta_{W}}\left(\frac{1}{2}-\sin \theta^{2} \theta_{W}\right) M_{\rho^{0}} f_{\rho^{0}} \epsilon_{\rho^{0}}^{\alpha} \frac{-1}{M_{Z}^{2}}\left(\frac{-i g}{\cos \theta_{W}} \frac{1}{4}\right) \bar{u}_{\nu_{\mu}} \gamma^{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A \\
= & \frac{g^{2}}{8 \sqrt{2} M_{Z}^{2} \cos ^{2} \theta_{W}}\left(1-2 \sin ^{2} \theta_{W}\right) M_{\rho^{0}} f_{\rho^{0}} \epsilon_{\rho^{0}}^{\alpha} \bar{u}_{\nu_{\mu}} \gamma^{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A \\
= & \frac{g^{2}}{8 \sqrt{2} M_{W}^{2}}\left(1-2 \sin ^{2} \theta_{W}\right) M_{\rho^{0}} f_{\rho^{0}} \epsilon_{\rho_{0}}^{\alpha} \bar{u}_{\nu_{\mu}} \gamma^{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A \\
= & \frac{G_{F}}{2}\left(1-2 \sin ^{2} \theta_{W}\right) M_{\rho^{0}} f_{\rho^{0}} \epsilon_{\rho_{0}}^{\alpha} \bar{u}_{\nu_{\mu}} \gamma^{\alpha}\left(1-\gamma^{5}\right) u_{\nu_{\mu}} A, \tag{3.115}
\end{align*}
$$

where $f_{\rho^{0}} \approx f_{\rho^{+}}, V_{u d}=1$, isospin conservation gives $M_{\rho^{0}}=M_{\rho^{+}}$and we use the high energy approximation $M_{\mu}=M_{\nu}=0$.

So we can see that for coherent $\rho^{0}$ :

$$
\begin{equation*}
\mathcal{M}\left(\nu_{\mu} \mathcal{A} \rightarrow \nu_{\mu} \rho^{0} \mathcal{A}\right)=\frac{1}{\sqrt{2}}\left(1-2 \sin ^{2} \theta_{W}\right) \mathcal{M}\left(\nu_{\mu} \mathcal{A} \rightarrow \mu^{-} \rho^{+} \mathcal{A}\right) \tag{3.116}
\end{equation*}
$$

and the coherent $\rho^{0}$ cross-section can be written as

$$
\begin{equation*}
\frac{d^{3} \sigma\left(\nu_{\mu} \mathcal{A} \rightarrow \nu_{\mu} \rho^{0} \mathcal{A}\right)}{d Q^{2} d \nu d t}=\frac{1}{2}\left(1-2 \sin ^{2} \theta_{W}\right)^{2} \frac{d^{3} \sigma\left(\nu_{\mu} \mathcal{A} \rightarrow \mu^{-} \rho^{+} \mathcal{A}\right)}{d Q^{2} d \nu d t} \tag{3.117}
\end{equation*}
$$

## Chapter 4

## The Neutrino Oscillation Magnetic Detector (NOMAD) EXPERIMENT

### 4.1 The NOMAD Neutrino Beam

The NOMAD experiment is designed to search for $\nu_{\mu} \rightarrow \nu_{\tau}$ and $\nu_{\mu} \rightarrow \nu_{e}$ oscillations in a predominantly $\nu_{\mu}$ beam at CERN. Figure 4.1 shows the predicted distributions of neutrino and anti-neutrino flux. The neutrino beam in NOMAD, with a 25 GeV average energy, was produced from the in-flight decays of the secondary mesons, such as $\pi^{ \pm}, K^{ \pm}, K^{0}$. The mesons originated from the 450 GeV protons from the Super Proton Synchrotron (SPS) incident on a beryllium target (made of 11 rods 10 cm long and 2 mm in diameter each separated by 9 cm gaps). The secondary pions and kaons were focused by a pair of coaxial magnetic lenses: a horn and a reflector (shown in Figure 4.2). In this system, charged particles were deflected by the toroidal field between two coaxial conductors carrying equal and opposite currents so that the focusing of particles of one sign implied defocusing particles of the opposite sign. Collimators reduced the anti-neutrinos contamination by intercepting the defocused secondaries. The mesons were allowed to decay within a 290 m long evacuated decay pipe. Shielding made from iron and earth followed which was used to range out muons and absorb hadrons. The average neutrino flight path to the NOMAD detector was 628 m , the detector being 836 m downstream of the Be-target.

The Monte Carlo simulation predicted the relative abundance of neutrino species: $\nu_{\mu}: \bar{\nu}_{\mu}: \nu_{e}: \bar{\nu}_{e}=1.00: 0.061: 0.0094: 0.0024$. with average energies of 23.5, 19.2,
37.1 , and 31.3 GeV [48].


Figure 4.1: Predicted neutrino $\left(\nu_{\mu}\right)$ and anti-neutrino $\left(\bar{\nu}_{\mu}\right)$ flux in NOMAD.

### 4.2 The NOMAD Detector



Figure 4.2: Schematic layout of the West Area Neutrino Facility(WANF) beam line [8].

The NOMAD detector was designed to measure and identify most of the particles, including charged and neutral particles, produced in the interaction of neutrino and target [5], which was composed of several sub-detectors. A top view of the NOMAD
detector is shown in Figure 4.3. In this Figure, the fiducial volume of the NOMAD detector consists of 44 drift chambers and with a low average density (around $\left.0.1 \mathrm{~g} \mathrm{~cm}^{-3}\right)$. The drift chambers are located within a dipole magnet providing a 0.4 T magnetic field orthogonal to the neutrino beam line. The existence of magnetic field allows for the determination of the momenta of charged tracks via their curvature with minimal degradation due to multiple scattering. The direction of the magnetic field is chosen as the X reference axis. The incoming neutrinos' direction is called Z axis. Y axis is orthogonal to both of them. On average, the equivalent material in the DC encountered by particles produced in a $\nu$-interaction was about half of a radiation length and a quarter of a hadronic interaction length $(\lambda)$. The fiducial mass of the NOMAD DC-target, 2.7 tons, was composed primarily of carbon ( $64 \%$ ), oxygen $(22 \%)$, nitrogen ( $6 \%$ ), and hydrogen ( $5 \%$ ) yielding an effective atomic number, $\mathrm{A}=12.8$, similar to carbon. There were nine modules of transition radiation detectors (TRD) located downstream of the target which are used to separate electrons from pions. Since Pions are about 270 times heavier than electrons, the chance of emitting a transition radiation X-ray is much smaller than electrons. With this TRD, about $99.9 \%$ of pions are rejected and $90 \%$ electrons are kept. The TRD was followed by an electromagnetic calorimeter (ECAL) including a preshower, a hadronic calorimeter and a muon chamber providing a clean identification of the muons. The energy resolution of this ECAL in NOMAD is $\frac{\sigma(E)}{E}=(1.04 \pm 0.01) \%+\frac{3.22 \pm 0.07}{\%} E(\mathrm{GeV})$. With the muon chamber, about $92 \%$ of the muons with momenta above 6 GeV can be reconstructed. In the NOMAD detector, the charged tracks could be reconstructed in the DC with an approximate momentum $(p)$ resolution of $\delta p / p 3.5 \%$. The experiment recorded over 1.7 million neutrino interactions in the active drift-chamber (DC) target in the range $O(1) \leq E_{\nu} \leq 300 \mathrm{GeV}$. Besides the drift chamber, the Front Calorimeter (FCAL) was installed to provide additional massive active target for neutrino interactions. This FCAL consisted of 23 iron plates and 4.9 cm thick and separated by


Figure 4.3: Diagram of the NOMAD detector (top view) [5]
1.8 cm gaps. Twenty out of the 22 gaps were instrumented with long scintillators read out on both ends. The depth of FCAL is about 5 nuclear interaction length [48]. The total mass of FCAL in NOMAD is about 17.7 tons. This large mass provides high statistics. There are two main topics are studied with this FCAL: One is the production of opposite sign muon pairs ("dimuons"); The other one is the search for the heavy neutrinos ("sterile neutrinos").

### 4.3 Reconstruction and Simulation

In the NOMAD experiment, The drift chamber can be used to determine the event topology and to measure the momenta of charged particles. when the neutrinos interact with the target in the drift chamber, the trajectories of charged particles are reconstructed from the coordinate measurements provided by the drift chamber. In order to provide good measurement of the tracks, a very high efficiency of the
track reconstruction is required. Also the measured track parameters do not deviate significantly from the true particle momenta, i.e. the reconstruction program should provide good momentum resolution. The amount of ghost tracks should be minimized. Since in the drift chamber, the amount of matter crossed by a particle between two measurement planes cannot be neglected, the effects of energy losses and multiple scattering must be carefully taken into account [7].

To reconstruct the tracks, first, the pattern recognition (track search) should be performed to decide which individual measurements provided by the detector should be associated together to form an object representing a particle trajectory. Second, a fitting procedure should be applied to this set of measurements in order to extract the parameters describing the trajectory out of which the physical quantities can be computed [7].

To find the particle tracks, first, it is needed to guess possible tracks from hit combinatories and provide initial track parameters. Second, it is attempting to build a track from the given parameters by repeatedly collecting hits, fitting and rejecting possible outliers. The track is claimed to be fitted when no more hits can be added to it [7].

To simulate the Monte Carlo events in NOMAD, Neglib which was built based on LEPTO 6.1 [29] and JETSET [47] is used as the event generator. Rein-Sehgal (RS) model [44], Berger-Sehgal(BS) model [15] are used for the coherent event simulation.

### 4.4 Neutrino Interaction Candidate in NOMAD Target

Figure 4.4 shows the $\nu_{\mu}$ charged current interaction candidate in NOMAD Drift Chamber, we see that there are many hadron tracks, enabling the momentum vector measurements. $\mu$ is kinematically separated from the hadron vector.

Figure 4.5 to Figure 4.7 show some other neutrino interactions candidates in NOMAD detector.


Figure 4.4: $\nu_{\mu}$-CC candidates in NOMAD.

Most difficult to measure among the 4 neutrino species


Figure 4.5: $\bar{\nu}_{e}$ - CC candidates in NOMAD.

### 4.5 Coherent Signature in NOMAD Target

Some candidate coherent interaction events are shown from Figure 4.8 to Figure 4.10. In the charged current coherent process, there are two tracks which are identified as the leading lepton and coherent meson. A coherent $\rho^{0}$ candidate event is shown in Figure 4.8. There is a single, forward $\rho^{0}$ produced with no accompanying particles in a neutral charged coherent $\rho^{0}$ event. The $\rho^{0}$ in the final state will promptly decay into a charged pion pair. Therefore, in the NOMAD detector, there will be two tracks detected.


Figure 4.6: QE candidates in NOMAD.


Figure 4.7: Resonance candidates in NOMAD.


Figure 4.8: Coherent $\rho^{0}$ candidate event picture.


Figure 4.9: Coherent $\pi^{+}$candidate event picture.


Figure 4.10: Coherent $\pi^{-}$candidate event picture.

## Chapter 5

## A High Resolution Fine Grained Tracker (FGT)

## as the Near Detector for ELBNF/DUNE

### 5.1 Introduction \& Salient Physics Goals

In the ELBNF/DUNE project, besides the far detector, a near detector known as HIRESMNU, is also proposed with high resolution within a dipole magnetic field. This near detector is based on the NOMAD detector concept but with some improvements. It combines large statistics with high resolution (for example, momentum and energy) in the reconstruction of neutrino events compared with previous experiments. HIRESMNU is expected to achieve high precision in the measurements of neutrino interactions, structure of nucleons/nuclei, and the elements of the neutrino mass matrix. It is also designed to search for new physics, such as, sterile neutrinos, high $\Delta m^{2}$ oscillations, light dark matter etc. Hopefully, with HIRESMNU, there will be some unexpected discoveries made by the far detector. The fiducial mass of this near detector is 7 tonnes, bigger than the 2.7 tonnes of the NOMAD detector. Different from the NOMAD detector, there is a $4 \pi$ ECAL coverage in the dipole magnetic field and a $4 \pi \mu$ detector coverage instead of only downstream.

A summary of performance for the fine grained tracker configuration is shown in Table 5.1. All the parameters of the HIRESMNU are not fixed yet, but they have to meet our goals with the design [1].

Constraining the systematic uncertainties in the oscillation studies: The precision of the near detector (ND) measurements will be essential for the neutrino

Table 5.1: Summary of performance for the fine grained tracker configuration [1].

| Performance Metric | Value |
| :--- | :---: |
| Vertex resolution | 0.1 mm |
| Angular resolution | 2 mrad |
| $E_{e}$ resolution | $5 \%$ |
| $E_{\mu}$ resolution | $5 \%$ |
| $\nu_{\mu} / \bar{\nu}_{\mu}$ ID | Yes |
| $\nu_{e} / \bar{\nu}_{e} \mathrm{ID}$ | Yes |
| $\mathrm{NC} \pi^{0} /$ CCe rejection | $0.1 \%$ |
| $\mathrm{NC} \gamma / \mathrm{CCe}$ rejection | $0.2 \%$ |
| $\mathrm{NC} \mu / \mathrm{CCe}$ rejection | $0.01 \%$ |

oscillation studies ( $\nu \mathrm{OSCL}$ ) in ELBNF/DUNE, so the associated systematic error should be less than the corresponding statistical error. The required systematic precision in the near detector will determine the detector parameters such as resolution, fiducial mass, and so on. To this end, we will pay particular attention to [18]:
(1): Measurement of the relative abundance and energy spectrum of all four species of neutrinos in the LBNE beam: $\nu_{\mu}, \bar{\nu}_{\mu}, \nu_{e}$ and $\bar{\nu}_{e}$ via the $i n$ situ identification of their CC-interactions.
(2): Determination of the absolute $\nu$-flux using the $\nu$-electron scattering.
(3) Identification and precise measurement of $\pi^{0}$, photon, electron and positron yields in $\nu$-induced neutral-current (NC) and charge-current (CC) interactions - the most important background to the $\nu_{e}$-appearance.
(4) Measurement of NC cross-section relative to CC as a function of the hadronic energy, Ehad, since NC processes constitute the largest background to the $\nu$-CC identification.
(5) Measurement of the $\pi^{ \pm}$content in CC and NC hadronic jets since the $\pi^{ \pm} \rightarrow \mu^{ \pm}$ are the principal background to the $\nu_{\mu}\left(\bar{\nu}_{\mu}\right)$-CC.
(6) Measurement of the differential cross-section for various exclusive, semi-exclusive and inclusive channels relevant for the $\nu$ OSCL studies, such as quasi-elastic (QE), res-
onance (Res) and deep-inelastic (DIS) interactions.
(7) Quantification of nuclear-target material cross-section that might affect the $\nu$-nucleus interactions when extrapolating the near detector measurements to the far detector.

Precision neutrino physics: The HIRESMNU in ELBNF/DUNE is designed with a generational advance in the precision measurement of physics parameters, such as the precise measurement of isospin physics, sum-rules, QCD tests, baryon-spin, strange meson and baryon production, charm mesons and electroweak constants. As a case study, we propose to investigate the feasibility of a measurement of the weakmixing angle, $\sin ^{2} \theta_{W}$ in the $\nu(\bar{\nu})-\mathrm{q}(\mathrm{DIS})$ channel at a momentum transfer (Q) in the neighborhood of 4 GeV with a precision approaching $0.2 \%$. The sought precision on $\sin ^{2} \theta_{W}$ in this experiment will be comparable to that attained by the collider experiments. The HIRESMNU ND will also permit searches for new physics with unprecedented sensitivity; the searches include high- $\Delta m^{2}$ oscillations in $\nu_{\mu} \rightarrow \nu_{e}$, $\nu_{\mu} \rightarrow \nu_{\tau}, \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$ and $\nu_{e} \rightarrow \nu_{\tau}$ channels, rare resonance, heavy neutrinos, and exotic boson [18].

These considerations imply the following requirements:

- Magnetized tracker to identify positive from negative particles throughout the curvature of the particle tracks $(\mathrm{B} \sim 0.4 \mathrm{~T})$.
- Low density medium to track charged particles $\left(\rho \sim 0.1 \mathrm{~g} \mathrm{~cm}^{-3}\right)$.
- Large statistics ( $\sim 10^{8}$ neutrino interactions).


### 5.2 Sub Detectors

The HIRESMNU offers a generational advance in the identification of particles and the precise measurement of their momenta. The most precise neutrino detector to
date is NOMAD whose energy range was $2.5 \leq E_{\nu} \leq 300 \mathrm{GeV}$; the energy range of interest in ELBNF/DUNE is $0.5 \leq E_{\nu} \leq 100 \mathrm{GeV}$.

Figure 5.1 shows the sketch of the proposed HIRESMNU detector. In this figure, the incident neutrino beam comes from the left side. In Figure 5.1, we see that,


Figure 5.1: Sketch of the proposed HIRESMNU detector showing the inner STT and the $4 \pi$ ECAL in the dipole magnet with the muon-ID detector(MRD). The internal magnetic volume is approximately $4.5 \mathrm{~m} \times 4.5 \mathrm{~m} \times 8 \mathrm{~m}$. Also shown is one module of the proposed STT [18]
different from NOMAD with an ECAL (Electomagnetic Calorimeter) at the downstream end, and with a muon-ID detector outside the magnetic field, in HIRESMNU, the near detector (ND) has a $4 \pi$ ECAL coverage including downstream (DS), sides (Barrel) and upstream (US) ends of the detector. The tracking volume will be fully surrounded by electromagnetic calorimetry. Relative to the NOMAD detector, the HIRESMNU near detector has an enhanced tracking detector, that is composed of straw tubes with 1 cm in diameter. Vertical ( $\mathbf{Y}$ ) and horizontal ( $\mathbf{X}$ ) straws will be al-
ternated and arranged in modules with each module containing a double straw layer. The HIRESMNU ND also has an improved muon identification capability. Compared to the $85 \%$ efficiency of muon detection in NOMAD, HIRESMNU will tag $98 \%$ of the muons in the $\nu_{\mu}$ charged current sample. The trigger in HIRESMNU will have neither geometry bias nor charge bias as NOMAD does. HIRESMNU will accumulate about 60 million neutrino interactions in 5 years, a factor of 30 more events than NOMAD had.

## Straw Tube Tracker (STT)

The Straw Tube Tracker (STT) is the particle tracker of HIRESMNU, and locates at the center of the detector. The design of the straw tube follows that of the COMPASS detector [38, 43] similar to the dimensions of HIRESMNU, which is also a low density tracking detector, $\rho \leq 0.1 \mathrm{~g} \mathrm{~cm}^{-3}$. The conceptual transition radiation measurement in $\nu$-interactions is based upon NOMAD-TRD [12]; the design of the transition radiation (TR) detection follows that of the ATLAS Transition Radiation Tracker [4, 3, 2]. The Straw Tube Tracker (STT) will have a total 160 modules. Vertical (Y) and horizontal (X) planes of straws will be alternated and arranged in modules. Each module will consist of a double straw layer (either XX or YY) [18]. Figure 5.2 shows the layout of the straw layers and cross-section of an STT module.

Some parameters for the fine grained tracker are listed in Table 5.2.
Nuclear targets will be installed in the upstream end of the particle tracker (STT). Figure 5.3 shows the sketch of a basis STT module for the measurement of nuclear effects. This nuclear target provides a statistical robust sample $(\times 5$ the far detector statistics) to quantitate the differences in $\nu$-nuclear interaction. Argon gas, same as in the far detector, is proposed to use in pressurized tubes at the upstream end of the STT [18].

Figure 5.4 shows the schematic of the ATLAS transition radiation tracker (TRT).

Table 5.2: Parameters for the fine-grained tracker [1].

| Performance Metric | Value |
| :--- | :---: |
| STT detector volume | $3 \times 3 \times 7.04 \mathrm{~m}^{3}$ |
| STT detector mass | 8 tons |
| Number of straws in STT | 123,904 |
| Inner magnetic volume | $4.5 \times 4.5 \times 8.0 \mathrm{~m}^{3}$ |
| Targets | $1.27-\mathrm{cm}$ thick argon $\sim 50 \mathrm{~kg}$, |
|  | water and others |
| Transition radiation radiators | 2.4 cm thick |
| ECAL $X_{0}$ | 10 barrel, 10 backward, |
|  | 18 forward |
| Number of scintillator bars in ECAL | 32,320 |
| Dipole magnet | 2.4 -MW power, |
|  | $60-\mathrm{cm}$ steel thickness |
| Magnetic field and uniformity | 0.4 T; |
|  | $<2 \%$ variation over inner volume |
| MuID configuration | 32 RPC planes interspersed |
|  | between 20-cm thick layers of steel |

## Cross-Section of STT Module



Figure 5.2: Layout of the straw layers and cross-section of an STT Module [18].

The endplug for the STT will be similar to the ATLAS-TRT, which could enable the transition radiation measurement with a mixture of Xenon (70\%) and $\mathbf{C O}_{\mathbf{2}}(30 \%)$ To protect STT against humidity, $\mathbf{C O}_{\mathbf{2}}$ will be flushed through the straws at a forcedflow rate of $\sim 100 \mathrm{~m}^{3} \mathrm{~h}^{-1}$.


Figure 5.3: Sketch of a basis STT module for the measurement of nuclear effects. Several modules can be placed in the upstream magnetic volume with different target materials $(\mathrm{Pb}, \mathrm{Fe}$ etc.) of the same thickness in radiation length [18].

## Electromagnetic Calorimeter (ECAL)

The tracking volume of HIRESMNU (STT) will be surrounded by a $4 \pi$ ECAL coverage: the forward or downstream (DS) module, the four side (top-bottom and leftright) of the Barrel module, and the upstream (US) module. This ECAl is a leadscintillator calorimeter based upon the T2K-ECAL and have transverse and longitu-


Figure 5.4: Schematic of the ATLAS STT Endplug [18].
dinal segmentation. The ECAL-surrounded STT will be embedded inside the dipole magnet [18].

The most important component of the ECAL in HIRESMNU near detector is the forward, or downstream (DS), module (shown in Figure 5.5) composed of 58 layers of $10-\mathrm{mm}$-thick (along z -direction) scintillator (Sci) followed by $1.75-\mathrm{mm}$-thick lead, corresponding to 10 radiation lengths $\left(X_{0}\right)$. The first layer will be composed of 160 horizontal scintillator bars (providing the Y-coordinate of particles), $400-\mathrm{cm}$ long, $2.5-\mathrm{cm}$ wide, and $1-\mathrm{cm}$ thick; followed by the Pb sheet. The $2.5-\mathrm{cm}$ width of the bar is informed by the Moliere-radius of a 2.5 GeV electron-typical of $\nu_{e}$-induced CC- which is approximately 2 cm . The second layer will be composed of 160 vertical scintillator bars (providing the X-coordinate), $400-\mathrm{cm}$ long, $2.5-\mathrm{cm}$ wide, and $1-\mathrm{cm}$ thick; followed by the Pb sheet. The third layer will be a Y-plane of bars, vertically shifted by 1.25 cm ; followed by Pb sheet; The fourth layer will be a X-plane of bars, horizontally shifted by 1.25 cm ; followed by Pb sheet. This arrangement will repeat itself to complete the DS-ECAL module [18].

Figure 5.5 shows a preliminary schematic of the DS-ECAL containing a total of

9280 Sci-bars with a total Sci-volume of $4.64 \mathrm{~m}^{3}$. Figure 5.6 shows the engineering details of the module.


Figure 5.5: Preliminary Schematic of the DownStream or Forward ECAL [18].

At a national lab or fabrication center in India the scintillator bars will be extruded with Kuraray wavelength-shifting fibers and later threaded through the middle of the bars. The fibers will be read out at each end by silicon photomultiplier (SiPM) type photosensors. The two readings allow for a position determination. It follows that the number of readout channels will be twice the number of scintillator bars. In the DS-ECAL, there will be 18,560 SiPMs. The T2K-ECAL used the Hamamatsu photosensors called MPPC which will be adequate for our use. It must be noted that the SiPM technology has undergone a rapid improvement, driving the costs lower and offering better performance.

Since the Barrel-ECAL will surround the sides of the STT, nominally there would be four modules, covering a $8 \mathrm{~m} \times 4 \mathrm{~m}$ area, corresponding to the top, bottom, left


Figure 5.6: DownStream(DS) or Forward ECAL [18].
and right sides. The conceptual design of the Barrel-ECAL is similar to that of the DS-ECAL. A Barrel-ECAL module, Figure 5.7, will have eight layers of alternating horizontal and vertical scintillator strips every 7 mm of lead, or $10-X_{0}$ deep. The upstream (US) ECAL will be identical to one of the Barrel-ECAL modules (see Figure 5.8). For the scintillator bars, our default assumption is that the dimensions of the bars in the Barrel-ECAL and US-ECAL will remain similar to those in the DS-ECAL, i.e., $400-\mathrm{cm}$ long, $2.5-\mathrm{cm}$ wide, and $1-\mathrm{cm}$ thick. The total number of bars in the eight modules of Barrel-ECAL and the one module of US-ECAL will be 11,520 (4-layers $\times 400 \mathrm{~cm} / 2.5 \mathrm{~cm} \times 9$ ) correspond ing to 30,000 readout channels.

We have given some thought to reducing the number of channels in the ECAL. Our preliminary optimization studies include (a) reducing the longitudinal granularity in the most important portion, the DS-ECAL; (b) reducing its transverse granularity; and (c) reducing the transverse granularity of the Barrel-ECAL. A detailed GEANT4-
based MC will provide us with better guidance.


Figure 5.7: Specifications of the Barrel ECAL [18].


Figure 5.8: UpStream(UP) ECAL [18].

Table 5.3: Parameters of the UA1/NOMAD dipole magnet [18].

| Item | Value |
| :--- | :---: |
| Dimension | $3.5 \times 3.5 \times 7 \mathrm{~m}^{3}$ |
| Maximum B-Field | 0.7 T |
| Maximum Current | $10,000 \mathrm{~A}$ |
| Resistance(40C) | 0.0576 |
| Voltage | 576 V |
| Mass | 900 T |
| Cooling water flow | $50 \mathrm{liters} / \mathrm{sec}$ |
| Pressure gredient | 15 atm (in .vs. out) |
| Temperature Diff | 30 C |

## Dipole Magnet

In HIRESMNU near detector, the STT particle tracking detector and ECAL modules will both locate inside a $0.4-\mathrm{T}$ dipole magnet with inner dimensions of $4.5-\mathrm{m}$ wide by $4.5-\mathrm{m}$ high by $8.0-\mathrm{m}$ long. The proposed dipole magnet is a larger version of the UA1 dipole magnet [5] which is used by NOMAD, and is currently in use by the near detector of T2K. Because the dipole magnet for HIRESMNU is very similar to the UA1 magnet yoke assembly, Figure 5.9 shows a photograph of the specifications of the UA1 magnet during the NOMAD operation. Table 5.3 summarizes the salient parameters of this magnet at the maximum operating B field of 0.7 T . The HIRESMNU magnet will operate at 0.4 T .

The principal differences between the UA1/NOMAD and the HIRESMNU magnets are: (a) size; (b) coil-UA1 used Al-coil to minimize the degredation of the energy resolution of the outgoing jets with resistivity $\rho_{A l}=2.8 \times 10^{-8} \Omega \mathrm{~m}$; and (c): HIRESMNU magnet does not need a hole in the coil to allow passage of the proton beam, as was the case with the UA1 magnet. The HIRESMNU dipole will be composed of $8+8$ ' C '-Sections, the iron yokes. The coil will be made of copper, with aluminium as another option. Figure 5.10 shows a conceptual engineering drawing of the proposed magnet and 'C' Section.


Figure 5.9: Specs of the UA1/NOMAD dipole magnetic [18].

## Resistive Plate Chamber as the Muon Detector (RPC)

In the HIRESMNU design, the muon-ID detector will identify muon tracks which will then be matched with tracks in the STT with measured momenta and charges. Thus, the muon-ID detector is not required to furnish muon momentum, only the $\mu$-ID. However, given the large rate of muons in the ND location, we need a muon-ID detector with such spatial and time resolutions as to precisely reconstruct $\mu$-track segments and permit an unambiguous match with the muon track in the STT. The muon-ID detector can be divided into two systems [18]:


Figure 5.10: Conceptual sketch of one of the C-sections that constitute the magnet return yoke (dimensions are in mm ). The vertical dimension is longer that the horizontal one in order to accommodate the magnet coil [18].

- The Muon Range Detector (MRD) instruments the gaps between the plates of the magnet return yoke. The main task of the MRD is to identify muons at low momenta exiting the sides of the detector. The MRD will reconstruct track segments within the magnet return yoke, including those of the stopping (ranging-out) muons [18].
- The External Muon Identifier (EMI) will identify high-energy forward muons. It is located outside the magnet, at the downstream end of the detector. The EMI will consist of two stations separated by a passive concrete/iron absorber. At each station, outside and downstream of the dipole magnet, it will reconstruct muon track segments to be matched with the STT tracks.

Due to the multiple scattering in the material (mainly iron) crossed by the muon tracks reaching both the MRD and the EMI detector, a space resolution in the range of $\simeq 0.75 \mathrm{~mm}$ will be adequate to accomplish the tracking task of the muon-ID detectors. For both the MRD and EMI detectors, we have selected the same Resistive Plate Chamber (RPC) technology as that developed for the LHC experiments, OPERA, BaBar and Argo. (An alternative under consideration is glass-based RPCs.) In particular, we follow the design of the RPC detectors used in the OPERA experiment to instrument the gaps within the magnet of the Muon Spectrometer. This type of application is similar to our MRD and has been operational for a few years in OPERA. A sketch of the detector is shown in Figure 5.11. Two electrodes, made of 2 mm Bakelite with linseed oil and volume resistivity $\rho>5 \times 10^{11} \Omega \mathrm{~cm}$ at $\mathrm{T}=20^{\circ} \mathrm{C}$, are kept 2 mm apart by means of polycarbonate spacers in a 10 cm lattice configuration. The external surface of the electrodes is painted with graphite of high surface resistivity and protected with a $190 \mu \mathrm{~m}$ thick PET layer that is applied during the installation on each side of the RPC to prevent high voltage discharge. The inner surface of the electrodes is coated with a few-micron-thick polymerized linseed oil layer. The total thickness of an RPC is between 6 and 7 mm [18].


Figure 5.11: Cross-section of a Resistive Plate Chamber with its associated strips for the read out of the induced signal [18].


Figure 5.12: Layout of the HIRESMNU with downstream, external muon-ID detector(EMI) and shielding. The EMI specifications are preliminary; detailed simulation studies of EMI will yield a more robust design [18].


Figure 5.13: Momentum and Energy resolution of HIRESMNU [18].

### 5.3 How the FGT Helps Accomplish the Above Goals

The FGT offers a generational advance in the precision of the individual particle momentum measurements and identification. The most precise neutrino detector to
date is NOMAD whose energy range is $2.5 \leq E_{\nu} \leq 300 \mathrm{GeV}$; the energy range of interest in LBNE is $0.5 \leq E_{\nu} \leq 100 \mathrm{GeV}$. The advancements of the FGT compared with current precision measurement are:

- An Enhanced Tracking Detector.
- $4 \pi$ Electromagnetic Calorimeter Coverage.
- Improved Muon-Identification.
- No Geometry Bias or Charge Bias Trigger.
- High Event Rate.


## Chapter 6

## Coherent $\pi^{-}$Measurement in NOMAD

### 6.1 Signal Signature



Figure 6.1: Feynman Diagram of Coherent Pion Production.

Figure 6.1 shows the Feynman diagram of coherent $\pi^{-}$process $\left(\bar{\nu}_{\mu}+\mathcal{N} \rightarrow \mu^{+}+\right.$ $\left.\pi^{-}+\mathcal{N}\right)$. In the coherent $\pi^{-}$events, there are only two charged particles in the final state, which means in the drift chamber of NOMAD detector, we are looking for the events with the following properties:

- There are only two tracks observed. The one which is longer and triggered the Muon chamber at downstream is identified as Muon.
- The events are all with a low value of $Q^{2}$ which is defined in Equation $(3.1)\left(Q^{2} \leq\right.$ $\left.1 G e V^{2}\right)$.
- In the selected signal events, the direction of the outgoing lepton is collinear with the incoming neutrino direction.


### 6.2 Background

The background of coherent $\pi^{-}$events comes from several sources:

- The biggest background comes from $\nu_{\mu}$ charged current events. In some of the $\nu_{\mu}$ charged current events, $\pi^{-}$are also produced, and could be identified as coherent $\pi^{-}$events especially when other particles (except $\pi^{-}$and $\mu^{+}$) have very small energy.
- Another background comes from coherent $\rho^{0}$ events: The decay of the $\rho^{0}$ produces $\pi^{+}$and $\pi^{-}$, which might also be classified as coherent Pion events.
- Neutral current events are also a background source in coherent $\pi^{-}$analysis when there are only two tracks observed and one of them is classified as Muon.


### 6.3 Neutrino and Anti-neutrino Beams

In the NOMAD experiment, the neutrino beam is produced by extracting part of the 450 GeV proton beam circulating in the SPS (Super Proton Synchrotron) and allowing it to interact with a beryllium target. Figure 4.2 show the schematic layout of the the WANF beam line[8]. The two toroidal magnetic lenses, referred to as the horn and the reflector, focused charged particles of a given sign (positive for a predominantly $\nu_{\mu}$ beam); The polarity of these magnetic elements could be changed within minutes in order to produce an anti-neutrino beam ( $\bar{\nu}_{\mu}$ beam) [8], which are called neutrino beam mode (positive focusing data: FocP) and anti-neutrino beam mode (negative focusing data: FocN).

The principle of the focusing is illustrated in Figure 6.2. The reflector provided additional focusing for high momentum particles by horn. The magnetic field was
provided by current sheets flowing in the inner and outer conductors of the lenses. The field was measured to be azimuthally symmetric to better than $1.5 \%$. Its value at a radial position r from the beam axis and for a current I is given by [8]

$$
\begin{equation*}
B=\frac{\mu_{0} I}{2 \pi r} \tag{6.1}
\end{equation*}
$$

The current ( 100 kA for the horn and 120 kA for the reflector) was provided by the discharge of capacitor banks and lasted 6.8 ms . The thickness of the inner conductor was minimized to reduce secondary interactions while maintaining adequate strength to withstand the magnetic forces.


Figure 6.2: Principle of the focusing. The lines are representative trajectories of particles of three different momenta [8].

Positively charged particles (mainly $\pi^{+}$and $K^{+}$mesons) produced around zero degrees with respect to the primary proton beam are focused into a near parallel beam by a system of magnetic lenses and subsequently decay producing neutrinos.

The effect of the horn and the reflector on particles of different sign is illustrated in Figures 6.3 and 6.4, which show angular distributions of positive and negative pions at a plane just upstream of the horn and immediately downstream of it. Upstream of the horn, pions of both positive and negative charges emerging from the target have very similar angular distributions, with the bulk of the particles within 10 mrad, which is the acceptance of the collimators. While traversing the horn, positive pions with both momentum around $50 \mathrm{GeV} / \mathrm{c}$ are focused into a near-parallel beam leading to an overall enhancement at small angles of up to a factor of 30 (Figure 6.3).


Figure 6.3: Distribution of the angle between the $\pi^{+}$momentum vector and the beam line direction, $P_{T} / P$, just upstream of the horn (top left),right after it(top right) and the ratio of the latter to the former(bottom) [8].

Negative pions are strongly defocused resulting in their reduction at small angles by as much as a factor of 5 (Figure 6.4). The reflector provides an additional focusing for positive particles of momentum both higher and power than $50 \mathrm{GeV} / \mathrm{c}$ that were respectively underfocused and overfocused by horn [8].

### 6.4 Neutrino Beam Mode

In the NOMAD data, all of the incoming neutrinos originating from the neutrino beam mode are dominated by the Muon neutrinos $\nu_{\mu}$, but still with a about $3 \%$ percentage contamination of the anti-neutrinos $\bar{\nu}_{\mu}$. At first, we are going to measure


Figure 6.4: Distribution of the angle between the $\pi^{-}$momentum vector and the beam line direction, $P_{T} / P$, just upstream of the horn (top left),right after it(top right) and the ratio of the latter to the former(bottom) [8].
the coherent $\pi^{-}$events in the neutrino beam mode.

## Normalization of Events Samples

In the analysis, to ensure the Monte Carlo matches well with the NOMAD experimental data, all the Monte Carlo components are initially normalized based on coherent events and inclusive charged current events.

The initial normalization factors in the analysis are shown in Table 6.1 and Table 6.2. Where $1,436,000$ is the number of the $\nu_{\mu}$ inclusive charged current events from the reference [52]. The ratio between neutral current and charged current $\nu_{\mu}$

Table 6.1: Normalization of Neutrino Beam Mode events [52, 8].

| Mode | Event Number |
| :--- | :---: |
| $\nu_{\mu} \mathrm{CC}$ | $1,436,000$ |
| $\nu_{\mu} \mathrm{NC}$ | $1,436,000 \times 0.37$ |
| $\bar{\nu}_{\mu} \mathrm{CC}$ | $1,436,000 \times 0.025$ |
| $\bar{\nu}_{\mu} \mathrm{NC}$ | $1,436,000 \times 0.025 \times 0.37$ |

Table 6.2: Ratios between interaction mode.

| Ratio | Value |
| :---: | :---: |
| NC/CC | 0.37 |
| non-DIS/DIS | 0.055 |
| QE/RES | 0.75 |

events also comes from this work, but not published yet. These initial normalization factors should provide us with a reasonable description of the data. Later we will fine-tune certain normalization factors to improve the fit.

## Preselection of $\bar{\nu}_{\mu}$ CC and Coh $\pi$ Events

To select the $\bar{\nu}_{\mu}$ charged current process, some variables cuts were applied during this analysis, including:

- Fermi Momentum Cut: Pfermi Cut( $<1.0 \mathrm{GeV}$ ): Fermi Momentum(Pfermi) cut is applied to all the Monte Carlo events and helps us remove the non-physical events.
- W2s Cut: Invariant hadronic mass (W2s: $W^{2}=(q+P)^{2}>1.96 \mathrm{GeV}^{2}$ ), which is only applied to deep inelastic charged current and neutral current events, which will remove most of the background from deep inelastic interactions.
- Fiducial Volume: The drift chamber (DC) target is $3 \times 3 \times 4 m^{3}$ and perpendicular to the neutrino beam direction. A reduced volume is necessary because some interactions in the magnet can create vertexes close to or outer the edges of
the drift chamber active area. In my research, the fiducial volume chosen is $|\mathrm{X}|<130 \mathrm{~cm},|\mathrm{Y}-5|<130 \mathrm{~cm}$, and $\mathrm{Z}<405 \mathrm{~cm}$ with minimum z varying depending on experiment setup.

Table 6.3: Minimum value for the fiducial volume in NOMAD.

| run No. | $\leq 8375$ | $[8376,9344]$ | $[9345,14164]$ | $\geq 14165$ |
| :--- | :---: | :---: | :---: | :---: |
| $\min \mathrm{z}(\mathrm{cm})$ | 265 | 115 | 5 | 35 |

- Muon Identification: Since this analysis is searching for coherent $\pi^{-}\left(\bar{\nu}_{\mu}+\right.$ $\left.\mathcal{N} \rightarrow \mu^{+}+\pi^{-}+\mathcal{N}\right)$, Muon identification $\left(\mu^{+}\right)$becomes very critical. The first identification is a loose identification of Phase II which means that a track in drift chamber can be matched to a hit in the Muon chamber within 40 cm in the first station or within 50 cm in the second condition. This identification is stored in the DSTs (Data Summary Type). Then, the selection of Muon can be tightened by considering the goodness of fit between the track and the hit in the Muon chamber, and accepting only those combinations with the total $\chi^{2}$ less than 20, over the four degrees of freedom of the fit.
- Tube/Veto Cut: Tube/Veto cut is applied to remove the charged particles and only let the neutral particles pass through the drift chamber. Then, the events originating mainly from the Muon contaminants can be rejected.
- Track Number Cut (ncand): In the $\bar{\nu}_{\mu}$ charged current interactions, there are two or more than two charged tracks observed. In coherent $\pi^{-}$interactions, there are only two charged tracks observed, including Muon and Pion.
- Muon Momentum $\left(\vec{P}_{\mu}\right)$ Cut: To ensure the quality of the momentum's measurement, the Muon momentum should be greater than 1.5 GeV .
- Charge Identification: In a coherent $\pi^{-}$event, the hadron is $\pi^{-}$. To select the events with a negative hadron, Charge Identification is applied.


Figure 6.5: Reconstruction of the event kinematics for individual events.

- Angle Cut: To select the coherent $\pi^{-}$events, 2 angle cuts have been applied. The angle $\theta$ which is the angle between the outgoing lepton $\mu$ and the meson $\pi$, should be less than $177.5^{\circ}$ and greater than 0.5 Rad .
- Missing Transverse Momentum $\left(P_{T}^{m}\right)$ Cut: Figure 6.5 shows the reconstruction of the event kinematics for individual event. To choose the coherent events, the transverse momentum of the lepton must be less that 0.5 GeV .
- Neutral Vertex and Cluster Cut: neutral vertex and cluster cut are used to cut off the the coherent $\rho^{0}$ background especially when the two neutral vertex or cluster are both from the primary vertex.

$$
\begin{gather*}
|t|=\left[\left(E_{\mu}-p_{\mu}^{z}\right)+\left(E_{\pi}-p_{\pi}^{z}\right)\right]^{2}+\left[\left(p_{\mu}^{x}+p_{\pi}^{x}\right)+\left(p_{\mu}^{y}+p_{\pi}^{y}\right)\right]^{2}  \tag{6.2}\\
\left|t_{\min }\right|=\left[\left(Q^{2}+m_{\pi}^{2}\right) / 2 E_{\pi}\right]^{2} \tag{6.3}
\end{gather*}
$$

$$
\begin{equation*}
t^{\prime}=\left|t-t_{\min }\right| \tag{6.4}
\end{equation*}
$$

In the analysis, besides the variables mentioned in the preselection section, there are three other variable are also used and very important in the analysis of coherent processes, they are $|t|,\left|t_{\text {min }}\right|$, and $\left|t^{\prime}\right|$ which are defined in Equation (6.2), Equation (6.3) and Equation (6.4). Where $|t|$ represents the magnitude of the square of nucleus' 4-momentum transferred to the nucleus.

Table 6.4: Cut table of Monte Carlo events in generated level.

|  | $N_{\bar{\nu}}^{C C}$ | $N_{\nu}^{C C}$ | $N^{N C}$ | TotBkg | $N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| Total | 47790.3 | 1645469 | 628612.1 | 2274081 | 2321871 |
| Pfermi $<1.0$ | 47380.6 | 1631201 | 623116.4 | 2254318 | 2301698 |
| W2s $>1.96($ CCDIS $)$ | 40610.1 | 1555351 | 589725.7 | 2145077 | 2185687 |
| $\mid$ GenX, (GenY-5) $\mid<130 \mathrm{~cm}$ | 36076.4 | 1436828 | 544853.6 | 1981682 | 2017758 |
| Zmin $<$ GenZ $<405 \mathrm{~cm}$ |  |  |  |  |  |
| $2.5<$ Evis $<300$ | 35900.0 | 1436001 | 544603.7 | 1980605 | 2016505 |

Table 6.4 shows the preselection of $\bar{\nu}_{\mu}$ charged current events in generated level. The normalization factor to the total inclusive $\bar{\nu}_{\mu}$ charged current events is $0.025 \times$ $1.436 \times 10^{6}$ after normalization, Z-weight, and flux(beam) reweight. $N_{\bar{\nu}_{\mu}}^{C C}$ represents the number of $\bar{\nu}_{\mu}$ charged current events including $\bar{\nu}_{\mu}$ charged current deep inelastic events $\left(\bar{\nu}_{\mu}\right.$ CCDIS $), \bar{\nu}_{\mu}$ charged current quasi-elastic events $\left(\bar{\nu}_{\mu} \mathrm{CCQE}\right)$, and $\bar{\nu}_{\mu}$ charged current resonance events ( $\left.\bar{\nu}_{\mu} \mathrm{CCRES}\right) . N_{\nu_{\mu}}^{C C}$ represents the number of $\nu_{\mu}$ charged current events including $\nu_{\mu}$ charged current deep inelastic events $\left(\nu_{\mu}\right.$ CCDIS $), \nu_{\mu}$ charged current quasi-elastic events $\left(\nu_{\mu} \mathrm{CCQE}\right)$, and $\bar{\nu}_{\mu}$ charged current resonance events ( $\nu_{\mu}$ CCRES). $N^{N C}$ represents the number of $\bar{\nu}_{\mu}$ neutral current events and $\nu_{\mu}$ neutral current events. It is obvious that, after all the normalization, the total $\nu_{\mu}$ charged current events has been normalized to $1,436,000$; and the total $\bar{\nu}_{\mu}$ charged current events has been normalized to 35,900 , which is equal to $1,436,000 \times 0.025$.

Table 6.5: Reconstructed variable cut table after normalization and reweighted.

|  | $N_{\bar{D}}^{C C}$ | $N_{\nu}^{C C}$ | $N^{N C}$ | TotBkg | $N_{M C}$ | $N_{\text {data }}$ | $N_{\text {data }} / N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Total | 38919.5 | 1434877 | 464599.6 | 1899476 | 1938396 | 4018980 | 2.073 |
| Pfermi $<1.0$ | 38595.7 | 1422826 | 460710.6 | 1883537 | 1922133 | 4018980 | 2.091 |
| W2s $>1.96$ (CCDIS) | 33404.4 | 1357502 | 459376.8 | 1816879 | 1850283 | 4018980 | 2.169 |
| FV cut | 27615.3 | 1202210 | 405951.5 | 1608162 | 1635777 | 2515946 | 1.538 |
| PhaseII | 25478.2 | 1065825 | 14718.7 | 1080544 | 1106022 | 1151014 | 1.041 |
| Nmu $=1$ | 25361.4 | 1058265 | 12319.5 | 1070584 | 1095946 | 1138941 | 1.039 |
| veto $/$ tube | 25234.0 | 1055009 | 12263.2 | 1067273 | 1092507 | 1109323 | 1.015 |
| ncand $>=2$ | 23411.5 | 1029563 | 12179.9 | 1041743 | 1065155 | 1079287 | 1.013 |
| $\left\|\vec{p}_{\mu}\right\|>2.5$ | 23277.8 | 1023853 | 11742.4 | 1035595 | 1058873 | 1072094 | 1.012 |
| $\mu^{+}$ | 23172.5 | 8257.5 | 6011.7 | 14269.3 | 37441.7 | 52834 | 1.411 |
| DeltaP $/ \mathrm{P}<=0.2$ | 23096.9 | 7311.7 | 5819.0 | 13130.7 | 36227.7 | 50085 | 1.383 |
| thetamupi $<177.5$ | 23094.4 | 7311.7 | 5818.4 | 13130.1 | 36224.5 | 47549 | 1.313 |
| Evis $<300 \mathrm{GeV}$ | 23093.2 | 7309.9 | 5817.5 | 13127.4 | 36220.6 | 47119 | 1.301 |
| Ehad $<300 \mathrm{GeV}$ | 23093.2 | 7309.3 | 5817.3 | 13126.6 | 36219.8 | 47117 | 1.301 |
| Nuhat $>0$ | 19838.1 | 2321.7 | 518.2 | 2840.0 | 22678.1 | 24898 | 1.098 |
| Ybj $<0.5$ | 16157.8 | 585.7 | 250.4 | 836.1 | 16994.0 | 18060 | 1.063 |
|  |  |  |  |  |  |  |  |

Table 6.5 shows the preselection of $\bar{\nu}_{\mu}$ charged current events in reconstructed level. From this table, it is obvious that there is still about $6.3 \%$ disagreement between Monte Carlo events and NOMAD data after all the cuts. We are going to vary the normalization factors to get better matches between Monte Carlo events and NOMAD data.

To ensure that the Monte Carlo matches the NOMAD data well, the following method was used to fit the Monte Carlo events with NOMAD data. The initial step needed to do is to choose the variables will be used to fit. The principle of choosing the variables used for fitting is they should give a good separation of signal from background. After checking the distributions of some kinematic variables, the three variables selected for fitting are as follows: $\hat{\nu}$ ( $\hat{\nu}$ is defined in Equation 6.5), Ybj ( Ybj is defined in Equation 6.6), and the missing transverse momentum $P_{T}^{m}$. The normalization factors for $\bar{\nu}_{\mu}$ charged current events, $\nu_{\mu}$ charged current events, and neutral current events are $k_{n o r m}^{\bar{\nu}_{\mu} C C}, k_{n o r m}^{\nu_{\mu} C C}$ and $k_{\text {norm }}^{N C}$.

$$
\begin{gather*}
\hat{\nu}=\frac{P_{T}^{l e p}-P_{T}^{m}}{P_{T}^{\text {hadron }}}  \tag{6.5}\\
y=\frac{p \cdot q}{k \cdot p}=\frac{M\left(E-E^{\prime}\right)}{E M}=1-\frac{E^{\prime}}{E}=\frac{\nu}{E}, \tag{6.6}
\end{gather*}
$$

where the variables, $\mathrm{p}, \mathrm{q}, \mathrm{E}, \mathrm{E}^{\prime}, \nu$ have been defined at the beginning of chapter 3 .
The procedure includes the following steps:

- Create the 3D distribution with respect to missing transverse momentum $P_{T}^{m}$, Bjorken variable Ybj , and $\hat{\nu}$;
- Divide the distribution into equi-populated $5 \times 5 \times 5=125$ cubes. Because the events are not distributed uniformly, the volumes of the cubes are not the same.
- Run through all the possible normalization:
- Set the default value of $k_{\text {norm }}^{\bar{\nu}_{\mu} C C}$ at 0.9 and vary it in steps of 0.001 from 0.9 to 1.1;
- Set the default value of $k_{\text {norm }}^{N C}$ at 1.9 and vary it in steps of 0.001 from 1.9 to 2.5 ;
- Set the default value of $k_{\text {norm }}^{\nu_{\mu} C C}$ at 1.5 and vary it in steps of 0.001 from 1.4 to 1.5 ;

Then the total number of Monte Carlo events after normalization becomes

$$
\begin{aligned}
N_{M C}= & N_{\bar{\nu}_{\mu}}^{C C} \times k_{\text {norm }}^{\bar{\nu}_{\mu} C C}+ \\
& N^{N C} \times k_{\text {norm }}^{N C}+ \\
& N_{\nu_{\mu}}^{C C} \times k_{\text {norm }}^{\nu_{\mu} C C}
\end{aligned}
$$

- Calculate the $\chi^{2}$ for each cube and sum over all of them.
$\chi^{2}=\frac{\left(N_{\text {data }}-N_{M C}\right)^{2}}{\sigma_{\text {data }}^{2}+\sigma_{M C}^{2}}$
While $\sigma_{\text {data }}=\sqrt{N_{\text {data }}}$, and $\sigma_{M C}=\sqrt{\sum_{j}\left(\sqrt{N_{M C}(j)}\right)^{2}}$, index j represents the Monte Carlo events $\left(\bar{\nu}_{\mu} C C, \nu_{\mu} \mathrm{CC}\right.$ and NC$)$.
- Get the parameter corresponding the minimum $\chi^{2}$.

$$
k_{\text {norm }}^{\bar{\nu}_{\mu} C C}=0.961 \pm 0.018
$$

(shown in Figure 6.6 )

$$
k_{N o r m}^{N C}=2.052 \pm 0.052
$$

(shown in Figure 6.7)

$$
k_{\text {norm }}^{\nu_{\mu} C C}=1.445 \pm 0.051
$$

(shown in Figure 6.8)
Since the total number of points I used for fit is 125 , the minimum $\chi^{2}$ equals to 979.61 , then the value $\chi^{2} / \mathrm{DOF}$ is 7.837 (DOF means degree of freedom). Compared to the $\chi^{2}$ before this 3D fit, it is reduced.


Figure 6.6: $\chi^{2}$ distribution with respect to $k_{n o r m}^{\bar{\nu}_{n} C C}$, fix $k_{\text {norm }}^{N C}=1.062, k_{n o r m}^{\nu_{\mu} C C}=1.445$ using 125 points.

- After fitting with the 3D distribution of $P_{T}^{m}, \mathrm{Ybj}$, and $\hat{\nu}, \mathrm{I}$ also created the 2D distributions of Ybj and $\hat{\nu}$, and fitted the MC events to the Data as a check. The normalization factors I got from this 2D fit are
$k_{n o r m}^{\bar{\nu}_{\mu} C C}=1.103 \pm 0.016 ;$
$k_{\text {norm }}^{\prime \nu_{\mu}}=0.989 \pm 0.016 ;$
where $k_{\text {norm }}^{\prime} \nu_{\mu}$ is applied on both the $\nu_{\mu}$ charged and neutral current events. The total points I used for this 2 D fit is 50 , and The minimum $\chi^{2}$ is 228.125 , then the $\chi^{2} /$ DOF equals to 4.56 . Which means the $\chi^{2} /$ DOF is reduced from 7.80 to


Figure 6.7: $\chi^{2}$ distribution with respect to $k_{\text {norm }}^{N C}$, fix $k_{\text {norm }}^{\bar{\nu}_{\mu} C C}=0.961, k_{\text {norm }}^{\nu_{\mu} C C}=1.445$ using 125 points.
4.56 by this 2D fit. Then the new full normalization factors I got after the 2 D fit are
$p_{n o r m}^{\bar{\nu}_{\mu} C C}=0.961 \times 1.103=1.06 ;$
$p_{\text {norm }}^{N C}=2.052 \times 0.989=2.03 ;$
$p_{n o r m}^{\nu_{\mu} C C}=1.445 \times 0.989=1.43 ;$

- After fitting with the 2 D distribution of Ybj and $\hat{\nu}$ cut, I refitted the $\bar{\nu}_{\mu}$ charged current events again with Ybj distribution as a check. For this 1D fit, all the events are divided into 5 equi-populated bins according to the Ybj distribution.

The minimum value of $\chi^{2}$ is 25.600 , then the value $\chi^{2} /$ DOF equals to 5.120 . Compared to $\chi^{2} /$ DOF before this 1D fit, it is reduced by a factor of 1.98 . $p_{\text {norm }}^{\bar{\nu}_{\mu} C C}=0.961 \times 1.103 \times 0.975=1.03 ;$


Figure 6.8: $\chi^{2}$ distribution with respect to $k_{\text {norm }}^{\nu_{\mu} C C}$, fix $k_{\text {norm }}^{\bar{\nu}_{\mu} C C}=0.961, k_{\text {norm }}^{N C}=2.052$ using 125 points.

$$
\begin{aligned}
& p_{\text {norm }}^{N C}=2.052 \times 0.989=2.03 \\
& p_{\text {norm }}^{\nu_{C} C C}=1.445 \times 0.989=1.43
\end{aligned}
$$

- Get the total normalization factor
$p_{n o r m}^{\bar{\nu}_{\mu} C C}=0.961 \times 1.103 \times 0.975=1.03$, which means the total $\bar{\nu}_{\mu}$ charged current events

$$
N_{\bar{\nu}_{\mu}}^{C C}=1436000 \times 0.025 \times 1.03=36977 ;
$$

This fit procedure is only used to define the center value of the background scale factors. The large value of $\chi^{2}$ from the fit can be explained by the systematic uncertainties which were not included in the $\chi^{2}$ calculation. These systematic uncertainties on the background subtraction will be estimated using a different ways of background normalization (which will be shown in section 6.8).

The first two columns of Table 6.6 show the number of generated and reconstructed $\bar{\nu}_{\mu}$ charged current events in 7 visible energy (Evis) bins. By calculating the ratio between the reconstructed and generated values in each bin, we can get the corresponding efficiency, which are shown in the last column of Table 6.6. The elements in the last column of Table 6.6 construct the 7 dimensional efficiency vector, and it is going to be used to calculated the corrected signal. Besides the efficiency vector, the efficiency matrix is also used to calculate the fully corrected signal.

Table 6.6: The efficiency (ratio between Reconstructed $\bar{\nu}_{\mu} \mathrm{CC}$ and Simulated $\bar{\nu}_{\mu} \mathrm{CC}$ events) in $7 E_{\nu}$ bins.

| Evis $(\mathrm{GeV})$ | Generated $N_{\bar{\nu}}^{C C}$ | Reconstructed $N_{\bar{\nu}}^{C C}$ | Efficiency |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 3970.6 | 1160.3 | 0.292 |
| $8.0-15.0$ | 6179.3 | 2566.7 | 0.415 |
| $15.0-20.0$ | 3933.6 | 1777.5 | 0.452 |
| $20.0-30.0$ | 6098.3 | 2905.4 | 0.476 |
| $30.0-50.0$ | 7618.3 | 3782.4 | 0.496 |
| $50.0-100.0$ | 6608.1 | 3288.7 | 0.498 |
| $100.0-300.0$ | 1491.9 | 677.0 | 0.454 |
| $2.5-300.0$ | 35900.0 | 16157.8 | 0.450 |

Table 6.7 and Table 6.8 show the generated, reconstructed $\bar{\nu}_{\mu}$ charged current events, charged current (CC) background, neutral current (NC) background, number of data before normalization with the factor from fit, and raw signal, reconstructed $\bar{\nu}_{\mu}$ charged current events, background after normalization with respect to Evis. With these numbers, we can calculated the fully corrected signal (The fully corrected signal is the $\bar{\nu}_{\mu}$ charged current events in this section). Where Normalized Signal equals to the reconstructed signal times the normalization factor I got from fit ( $p_{\text {or }_{\mu}^{\prime} \bar{\nu}_{\mu} C C}=1.03$, $p_{\text {norm }}^{N C}=2.03, p_{\text {norm }}^{\nu_{\mu} C C}=1.43$ ). Raw signal equals to the Number of data in each bin ( $N_{\text {data }}$ ) minus total normalized background in the same bin (Tot-Bkg). Using the Raw signal in each Evis bin divide by the efficiency in the corresponding bin, we get the fully corrected signal in the 7 Evis bins shown in Table 6.9.

Table 6.7: Reconstructed $\bar{\nu}_{\mu}$ CC events without normalization with the factor from fit.

| Evis(GeV) | $N_{\bar{\nu}_{\mu}}^{C C}$ generated | $N_{\bar{\nu}_{\mu}}^{C C}$ reconstructed | CC-Bkg | NC-bkg | Tot-Bkg | $N_{M C}$ | $N_{\text {data }}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 3970.6 | 1160.3 | 21.1 | 34.2 | 55.3 | 1215.6 | 1260 |
| $8.0-15.0$ | 6179.3 | 2566.7 | 77.7 | 44.0 | 121.7 | 2688.4 | 2513 |
| $15.0-20.0$ | 3933.6 | 1777.5 | 54.3 | 27.5 | 81.8 | 1859.2 | 1733 |
| $20.0-30.0$ | 6098.3 | 2905.4 | 95.3 | 39.3 | 134.6 | 3040.0 | 2997 |
| $30.0-50.0$ | 7618.3 | 3782.4 | 129.7 | 50.1 | 179.8 | 3962.2 | 4067 |
| $50.0-100.0$ | 6608.1 | 3288.7 | 151.1 | 42.2 | 193.3 | 3481.9 | 4217 |
| $100.0-300.0$ | 1491.9 | 677.0 | 56.6 | 13.1 | 69.7 | 756.7 | 1273 |
| $2.5-300.0$ | 35900.0 | 16157.8 | 585.7 | 250.4 | 856.1 | 16994.0 | 18060 |

Table 6.8: Raw signal and fully corrected data after normalization.

| Evis(GeV) | $N_{\bar{\nu}_{\mu}}^{C C}$ Reconstructed | Normalized Signal | Raw Signal | Tot-Bkg | $N_{M C}$ | $N_{\text {data }}$ | $N_{\text {data }} / N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1160.3 | 1199.2 | 1160.4 | 99.6 | 1298.8 | 1260 | 0.970 |
| $8.0-15.0$ | 2566.7 | 2652.6 | 2312.7 | 200.3 | 2852.9 | 2513 | 0.881 |
| $15.0-20.0$ | 1777.5 | 1837.0 | 1599.6 | 133.4 | 1970.4 | 1733 | 0.880 |
| $20.0-30.0$ | 2905.4 | 3002.7 | 2781.0 | 216.0 | 3218.7 | 2997 | 0.931 |
| $30.0-50.0$ | 3782.4 | 3853.2 | 3780.0 | 287.0 | 4140.2 | 4067 | 0.982 |
| $50.0-100.0$ | 3288.7 | 3398.8 | 3915.4 | 301.6 | 3700.4 | 4217 | 1.140 |
| $100.0-300.0$ | 677.0 | 699.7 | 1165.5 | 107.5 | 807.2 | 1273 | 1.577 |
| $2.5-300.0$ | 16157.8 | 16698.8 | 16714.7 | 1345.3 | 18044.1 | 18060 | 1.001 |

Table 6.9: Fully corrected signal got from efficiency vector (shown in Table 6.6).

| Evis(GeV) | Raw Signal | Efficiency | Full Corrected Signal |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 1160.4 | 0.292 | 3970.9 |
| $8.0-15.0$ | 2312.7 | 0.415 | 5567.8 |
| $15.0-20.0$ | 1599.6 | 0.452 | 3539.9 |
| $20.0-30.0$ | 2781.0 | 0.476 | 5793.8 |
| $30.0-50.0$ | 3780.0 | 0.496 | 7613.5 |
| $50.0-100.0$ | 3915.4 | 0.498 | 7830.8 |
| $100.0-300.0$ | 1165.5 | 0.454 | 2568.4 |
| $2.5-300.0$ | 16714.7 | 0.450 | 37137.3 |

Where, in Table 6.9, the Fully Corrected Signal equals to the Raw Signal divide by Efficiency ( Ratio of Reconstructed Signal and Generated Signal). Aside from the efficiency vector, the fully corrected signal in 7 Evis bins are also calculated from the efficiency matrix, which is shown in Table 6.13 , which are going to be used as denominator of ratio between the coherent $\pi^{-}$and $\bar{\nu}_{\mu}$ charged current events.

Table 6.10 shows the number of $\bar{\nu}_{\mu}$ charged current events as a function of $E_{\nu}$ and Evis before multiple the normalization factor got from the fit. Using the elements in this table times the normalization factors and divide by the generated $\bar{\nu}_{\mu}$ charged current events in Table 6.6, we can get the efficiency matrix which are shown in Table 6.11.

After get the efficiency matrix, to calculate the Raw signal, the background after normalization is still needed to be considered. The background matrix after normalization is shown in Table 6.12.

Table 6.11 shows the efficiency matrix of $\bar{\nu}_{\mu}$ charged current event selection in neutrino beam mode analysis. Table 6.12 shows the background matrix in each Evis and $E_{\nu}$ bin. The Raw Signal in each Evis bin can be calculated from the number of NOMAD data minus the number of background after normalization in each bin. The fully corrected signal can be calculated using the number of Raw Signal divide by the efficiency in the same Evis bin, which are shown in Table 6.13.

Table 6.10: $\bar{\nu}_{\mu}$ CC events as a function of $E_{\nu}$ (Evis) before normalization with the factor got from fit.

| $E_{\nu} \backslash$ Evis | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ | $2.5-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 827.4 | 9.4 | 0.1 | 0 | 0 | 0 | 0 | 837.0 |
| $8.0-15.0$ | 331.0 | 2028.6 | 22.6 | 0.1 | 0 | 0 | 0 | 2382.3 |
| $15.0-20.0$ | 1.4 | 505.2 | 1221.3 | 30.0 | 0.1 | 0.1 | 0 | 1758.1 |
| $20.0-30.0$ | 0.4 | 21.8 | 527.5 | 2354.7 | 44.8 | 0.4 | 0 | 2949.6 |
| $30.0-50.0$ | 0.1 | 1.6 | 5.7 | 517.3 | 3340.8 | 61.2 | 0 | 3926.8 |
| $50.0-100.0$ | 0 | 0.1 | 0.2 | 3.1 | 396.4 | 3099.6 | 41.3 | 3540.6 |
| $100.0-300.0$ | 0 | 0 | 0 | 0 | 0.3 | 127.4 | 635.6 | 763.4 |
| $2.5-300.0$ | 1160.3 | 2566.7 | 1777.5 | 2905.4 | 3782.4 | 3288.7 | 677.0 | 16157.8 |

Table 6.11: $\bar{\nu}_{\mu}$ CC efficiency matrix.

| $E_{\nu} \backslash$ Evis | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ | $2.5-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0.21 | 0 | 0 | 0 | 0 | 0 | 0 | 0.02 |
| $8.0-15.0$ | 0.08 | 0.33 | 0.01 | 0 | 0 | 0 | 0 | 0.07 |
| $15.0-20.0$ | 0 | 0.08 | 0.31 | 0 | 0 | 0 | 0 | 0.05 |
| $20.0-30.0$ | 0 | 0 | 0.13 | 0.39 | 0.01 | 0 | 0 | 0.08 |
| $30.0-50.0$ | 0 | 0 | 0 | 0.08 | 0.44 | 0.01 | 0 | 0.11 |
| $50.0-100.0$ | 0 | 0 | 0 | 0 | 0.05 | 0.47 | 0.03 | 0.10 |
| $100.0-300.0$ | 0 | 0 | 0 | 0 | 0 | 0.02 | 0.43 | 0.02 |
| $2.5-300.0$ | 0.29 | 0.42 | 0.45 | 0.48 | 0.50 | 0.50 | 0.45 | 0.45 |

Table 6.12: $\bar{\nu}_{\mu}$ Background matrix(the elements in this table is the total background in each bin).

| $E_{\nu} \backslash$ Evis | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ | $2.5-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 13.2 | 0 | 0 | 0 | 0 | 0 | 0 | 13.2 |
| $8.0-15.0$ | 38.0 | 81.4 | 2.0 | 0 | 0 | 0 | 0 | 121.4 |
| $15.0-20.0$ | 0 | 54.3 | 42.6 | 0 | 0 | 0 | 0 | 96.9 |
| $20.0-30.0$ | 0 | 0 | 53.4 | 98.5 | 8.4 | 0 | 0 | 160.3 |
| $30.0-50.0$ | 0 | 0 | 0 | 75.6 | 142.8 | 16.2 | 0 | 234.6 |
| $50.0-100.0$ | 0 | 0 | 0 | 0 | 103.3 | 170.1 | 10.5 | 283.9 |
| $100.0-300.0$ | 0 | 0 | 0 | 0 | 0 | 113.1 | 96.1 | 209.2 |
| $2.5-300.0$ | 51.2 | 135.7 | 98.0 | 174.1 | 254.5 | 299.4 | 106.6 | 1119.5 |

Table 6.13: Fully corrected signal got from efficiency matrix.

| Evis(GeV) | Raw signal | Efficiency | Fully Corrected Signal |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 1208.8 | 0.29 | 4168.3 |
| $8.0-15.0$ | 2377.3 | 0.42 | 5660.2 |
| $15.0-20.0$ | 1635.0 | 0.45 | 3633.3 |
| $20.0-30.0$ | 2822.9 | 0.48 | 5881.0 |
| $30.0-50.0$ | 3812.5 | 0.50 | 7625.0 |
| $50.0-100.0$ | 3917.6 | 0.50 | 7835.2 |
| $100.0-300.0$ | 1166.4 | 0.45 | 2592.0 |
| $2.5-300.0$ | 16940.5 | 0.45 | 37645.6 |

## Artificial Neural Network Method

To separate signal from background, besides the cuts on kinematic variables, neural network is trained using Monte Carlo events for future separation. Artificial neural networks (ANN or just NN) are widely used in physics data analysis [21]. Most of the neural networks are multilayer perceptions: input layer, at least one hidden layer and output layer. All the layers are made up of interconnected neurons.

1. Input layer. The neurons in a input layer receive the inputs, normalize them and forward them to the first hidden layer [21].
2. Hidden layer. The input to each neuron in a hidden layer is a linear combination of the outputs of the previous layer. The output is a sigmoid function of that combination. A sigmoid function is defined as

$$
S(x)=\frac{1}{1+e^{-x}}
$$

3. Outputlayer. Each neuron in any subsequent layer is computed as a linear combination of the outputs of the previous layer.

In physics analysis with neural network, at first, we need to take multiple kinematic variables $\left(I_{1}, I_{2} \ldots\right)$ as input and use a single output to indicate signal or background. $\left(f=f\left(I_{1}, I_{2} \ldots\right)\right)$ The neural network should be trained and validated with independent Monte Carlo samples before being used in analysis, to make sure that there
is no over training. The neural network is trained using Monte Carlo events(signal and background) with output equals to 1 for signal and 0 for the background. The aim of using neutral network is to minimize the total error of weighted examples, which is defined as the sum of in quadrature, divided by two, of the error on each individual output neuron. The output of the neural network is the probability of signal against background.


Figure 6.9: An example of the structure of an artificial neutral network [21].

We use TMultiLayerPerceptron class in root to build neural networks. The structure of NN is showed in figure 6.9.

## Kinematic Analysis

When the neural network is being used to separate the Coherent $\pi^{-}$signal from background, the first thing we need to decide is which values will be inputs. In this analysis, the following three variables were used as input:

- $p_{T}^{m}$ : missing transverse momentum, which is the measurement of the neutrino beam divergence.
- $X_{b j}$ : the fraction of the nucleon's momentum carried by the struck quark.
- $\zeta_{\pi}: E_{\pi} \times\left(1-\cos \left(\theta_{\pi}\right)\right)$, which means the forwardness of the outgoing Pions.

The principle of choosing input variables includes good separation power and good agreement between Monte Carlo events and data. Using the neural network, there is only one output variable in the coherent $\pi^{-}$analysis, which could be used to separate the signal from background.


Figure 6.10: $\quad P_{T}^{m}$ distribution of the neutrino beam mode data (positive focusing data: FocP).

The distributions of $P_{T}^{m}, \mathrm{Xbj}$, and $\zeta$ are shown from Figure 6.10 to Figure 6.12. In these figures, the total Monte Carlo events (Tot MC) is the sum of the signal (Coherent $\pi^{-}$) and total background (Tot Bkg).


Figure 6.11: Xbj distribution of the neutrino beam mode data (positive focusing data: FocP).


Figure 6.12: $\zeta$ distribution of the neutrino beam mode data (positive focusing data: FocP).


Figure 6.13: The NN distribution comparison of background and signal.


Figure 6.14: The NN distribution comparison of Data and MC.

Figure 6.13 shows the output of the neural network including the distributions of signal (Coherent $\pi^{-}$), $\bar{\nu}_{\mu}$ CC background, $\nu_{\mu}$ CC background, NC background, Coherent $\rho^{-}$, total background (TotBkg: sum of CC background, NC background and background from coherent $\rho^{-}$) as a function of NN output value, and so on. NN-output is a value from 0 to 1 , it is obvious that Background dominates the region NN output less than 0.3, and Signal dominates the region NN output greater than 0.7 .

With the output of neural network, the events in background and signal region can be normalized. Let's use BN to represent the Background Normalization factor and SN represent the Signal Normalization factor. They are calculated as following steps:

At first, all the events are divided into 7 bins according to visible energy (Evis) from 2.5 to 300 GeV (Evis $=\{2.5,8,15,20,30,50,100,300\})$.

Then fit the Monte Carlo events to NOMAD data bin by bin. The background normalization factor (BN), signal normalization factor (SN), and the number of corrected signal $\left(N_{\text {corr-sig }}\right)$ in each bin are calculated with the following formulas (BNtemp is a temporary factor used to calculate SN ):

$$
\begin{aligned}
& \text { BNtemp }=\frac{N_{d a t a}^{b[i]}}{N_{b k g}^{b i l}} \\
& \mathrm{SN}=\frac{\left(N_{\text {data }}^{s[i a}-b n t e m p \times N_{b k g}^{s[i]}\right)}{N_{s i g}^{s[i]}} ; \\
& \mathrm{BN}=\frac{\left(N_{\text {data }}^{b[i]}-s n t e m p \times N_{s i g}^{b[i]}\right)}{N_{b a c}^{b i]}} ; \\
& N_{\text {corr-sig }}=N_{\text {data }}^{s[i]}-N_{b k g}^{s[i]} \times B N
\end{aligned}
$$

Where i is the index of Evis bin number. $N_{\text {data }}^{s[i]}$ is the number of NOMAD data events in signal region. $N_{\text {data }}^{b[i]}$ is the number of NOMAD data events in background region. $N_{b k g}^{s}[i]$ is the number of background events in signal region. $N_{s i g}^{b}[i]$ is the number of signal events in background region. $N_{b k g}^{b}[i]$ is the number of background events in background region. $N_{\text {sig }}^{s}[i]$ is the number of signal events in signal region.

## Coherent $\pi^{-}$in Neutrino Mode

After we got the number of the $\bar{\nu}_{\mu}$ charged current events in 7 visible energy (Evis) bins, in this section, we will measure the number of coherent $\pi^{-}$in each Evis bin. Both of them are going to be used to calculate the ratio between coherent $\pi^{-}$and $\bar{\nu}_{\mu}$ charged current interactions.

Table 6.14: Normalization of the MC events.

| $\bar{\nu}_{\mu}$ CC | 35900 |
| :--- | :--- |
| $\bar{\nu}_{\mu}$ CC DIS | 34028.5 |
| $\bar{\nu}_{\mu}$ NC DIS | 341.05 |
| $\bar{\nu}_{\mu}$ QE | 802.1 |
| $\bar{\nu}_{\mu}$ RES | 1069.5 |
| $\bar{\nu}$ cohPi | 417.1 |
| $\bar{\nu}$ cohRho | 250.3 |
| $\nu_{\mu}$ CC | 1436001 |
| $\nu_{\mu}$ QE | 32084.0 |
| $\nu_{\mu}$ RES | 42778.6 |
| $\nu_{\mu}$ CC DIS | 1361138 |
| $\nu_{\mu}$ NC DIS | 531320.7 |

Table 6.14 shows the numbers of normalized Monte Carlo events with beamweight (flux reweight) and Z-weight. These numbers are calculated from:

The total number of $\nu_{\mu}$ charged current events within Fiducial Volume is normalized to 1436000 [52]; The total number of $\bar{\nu}_{\mu}$ charged current events equals to 0.025 $\times 1436000[8]$;

The ratio between neutral current and charged current events $N^{N C} / N^{C C}$ equals to 0.37 , ratio between non-deep inelastic and deep inelastic events $N^{n o n-D i s} / N^{\text {Dis }}$ equals to 0.055 ; ratio between quasi-elastic and resonance events $N^{Q E} / N^{R E S}$ equals to 0.75 ; ratio between coherent $\pi$ and quasi-elastic events $N^{C o h-\pi} / N^{Q E}$ equals to $0.26 ; N^{C o h-\rho} / N^{C o h-\pi}$ equals to 0.6 ;

Table 6.15: Summary of event reduction in different data and MC samples for the preselection cuts.

|  | $N_{\bar{V}}^{C C}$ | $N_{\nu}^{C C}$ | $N^{N C}$ | Coh $\rho^{-}$ | TotBkg | Coh $\pi^{-}$ | $N_{\text {data }}$ | Ratio |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Total | 47790.3 | 1645469 | 628612.1 | 289.6 | 2322161 | 735.8 | - | - |
| Pfermi | 47380.6 | 1631201 | 623116.4 | 289.6 | 2301988 | 735.8 | - | - |
| W2S | 40610.1 | 1555351 | 589725.7 | 289.6 | 2185977 | 735.8 | - | - |
| Gen-FV | 36076.4 | 1436828 | 544853.6 | 257.6 | 2018015 | 638.3 | - | - |
| $E_{\pi}>0.5($ Coh only $)$ | 36076.4 | 1436828 | 544853.6 | 250.3 | 2018009 | 417.1 | - | - |
| $2.5<$ Evis $<300$ | 35900.0 | 1436001 | 544603.7 | 250.3 | 2016755 | 417.1 | - | - |
| total | 38919.5 | 1434877 | 464599.6 | 247.4 | 1938643 | 590.2 | 4018980 | 2.07 |
| pfermi | 38595.7 | 1422826 | 460710.6 | 247.4 | 1922380 | 590.2 | 4018980 | 2.08 |
| W2 | 33404.4 | 1357502 | 460710.6 | 247.4 | 1850530 | 590.2 | 4018980 | 2.17 |
| Rec-Fidu | 29777.2 | 1255851 | 427703.2 | 219.6 | 1712318 | 528.1 | 3135328 | 1.83 |
| $E_{\pi}>0.5($ Coh only $)$ | 29777.2 | 1255851 | 426469.6 | 201.7 | 1712300 | 400.1 | 3135328 | 1.83 |
| Ph2mu | 27445.9 | 1111808 | 15116.2 | 175.5 | 1154546 | 357.6 | 1240972 | 1.07 |
| Tube $/$ Veto | 27282.6 | 1107721 | 15038.4 | 174.5 | 1150216 | 355.3 | 1185892 | 1.03 |
| Ncand $=2$ | 4218.0 | 133866.6 | 567.5 | 141.1 | 138793.2 | 309.2 | 145171 | 1.04 |
| $\mu^{+} / E_{\mu}$ | 4212.1 | 253.9 | 305.4 | 139.7 | 4911.1 | 307.9 | 10163 | 1.95 |
| Had- | 2568.2 | 150.4 | 230.0 | 138.5 | 3087.0 | 305.8 | 6607 | 1.95 |
| $\theta_{\mu^{+}, \pi^{-}}<177.5$ | 2566.1 | 150.4 | 229.3 | 138.5 | 3084.2 | 305.6 | 4612 | 1.36 |
| Fit Matrix error | 2548.0 | 140.1 | 224.4 | 137.2 | 3049.6 | 303.5 | 4192 | 1.25 |
| $P_{\pi}>1.0$ | 967.8 | 111.6 | 137.5 | 129.2 | 1346.0 | 249.8 | 2180 | 1.37 |
| $\theta_{\mu^{+}, \pi^{-}<0.5}^{P_{T}^{m}<0.5}$ | 689.8 | 85.0 | 132.3 | 128.6 | 1035.7 | 239.1 | 1667 | 1.31 |
| $\left(p_{\pi}-p_{\text {neu }}\right) /\left(p_{\pi}+p_{\text {neu }}\right)>0$ | 296.5 | 30.1 | 12.2 | 116.1 | 454.9 | 231.4 | 754 | 1.10 |
| nV0 cut | 230.0 | 15.5 | 8.1 | 65.8 | 319.4 | 230.9 | 568 | 1.03 |
| Nclu cut | 222.8 | 14.6 | 7.9 | 56.6 | 301.9 | 230.7 | 556 | 1.04 |
| mgg<0.5 | 190.7 | 11.2 | 6.4 | 25.1 | 233.5 | 225.0 | 470 | 1.03 |
| $19512<$ Run $<21270$ | 190.7 | 11.2 | 6.4 | 25.1 | 233.5 | 225.0 | 470 | 1.03 |
|  | 190.7 | 11.2 | 6.4 | 25.1 | 233.5 | 225.0 | 415 | 0.91 |

Table 6.15 shows the summary of event reduction in different data and Monte Carlo samples for the preselection cuts. From this table, we can see that even after all the variable cuts, there is still about $9 \%$ disagreement between the Monte Carlo events and NOMAD data.

To get better matches between Monte Carlo events and NOMAD data, neural network training is used in this analysis which has been introduced in previous sections.


Figure 6.15: The distribution of sensitivity of the neural network in coherent $\pi^{-}$ analysis of neutrino beam mode.

Figure 6.15 shows the distributions of the sensitivity which is defined as

$$
\begin{equation*}
\text { Sensitivity }=\frac{N^{\text {Corr-Sig }}}{\sqrt{N^{\text {Corr-Sig }}+N^{\text {Norm-Bkg }}}} \tag{6.7}
\end{equation*}
$$

where $N^{\text {Corr-Sig }}$ is the number of corrected signal events, $N^{\text {Norm-Bkg }}$ is the number of normalized background events. This figure is plotted from the result in Table 6.16. The sensitivity gave us a tool to find the optimal cut values to define the signal and
background regions. Combine the distribution of the sensitivity and Figure 6.13, we finally chose 0.7 as the cut value of signal region and 0.4 as the cut value of background region, then the signal region is from 0.7 to 1 and the background region is from 0 to 0.4 . The background region is going to be used to calibrate the background of NOMAD data. The region from 0.4 to 0.7 is not ideal region either for background normalization or measurement of signal. Because the ratio of signal to background is too close to 1 .

Some kinematic variable distributions in background (control) region and signal region after neural network are shown from Figure 6.16 to Figure 6.35. It is obvious that there is a good agreement between Monte Carlo events and NOMAD data after the neural network analysis. In these figures, " Dt " represents NOMAD data; "Total MC" represents total Monte Carlo events; "Tot Bkg" represents total background events; " $\bar{\nu} \mathrm{CC}$ " represents $\bar{\nu}_{\mu}$ charged current events; " $\nu \mathrm{CC}$ " represents $\nu_{\mu}$ charged current events; "NC" represents the combination of $\bar{\nu}_{\mu}$ and $\nu_{\mu}$ neutral current events; "Coh $\pi^{-"}$ represents the coherent $\pi^{-}$events, which is also the signal in this analysis, "Coh $\rho^{-"}$ represents the coherent $\rho^{-}$events. All the variable distributions are consistent with the theoretical prediction. For example, the signal (coherent $\pi^{-}$events) is with a lower value of $Q^{2}$ compared to the background. The Bjorken variable Xbj is defined as

$$
\begin{equation*}
X b j=\frac{Q^{2}}{2 M \nu} \tag{6.8}
\end{equation*}
$$

when the target is at rest. Xbj is also very small compared to the background. The missing transverse momentum $P_{T}^{m}$ of signal is smaller compared to the background. $\zeta_{\pi}$ which represent the forwardness of the outgoing hadron is also a smaller value compared to the background, which means the meson Pion is very forward outgoing. The magnitudes of the square of the 4 momentum transfer to the nucleus $|\mathrm{t}|$ of signal is also very small compared to the background. The angle between the leading lepton $\mu^{+}$and the outgoing meson $\pi^{-}$of signal is smaller compared to the background. All
other variable distributions also gave a good agreement between the Monte Carlo events and NOMAD data and are also consistent with the theoretical predictions. In next step, the neural network analysis will be used to separate the signal from background.

Table 6.16 gives out the result of total number of background (Tot-Bkg), background normalization factor (BN), normalized background, number of Data, Raw signal (Raw Sig: which equals to the number of data minus the normalized background), Efficiency (Eff: which equals to the raw signal divide by the number of Data), and corrected signal (Corr-Sig: which equals to raw signal divide by efficiency) with different NN cut values. The corrected signal error includes the statistical error and the error comes from the background normalization.

Figure 6.36 shows a coherent $\pi^{-}$sample picture after all the kinematic cuts and neural network output cut of NOMAD events. In this figure, there are only two tracks are observed. The longer one is identified as Muon $\left(\mu^{+}\right)$which is also called leading lepton. The shorter one is identified as Pion $\left(\pi^{-}\right)$. Some values of this event are listed in Table 6.17. These values are consistent with the theoretical prediction described before, such as the small 4-momentum transfer $Q^{2}$, the small missing transverse momentum $P_{T}^{m}$, the small value of forwardness of meson $\zeta_{\pi}$, the small value of angle between Muon and Pion.


Figure 6.16: The $Y_{b j}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.17: The $Y_{b j}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.18: The $X_{b j}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.19: The $X_{b j}$ distribution from different contributions, $\nu$ - $\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}-$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.20: The $\zeta_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.21: The $\zeta_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.22: The $Q^{2}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.23: The $Q^{2}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.24: The $P_{T}^{m}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.25: The $P_{T}^{m}$ distribution from different contributions, $\nu$ - $\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.26: The $t$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.27: The $t$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.28: The $t^{\prime}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.29: The $t^{\prime}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.30: The $E_{\pi}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.31: The $E_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.32: The $\Phi_{h a d}^{P T}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.33: The $\Phi_{h a d}^{P T}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.34: The angle $\theta$ between Muon and Pion distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$-Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC(histogram).


Figure 6.35: The angle $\theta$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, $\mathrm{NC}, \bar{\nu}$-Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC(histogram).

Table 6.16: The NN cut table (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

| NN | Tot-Bkg | BN | Norm-Bkg | Data | Raw Sig. | Eff. | Corr-Sig. (Err = Stat., BN) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.4 | 70.6 | 0.95 | 67.1 | 238 | 171 | 0.48 | $356 \pm 36.72 \pm 17.12$ |
| 0.41 | 68.9 | 0.95 | 65.5 | 235 | 169 | 0.478 | $354.9 \pm 36.63 \pm 16.81$ |
| 0.42 | 67.7 | 0.95 | 64.3 | 233 | 169 | 0.475 | $355.1 \pm 36.62 \pm 16.59$ |
| 0.43 | 65.2 | 0.948 | 61.8 | 233 | 171 | 0.473 | $362.2 \pm 36.66 \pm 16.07$ |
| 0.44 | 64.1 | 0.948 | 60.7 | 231 | 170 | 0.47 | $362 \pm 36.65 \pm 15.87$ |
| 0.45 | 62 | 0.949 | 58.9 | 226 | 167 | 0.467 | $357.5 \pm 36.42 \pm 15.45$ |
| 0.46 | 60.6 | 0.949 | 57.5 | 224 | 166 | 0.465 | $358.2 \pm 36.41 \pm 15.19$ |
| 0.47 | 58.8 | 0.948 | 55.7 | 223 | 167 | 0.462 | $362.4 \pm 36.46 \pm 14.82$ |
| 0.48 | 57.1 | 0.948 | 54.1 | 220 | 166 | 0.459 | $361.4 \pm 36.37 \pm 14.49$ |
| 0.49 | 56 | 0.949 | 53.1 | 217 | 164 | 0.456 | $359.4 \pm 36.32 \pm 14.3$ |
| 0.5 | 55 | 0.948 | 52.1 | 216 | 164 | 0.453 | $361.9 \pm 36.43 \pm 14.14$ |
| 0.51 | 53.8 | 0.946 | 50.9 | 216 | 165 | 0.45 | $367.1 \pm 36.59 \pm 13.92$ |
| 0.52 | 52.3 | 0.944 | 49.4 | 215 | 166 | 0.446 | $370.9 \pm 36.7 \pm 13.66$ |
| 0.53 | 51.6 | 0.945 | 48.8 | 212 | 163 | 0.443 | $368.5 \pm 36.72 \pm 13.57$ |
| 0.54 | 50 | 0.943 | 47.2 | 211 | 164 | 0.439 | $373.3 \pm 36.88 \pm 13.28$ |
| 0.55 | 48.3 | 0.942 | 45.5 | 209 | 163 | 0.435 | $376 \pm 36.94 \pm 12.94$ |
| 0.56 | 47.1 | 0.941 | 44.3 | 208 | 164 | 0.43 | $380.3 \pm 37.14 \pm 12.73$ |
| 0.57 | 45.6 | 0.94 | 42.9 | 206 | 163 | 0.426 | $382.5 \pm 37.23 \pm 12.47$ |
| 0.58 | 44.6 | 0.94 | 41.9 | 203 | 161 | 0.422 | $381.9 \pm 37.32 \pm 12.3$ |
| 0.59 | 42.9 | 0.94 | 40.3 | 200 | 160 | 0.418 | $382.4 \pm 37.34 \pm 11.97$ |
| 0.6 | 41.9 | 0.942 | 39.5 | 195 | 156 | 0.413 | $376.6 \pm 37.26 \pm 11.81$ |
| 0.61 | 40.5 | 0.942 | 38.1 | 192 | 154 | 0.408 | $377.1 \pm 37.36 \pm 11.56$ |
| 0.62 | 39.6 | 0.942 | 37.3 | 189 | 152 | 0.403 | $376.5 \pm 37.49 \pm 11.44$ |
| 0.63 | 38 | 0.944 | 35.8 | 184 | 148 | 0.398 | $372.3 \pm 37.4 \pm 11.11$ |
| 0.64 | 36.5 | 0.945 | 34.5 | 179 | 144 | 0.392 | $368.6 \pm 37.41 \pm 10.86$ |
| 0.65 | 35 | 0.946 | 33.1 | 175 | 142 | 0.386 | $367.8 \pm 37.49 \pm 10.56$ |
| 0.66 | 33.9 | 0.945 | 32 | 172 | 140 | 0.38 | $368.2 \pm 37.67 \pm 10.37$ |
| 0.67 | 32.3 | 0.943 | 30.4 | 170 | 140 | 0.373 | $373.9 \pm 38.02 \pm 10.07$ |
| 0.68 | 30.5 | 0.943 | 28.8 | 166 | 137 | 0.367 | $374.3 \pm 38.14 \pm 9.686$ |
| 0.69 | 29.4 | 0.944 | 27.7 | 161 | 133 | 0.36 | $370.6 \pm 38.26 \pm 9.509$ |
| 0.7 | 27.7 | 0.947 | 26.2 | 154 | 128 | 0.351 | $363.8 \pm 38.26 \pm 9.183$ |
| 0.71 | 26 | 0.944 | 24.6 | 152 | 127 | 0.343 | $371.4 \pm 38.77 \pm 8.836$ |
| 0.72 | 25.1 | 0.945 | 23.7 | 147 | 123 | 0.335 | $368.3 \pm 39.06 \pm 8.735$ |
| 0.73 | 23.8 | 0.947 | 22.5 | 141 | 118 | 0.326 | $362.8 \pm 39.18 \pm 8.49$ |
| 0.74 | 22.6 | 0.952 | 21.5 | 133 | 111 | 0.318 | $351.1 \pm 39.14 \pm 8.292$ |
| 0.75 | 20.8 | 0.95 | 19.7 | 129 | 109 | 0.308 | $354.7 \pm 39.58 \pm 7.852$ |
| 0.76 | 19.3 | 0.949 | 18.3 | 125 | 107 | 0.299 | $357.1 \pm 40.04 \pm 7.524$ |
| 0.77 | 17.9 | 0.951 | 17 | 119 | 102 | 0.289 | $352.9 \pm 40.31 \pm 7.206$ |
| 0.78 | 16.6 | 0.955 | 15.9 | 111 | 95.1 | 0.279 | $341 \pm 40.34 \pm 6.937$ |
| 0.79 | 15.6 | 0.96 | 15 | 103 | 88 | 0.268 | $328.4 \pm 40.47 \pm 6.771$ |
| 0.8 | 14.5 | 0.956 | 13.9 | 101 | 87.1 | 0.256 | $340.1 \pm 41.8 \pm 6.614$ |
| 0.81 | 13.4 | 0.959 | 12.9 | 94 | 81.1 | 0.245 | $331.3 \pm 42.17 \pm 6.389$ |
| 0.82 | 12.7 | 0.956 | 12.2 | 91 | 78.8 | 0.233 | $338.5 \pm 43.54 \pm 6.365$ |
| 0.83 | 11.7 | 0.949 | 11.2 | 90 | 78.8 | 0.22 | $357.8 \pm 45.55 \pm 6.209$ |
| 0.84 | 11.2 | 0.949 | 10.6 | 85 | 74.4 | 0.208 | $358.2 \pm 46.97 \pm 6.264$ |
| 0.85 | 9.86 | 0.949 | 9.36 | 79 | 69.6 | 0.194 | $358.1 \pm 48.23 \pm 5.908$ |
| 0.86 | 8.48 | 0.95 | 8.05 | 72 | 63.9 | 0.18 | $355.9 \pm 49.7 \pm 5.496$ |
| 0.87 | 7.34 | 0.949 | 6.97 | 65 | 58 | 0.163 | $357.1 \pm 52.08 \pm 5.258$ |
| 0.88 | 6.09 | 0.943 | 5.74 | 60 | 54.3 | 0.145 | $375.1 \pm 55.92 \pm 4.905$ |
| 0.89 | 4.93 | 0.94 | 4.64 | 52 | 47.4 | 0.124 | $381.3 \pm 60.45 \pm 4.625$ |



Figure 6.36: Coherent $\pi^{-}$event picture originated by $\bar{\nu}_{\mu}$ contamination in the neutrino mode.

Table 6.17: Kinematic information of a coherent $\pi^{-}$event corresponding to Figure 6.36 survived from preselection and Neural Network.

| Anti-neutrino Beam Mode |  |  |  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Run | Event | NN(lh) | XVR | YVR | ZVR |  |  |
| 19796 | 11602 | 0.606 | 24.3 | -37.8 | 291 |  |  |
| $P_{T}^{m}$ | $\Phi_{L H}$ | $\Phi_{m P t H}$ | $\Theta_{\mu \pi}$ | t | $\zeta$ | Xbj | Ybj |
| 0.0919 | 177 | 20.6 | 0.308 | 0.0256 | 0.0857 | 0.125 | 0.0799 |
| Ncand | Nprim | Nsecond | Nvzero | nhitMuon | nhithad |  |  |
| 2 | 2 | 0 | 0 | 14 | 13 |  |  |
| particle | E | $\|\mathrm{P}\|$ | Px | Py | Pz | $\Theta$ | $\Phi$ |
| Muon | 25.3 | 25.3 | -0.271 | -0.633 | 25.3 | 1.56 | 66.9 |
| Pion | 2.19 | 2.19 | 0.21 | 0.568 | 2.1 | 16.1 | 69.7 |

## Determination of the Coherent $\pi^{-}$Cross-section

Instead of calculating the absolute cross-section of coherent $\pi^{-}$interactions in NOMAD, the R defined in Equation 6.9 is calculated as the ratio between the crosssection of coherent $\pi^{-}$events and inclusive $\bar{\nu}_{\mu}$ charged current cross-section.

$$
\begin{equation*}
R=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)} . \tag{6.9}
\end{equation*}
$$

We calculated R instead of the absolute cross-sections for two reasons: one is normalization; the other one is, some systematic uncertainties can be removed because they share the same systematic uncertainties. The numbers of inclusive $\bar{\nu}_{\mu}$ charged current events in each bin have been shown in Table 6.13. Next step, we are going to calculate the coherent $\pi^{-}$in each bin.

Table 6.18: BN and SN table in 7 visible energy (Evis) bins, calculated from variable BN depends on the $\operatorname{Evis}\left(\mathrm{R}=\frac{\sigma\left(C_{o h \pi^{-}}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis(GeV) | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0.7865 | 0.6167 | 0.2622 | $14.520 \pm 5.224$ |
| $8.0-15.0$ | 0.835 | 0.8941 | 0.1996 | $16.490 \pm 3.333$ |
| $15.0-20.0$ | 1.213 | 0.5591 | 0.3541 | $7.851 \pm 3.217$ |
| $20.0-30.0$ | 0.8609 | 1.179 | 0.2712 | $12.760 \pm 2.608$ |
| $30.0-50.0$ | 0.8089 | 0.9373 | 0.2715 | $7.362 \pm 1.795$ |
| $50.0-100.0$ | 1.658 | 0.9211 | 0.627 | $4.094 \pm 1.668$ |
| $100.0-300.0$ | 0.4954 | 1.943 | 1.251 | $3.894 \pm 2.341$ |
| $2.5-300.0$ | 0.9444 | 0.8886 | 0.1165 | $9.845 \pm 1.047$ |

Table 6.18 and Table 6.19 show the result of R (ratio between coherent $\pi^{-}$and $\bar{\nu}_{\mu}$ charged current events.) in 7 visible energy (Evis) bins which are calculated from variable BN and fixed BN .

Figure 6.37 shows the distribution of BN as a function of visible energy (Evis) in 7 bins, which are corresponding to the result in Table 6.18.

Table 6.19: BN and SN table in 7 visible energy (Evis) bins, using a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis(GeV) | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0.9444 | 0.6167 | 0.1165 | $13.770 \pm 5.257$ |
| $8.0-15.0$ | 0.9444 | 0.8941 | 0.1165 | $16.080 \pm 3.327$ |
| $15.0-20.0$ | 0.9444 | 0.5591 | 0.1165 | $8.492 \pm 2.966$ |
| $20.0-30.0$ | 0.9444 | 1.179 | 0.1165 | $12.600 \pm 2.588$ |
| $30.0-50.0$ | 0.9444 | 0.9373 | 0.1165 | $7.138 \pm 1.780$ |
| $50.0-100.0$ | 0.9444 | 0.9211 | 0.1165 | $4.748 \pm 1.406$ |
| $100.0-300.0$ | 0.9444 | 1.943 | 0.1165 | $3.615 \pm 2.294$ |
| $2.5-300.0$ | 0.9444 | 0.8886 | 0.1165 | $9.845 \pm 1.047$ |



Figure 6.37: The distribution of BN as a function of visible energy (Evis) in 7 bins(the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

Table 6.20: Signal in signal region, and Generated signal information calculated from variable BN.

| Evis(GeV) | Sig-S | Sig-Gen | Efficiency |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 19.262 | 98.158 | 0.197 |
| $8.0-15.0$ | 39.467 | 104.391 | 0.378 |
| $15.0-20.0$ | 20.909 | 51.016 | 0.410 |
| $20.0-30.0$ | 27.114 | 63.615 | 0.426 |
| $30.0-50.0$ | 26.177 | 59.890 | 0.437 |
| $50.0-100.0$ | 15.043 | 34.827 | 0.432 |
| $100.0-300.0$ | 1.908 | 5.195 | 0.367 |
| $2.5-300.0$ | 150 | 417.1 | 0.360 |

Table 6.21: Norm-bkg, Corr-sig as a function of Evis in 7 bins calculated from variable BN.

| Evis | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 3.89 | 0.787 | 3.06 | 15 | 11.9 | 0.197 | $60.53 \pm 21.15 \pm 5.17$ |
| $8.0-15.0$ | 8.04 | 0.835 | 6.71 | 42 | 35.3 | 0.378 | $93.34 \pm 18.38 \pm 4.243$ |
| $15.0-20.0$ | 3.55 | 1.21 | 4.31 | 16 | 11.7 | 0.41 | $28.52 \pm 11.28 \pm 3.068$ |
| $20.0-30.0$ | 4.67 | 0.861 | 4.02 | 36 | 32 | 0.426 | $75.03 \pm 15.04 \pm 2.972$ |
| $30.0-50.0$ | 5.52 | 0.809 | 4.46 | 29 | 24.5 | 0.437 | $56.13 \pm 13.25 \pm 3.428$ |
| $50.0-100.0$ | 3.1 | 1.66 | 5.14 | 19 | 13.9 | 0.432 | $32.08 \pm 12.27 \pm 4.502$ |
| $100.0-300.0$ | 0.591 | 0.495 | 0.293 | 4 | 3.71 | 0.367 | $10.09 \pm 5.724 \pm 2.015$ |
| $2.5-300.0$ | 29.4 | 0.944 | 27.7 | 161 | 133 | 0.36 | $370.6 \pm 38.26 \pm 9.509$ |

Table 6.20 shows the Efficiency as a function of visible energy (Evis) in 7 bins, which are calculated from ratio of the signal in signal region (Sig-S) after all the kinematic variable cuts and generated signal (Sig-Gen).

Table 6.21 shows the result of total background (Tot-bkg), BN, normalized background (Norm-bkg), data, raw signal (Raw-sig), Efficiency (Eff), corrected signal (Corr-Sig) in 7 Evis bins. Norm-bkg equals to Tot-bkg times BN. Raw-sig equals to data minus Norm-bkg. The Corr-Sig equals to the Raw-sig divide by Efficiency. Table 6.22 and Table 6.23 show the result of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ calculated from variable BN of neutrino beam mode data (Positive focusing data: FocP) in 7 Evis and $E_{\nu}$ bins.

Figure 6.38 and Figure 6.39 show the distributions of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R} \times \mathrm{E}$ in both linear scale (top) and log scale (bottom), where R is calculated using a variable BN.

Table 6.22: Corrected signal (Corr-Sig) as a function of Evis in 7 bins calculated from variable $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$-CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 60.53 | 4168.3 | $14.520 \pm 5.224$ |
| $8-15$ | 11.84 | 93.34 | 5660.2 | $16.490 \pm 3.333$ |
| $15-20$ | 17.40 | 28.52 | 3633.3 | $7.851 \pm 3.217$ |
| $20-30$ | 24.60 | 75.03 | 5881.0 | $12.760 \pm 2.608$ |
| $30-50$ | 38.47 | 56.13 | 7625.0 | $7.362 \pm 1.795$ |
| $50-100$ | 71.54 | 32.08 | 7835.2 | $4.094 \pm 1.668$ |
| $100-300$ | 142.70 | 10.09 | 2592.0 | $3.894 \pm 2.341$ |
| $2.5-300$ | 25.00 | 370.6 | 37645.6 | $9.845 \pm 1.047$ |



Figure 6.38: $\mathrm{R}=\frac{\sigma\left(\mathrm{Coh} \mathrm{\pi}^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ distribution in both linear scale (top) and log scale (bottom), calculated using a variable BN which depends on the Evis(the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).


Figure 6.39: $\mathrm{R} \times<\mathrm{E}>$ distribution in both linear scale (top) and log scale (bottom), calculated using a variable BN which depends on the Evis(the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

Table 6.23: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins calculated from variable $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| $E_{\nu}$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | $59.961 \pm 21.310$ | 4168.3 | $14.385 \pm 5.112$ |
| $8-15$ | 11.84 | $92.977 \pm 19.428$ | 5660.2 | $16.426 \pm 3.432$ |
| $15-20$ | 17.40 | $31.166 \pm 11.704$ | 3633.3 | $8.578 \pm 3.221$ |
| $20-30$ | 24.60 | $72.886 \pm 15.214$ | 5881.0 | $12.393 \pm 2.587$ |
| $30-50$ | 38.47 | $56.672 \pm 13.981$ | 7625.0 | $7.432 \pm 1.834$ |
| $50-100$ | 71.54 | $32.738 \pm 13.313$ | 7835.2 | $4.178 \pm 1.699$ |
| $100-300$ | 142.70 | $9.323 \pm 5.539$ | 2592.0 | $3.597 \pm 2.137$ |
| $2.5-300$ | 25.00 | $370.6 \pm 39.42$ | 37645.6 | $9.845 \pm 1.047$ |

Table 6.24: Norm-bkg, Corr-sig as a function of Evis in 7 bins, calculated from a fixed BN.

| Evis | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8$ | 3.89 | 0.944 | 3.67 | 15 | 11.3 | 0.197 | $57.42 \pm 21.79 \pm 2.297$ |
| $8-15$ | 8.04 | 0.944 | 7.59 | 42 | 34.4 | 0.378 | $91.01 \pm 18.67 \pm 2.477$ |
| $15-20$ | 3.55 | 0.944 | 3.35 | 16 | 12.6 | 0.41 | $30.85 \pm 10.73 \pm 1.01$ |
| $20-30$ | 4.67 | 0.944 | 4.41 | 36 | 31.6 | 0.426 | $74.12 \pm 15.17 \pm 1.277$ |
| $30-50$ | 5.52 | 0.944 | 5.21 | 29 | 23.8 | 0.437 | $54.42 \pm 13.5 \pm 1.471$ |
| $50-100$ | 3.1 | 0.944 | 2.93 | 19 | 16.1 | 0.432 | $37.21 \pm 10.99 \pm 0.8365$ |
| $100-300$ | 0.591 | 0.944 | 0.558 | 4 | 3.44 | 0.367 | $9.371 \pm 5.943 \pm 0.1876$ |
| $2.5-300$ | 29.4 | 0.944 | 27.7 | 161 | 133 | 0.36 | $370.6 \pm 38.26 \pm 9.509$ |

Table 6.25 and Table 6.26 show the result of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ calculated from a fixed BN of neutrino beam mode data (Positive focusing data: FocP) in 7 Evis and $E_{\nu}$ bins.

Table 6.25: Corrected signal (Corr-sig) as a function of Evis in 7 bins, calculated from a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis(GeV) | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 57.42 | 4168.3 | $13.770 \pm 5.257$ |
| $8-15$ | 11.84 | 91.01 | 5660.2 | $16.080 \pm 3.327$ |
| $15-20$ | 17.40 | 30.85 | 3633.3 | $8.492 \pm 2.966$ |
| $20-30$ | 24.60 | 74.12 | 5881.0 | $12.600 \pm 2.588$ |
| $30-50$ | 38.47 | 54.42 | 7625.0 | $7.138 \pm 1.780$ |
| $50-100$ | 71.54 | 37.21 | 7835.2 | $4.748 \pm 1.406$ |
| $100-300$ | 142.70 | 9.371 | 2592.0 | $3.615 \pm 2.294$ |
| $2.5-300$ | 25.00 | 370.6 | 37645.6 | $9.845 \pm 1.047$ |

Table 6.26: Corrected signal (Corr-sig) as a function of $E_{\nu}$ in 7 bins calculated from a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| $E_{\nu}$ | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | $56.921 \pm 21.438$ | 4168.3 | $13.656 \pm 5.143$ |
| $8-15$ | 11.84 | $90.749 \pm 19.359$ | 5660.2 | $16.321 \pm 3.420$ |
| $15-20$ | 17.40 | $33.195 \pm 10.867$ | 3633.3 | $9.136 \pm 2.991$ |
| $20-30$ | 24.60 | $72.066 \pm 15.065$ | 5881.0 | $12.254 \pm 2.562$ |
| $30-50$ | 38.47 | $55.295 \pm 13.771$ | 7625.0 | $7.252 \pm 1.806$ |
| $50-100$ | 71.54 | $37.439 \pm 11.369$ | 7835.2 | $4.778 \pm 1.451$ |
| $100-300$ | 142.70 | $8.743 \pm 5.409$ | 2592.0 | $3.373 \pm 2.087$ |
| $2.5-300$ | 25.00 | $370.6 \pm 39.42$ | 37645.6 | $9.845 \pm 1.047$ |

Figure 6.40 and Figure 6.41 show the distributions of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R} \times \mathrm{E}$ distributions in both linear scale (top) and log scale (bottom), where R is calculated from a fixed BN.


Figure 6.40: $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ distribution in both linear scale (top) and log scale (bottom), calculated from a fixed BN (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).


Figure 6.41: $\mathrm{R} \times<\mathrm{E}>$ distribution in both linear scale (top) and log scale (bottom), calculated from a fixed BN (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

### 6.5 Anti-neutrino Beam Mode

Similar to the analysis of the Neutrino Beam Mode NOMAD data, the NOMAD data generated from Anti-neutrino Beam Mode is also analyzed with the same method. With the anti-neutrino beam mode data, the ratio R which is the ratio between the coherent $\pi^{-}$events and $\bar{\nu}_{\mu}$ charged current events is also calculated. Combine the results of R from neutrino and anti-neutrino beam mode together, the averaged value of R is calculated and compared to the result of coherent $\pi^{+}$.

## Normalization of Event Samples

Same as the process in the neutrino beam mode analysis, in the anti-neutrino beam mode analysis, all the Monte Carlo events are also normalized based on coherent events and inclusive charged current events.

Table 6.27: Normalization of (Anti-)Neutrino Beam Mode events [8].

| Mode | Event Number |
| :---: | :---: |
| $\nu_{\mu} \mathrm{CC}$ | 13,000 |
| $\nu_{\mu} \mathrm{NC}$ | $13,000 \times 0.35$ |
| $\bar{\nu}_{\mu} \mathrm{CC}$ | 35,000 |
| $\bar{\nu}_{\mu} \mathrm{NC}$ | $35,000 \times 0.35$ |

Table 6.28: Ratios between interaction mode.

| Ratio | Value |
| :---: | :---: |
| NC/CC | 0.37 |
| non-DIS/DIS | 0.055 |
| QE/RES | 0.75 |

## Preselection of $\bar{\nu}_{\mu}$ CC Events

Table 6.29: Cut table of Monte Carlo events in generated level.

|  | $N_{\bar{U}}^{C C}$ | $N_{\nu}^{C C}$ | $N^{N C}$ | TotBkg | $N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| Total | 45462.8 | 14987.6 | 21038.9 | 36026.5 | 81489.3 |
| Pfermi $<1.0$ | 45073.8 | 14859.9 | 20853.3 | 35713.2 | 80787.0 |
| W2s $>1.96$ (CCDIS) | 39409.3 | 14157.7 | 18721.1 | 32878.8 | 72288.1 |
| $\mid$ GenX, (GenY-5) $\mid<130 \mathrm{~cm}$ | 35012.3 | 13073.7 | 16851.4 | 29925.1 | 64937.4 |
| Zmin $<$ GenZ $<405 \mathrm{~cm}$ |  |  |  |  |  |
| $2.5<$ Evis $<300$ | 35000 | 13000 | 16800 | 29800 | 64800 |

Table 6.29 shows the preselection of $\bar{\nu}_{\mu}$ charged current events in generated level. From this table, we can see that all the $\bar{\nu}_{\mu}$ charged current events has been normalized to 35,000 , and $\nu_{\mu}$ charged current events has been normalized to 13,000 which are consistent with the prediction shown in Table 6.27.

Table 6.30 shows the preselection of $\bar{\nu}_{\mu}$ charged current events in reconstructed level with kinematic variable cuts. Similar to the result in the neutrino beam mode data analysis, there is also a disagreement about $15 \%$ between the normalized Monte Carlo events and anti-neutrino beam mode NOMAD data. To get a better agreement, the Monte Carlo events was refitted. Through this refitting process, the new normalization factors can be calculated and used to normalize the Monte Carlo events.

Table 6.30: Reconstructed variable cut table after normalization and reweighted.

|  | $N_{\bar{\nu}}^{C C}$ | $N_{\nu}^{C C}$ | $N^{N C}$ | TotBkg | $N_{M C}$ | $N_{\text {data }}$ | $N_{\text {data }} / N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Total | 37718.5 | 12846.4 | 13049.0 | 50767.6 | 63613.9 | 189609 | 2.981 |
| Pfermi $<1.0$ | 37404.5 | 12740.5 | 12939.0 | 50343.5 | 63084.1 | 189609 | 3.006 |
| W2s $>1.96$ (CCDIS) | 32812.3 | 12175.0 | 12862 | 45674.3 | 57849.3 | 189609 | 3.278 |
| FV cut | 27076.9 | 10795.2 | 10953.7 | 38030.6 | 48825.8 | 103952 | 2.129 |
| PhaseII | 25286.1 | 9615.8 | 330.9 | 25617.0 | 35232.7 | 43627 | 1.238 |
| Nmu $=1$ | 25199.1 | 9523.1 | 282.3 | 25481.4 | 35004.6 | 43257 | 1.236 |
| veto/tube | 25064.6 | 9495.8 | 281.0 | 25345.6 | 34841.4 | 41900 | 1.203 |
| ncand $>=2$ | 23152.5 | 9287.8 | 278.6 | 23431.1 | 32718.9 | 38213 | 1.168 |
| $\left\|\vec{p}_{\mu}\right\|>2.5$ | 23055.4 | 9235.7 | 267.6 | 23323.0 | 32558.7 | 38009 | 1.167 |
| $\mu^{+}$ | 22969.9 | 101.0 | 143.0 | 244.0 | 23213.9 | 27055 | 1.165 |
| DeltaP $/ \mathrm{P}<=0.2$ | 22926.1 | 86.6 | 138.9 | 225.5 | 23151.7 | 26879 | 1.161 |
| thetamupi $<177.5$ | 22923.4 | 86.6 | 138.9 | 225.5 | 23149.0 | 26545 | 1.147 |
| Evis $<300 \mathrm{GeV}$ | 22922.4 | 86.6 | 138.9 | 225.5 | 23147.9 | 26479 | 1.144 |
| Ehad $<300 \mathrm{GeV}$ | 22922.4 | 86.6 | 138.9 | 225.5 | 23147.9 | 26479 | 1.144 |
| $19512<\mathrm{Run}<21270$ | 22922.4 | 86.6 | 138.9 | 225.5 | 23147.9 | 26479 | 1.144 |

To ensure the Monte Carlo events match the NOMAD data well, the same fitting method as used in analysis of neutrino beam mode is also applied to anti-neutrino mode. During this fitting process, two variables are used: Ybj and $\hat{\nu}$. The normalization factors for $\bar{\nu}_{\mu}$ charged current events, $\nu_{\mu}$ charged current events and neutral current events are $k_{\text {norm }}^{\nu_{\mu}}, k_{\text {norm }}^{\nu_{\mu} C C}$, and $k_{\text {norm }}^{N C}$.

The procedure includes the following steps:

- Create the 2D distribution of Ybj , and $\hat{\nu}$;
- Divide the distribution into equal-populated $5 \times 10=50$ cubes. Because the events do not distribute uniformly, the volumes of the cubes are not the same.
- Run through all the possible normalization,
- Set the default of $k_{\text {norm }}^{\bar{\nu}_{\mu} C C}$ at 1.15 and vary it in steps of 0.001 from 1.1 to 1.2;
- Set $k_{\text {norm }}^{N C}=4.7$ and vary it in steps of 0.001 from 4.6 to 4.8 ;
- Set $k_{\text {norm }}^{\nu_{\mu} C C}=0.1$ and vary it in steps of 0.001 from 0.0 to 0.2 ;

Then the total number of Monte Carlo events after normalization becomes

$$
\begin{aligned}
N_{M C}= & N_{\bar{\nu}_{\mu}}^{C C} \times k_{\text {norm }}^{\bar{\nu}_{\mu} C C}+ \\
& N^{N C} \times k_{\text {norm }}^{N C}+ \\
& N_{\nu_{\mu}}^{C C} \times k_{\text {norm }}^{\nu_{\mu} C C}
\end{aligned}
$$

- Calculate the $\chi^{2}$ for each cube and sum over all of them.

$$
\chi^{2}=\frac{\left(N_{\text {data }}-N_{M C}\right)^{2}}{\sigma_{\text {data }}^{2}+\sigma_{M C}^{2}}
$$

While $\sigma_{\text {data }}=\sqrt{N_{\text {data }}}$, and $\sigma_{M C}=\sqrt{\sum_{j}\left(\sqrt{N_{M C}(j)}\right)^{2}}$, index i represent the Monte Carlo events ( $\bar{\nu}_{\mu} \mathrm{CC}, \nu_{\mu} \mathrm{CC}$ and NC );

- Get the parameter corresponding the minimum $\chi^{2}$.

$$
\begin{array}{ll}
k_{\text {norm }}^{\bar{\nu}_{\mu} C C}=1.114 \pm 0.015 & \\
& k_{\text {norm }}^{N C}=4.67 \\
& k_{\text {norm }}^{\nu_{\mu} C C}=0.0
\end{array}
$$

Since the total points I used for this 2D fit is equal to 50 and the minimum $\chi^{2}$ is 247.2 , then the value $\chi^{2} / \mathrm{DOF}$ equals to 4.944 .

The normalization factors got from this fit procedure will be used to calculated the number of $\bar{\nu}_{\mu}$ charged current in the anti-neutrino beam mode. Same method was used in this anti-neutrino beam mode data analysis as in the neutrino beam mode data analysis.

Table 6.31 shows the number of $\bar{\nu}_{\mu}$-CC events, background from charged current events (CC-Bkg), background from neutral current events (NC-Bkg), total background (Tot-Bkg), total Monte Carlo events (Tot-MC), number of data ( $N_{\text {data }}$ ) after all the cuts in 7 Evis bins without normalization (not multiply Norm factors I got from the fit).

The first two columns of Table 6.33 show the generated and reconstructed $\bar{\nu}_{\mu}$ charged current events in 7 visible energy (Evis) bins. Analogous to the calculations in the neutrino beam mode data analysis, the efficiency vector can also be calculated from the ratio between the elements in these two columns (Reconstructed $\bar{\nu}_{\mu}$ charged current events divide by the Generated $\bar{\nu}_{\mu}$ charged current events) which is shown in the last column of Table 6.33.

Table 6.31: Reconstructed $\bar{\nu}_{\mu}$ charged current events without normalization.

| Evis(GeV) | $N_{\bar{\nu}_{\mu}}^{C C}$ generated | $N_{\bar{\nu}_{\mu}}^{C C}$ constructed | CC-Bkg | NC-bkg | Tot-Bkg | $N_{M C}$ | $N_{\text {data }}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1583.3 | 1168.5 | 0.8 | 6.5 | 7.3 | 1175.8 | 1042 |
| $8.0-15.0$ | 8550.9 | 5598.0 | 3.6 | 18.3 | 22.0 | 5620.0 | 5165 |
| $15.0-20.0$ | 6424.7 | 4016.0 | 3.9 | 14.7 | 18.7 | 4034.7 | 4039 |
| $20.0-30.0$ | 7976.4 | 5038.1 | 10.2 | 25.0 | 35.1 | 5073.2 | 5891 |
| $30.0-50.0$ | 5868.9 | 3988.4 | 21.0 | 35.8 | 56.9 | 4045.3 | 5262 |
| $50.0-100.0$ | 3879.0 | 2652.2 | 29.3 | 31.4 | 60.7 | 2712.8 | 4024 |
| $100.0-300.0$ | 716.8 | 461.3 | 17.7 | 7.2 | 24.9 | 486.1 | 1056 |
| $2.5-300.0$ | 35000 | 22922.4 | 86.6 | 138.9 | 225.5 | 23147.9 | 26479 |

Table 6.32: Raw signal and fully corrected data after normalization.

| Evis(GeV) | $\bar{\nu}_{\mu}$-CC reconstructed | normalized signal | Raw signal | Tot-Bkg | $N_{M C}$ | $N_{\text {data }}$ | $N_{\text {data }} / N_{M C}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1168.5 | 1301.7 | 1011.6 | 30.4 | 1332.1 | 1042 | 0.782 |
| $8.0-15.0$ | 5598.0 | 6236.2 | 5079.5 | 85.5 | 6321.7 | 5165 | 0.817 |
| $15.0-20.0$ | 4016.0 | 4473.8 | 3970.4 | 68.6 | 4542.4 | 4039 | 0.889 |
| $20.0-30.0$ | 5038.1 | 5612.4 | 5774.2 | 116.8 | 5729.2 | 5891 | 1.028 |
| $30.0-50.0$ | 3988.4 | 4443.1 | 5094.8 | 167.2 | 4610.3 | 5262 | 1.141 |
| $50.0-100.0$ | 2652.2 | 2954.5 | 3877.4 | 146.6 | 3101.1 | 4024 | 1.298 |
| $100.0-300.0$ | 461.3 | 513.8 | 1022.4 | 33.6 | 547.4 | 1056 | 1.929 |
| $2.5-300.0$ | 22922.4 | 25535.5 | 25830.3 | 648.7 | 26184.2 | 26479 | 1.011 |

Table 6.33: The efficiency (ratio between Rec- $\bar{\nu}_{\mu} \mathrm{CC}$ and $\mathrm{Gen}-\bar{\nu}_{\mu} \mathrm{CC}$ events) in $7 E_{\nu}$ bins.

| Evis $(\mathrm{GeV})$ | Generated $N_{\bar{\nu}_{\mu}}^{C C}$ | Reconstructed $N_{\bar{\nu}_{\mu}}^{C C}$ | Efficiency |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 1583.3 | 1168.5 | 0.738 |
| $8.0-15.0$ | 8550.9 | 5598.0 | 0.655 |
| $15.0-20.0$ | 6424.7 | 4016.0 | 0.625 |
| $20.0-30.0$ | 7976.4 | 5038.1 | 0.632 |
| $30.0-50.0$ | 5868.9 | 3988.4 | 0.680 |
| $50.0-100.0$ | 3879.0 | 2652.2 | 0.684 |
| $100.0-300.0$ | 716.8 | 461.3 | 0.644 |
| $2.5-300.0$ | 35000 | 22922.4 | 0.655 |

Table 6.32 shows the reconstructed $\bar{\nu}_{\mu}$, normalized signal, Raw signal which equals to the number of data minus the total background, total background (Tot-Bkg), total number of MC events ( $N_{M C}$ ) which equals to the sum of normalized signal and the total background, number of data $\left(N_{\text {data }}\right)$. These numbers will be used to calculated the fully corrected signal.

Table 6.34: Fully corrected signal get from efficiency vector (shown in Table 6.33).

| Evis(GeV) | Raw signal | Efficiency | Full Corrected Signal |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 1011.6 | 0.738 | 1370.7 |
| $8.0-15.0$ | 5079.5 | 0.655 | 7755.0 |
| $15.0-20.0$ | 3970.4 | 0.625 | 6352.6 |
| $20.0-30.0$ | 5774.2 | 0.632 | 9136.4 |
| $30.0-50.0$ | 5094.8 | 0.680 | 7492.4 |
| $50.0-100.0$ | 3877.4 | 0.684 | 5668.7 |
| $100.0-300.0$ | 1022.4 | 0.644 | 1587.6 |
| $2.5-300.0$ | 25830.3 | 0.655 | 39435.6 |

Using the Raw signal in Table 6.32 divide by the efficiency in the same energy bin shown in Table 6.33, we get the fully corrected signal which are shown in Table 6.34.

Table 6.35: $\bar{\nu}_{\mu}$ charged current events as a function of $E_{\nu}$ (Evis) before normalization.

| $E_{\nu} \backslash$ Evis | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ | $2.5-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 572.9 | 15.7 | 0.3 | 0.2 | 0.5 | 0.1 | 0.1 | 589.7 |
| $8.0-15.0$ | 585.0 | 4214.1 | 74.0 | 7.2 | 3.5 | 1.8 | 1.0 | 4886.6 |
| $15.0-20.0$ | 6.9 | 1288.5 | 2779.0 | 109.8 | 6.6 | 3.0 | 1.7 | 4195.5 |
| $20.0-30.0$ | 2.6 | 72.5 | 1143.9 | 4214.7 | 116.3 | 8.4 | 4.9 | 5563.4 |
| $30.0-50.0$ | 0.8 | 5.8 | 16.9 | 695.5 | 3485.1 | 86.7 | 5.8 | 4296.7 |
| $50.0-100.0$ | 0.2 | 1.2 | 1.6 | 10.1 | 374.0 | 2445.2 | 52.1 | 2884.4 |
| $100.0-300.0$ | 0.1 | 0.2 | 0.2 | 0.6 | 2.5 | 106.8 | 395.7 | 506.1 |
| $2.5-300.0$ | 1168.5 | 5598.0 | 4016.0 | 5038.1 | 3988.5 | 2652 | 461.3 | 22922.4 |

Table 6.36: Background as a function of $E_{\nu}$ and Evis before normalization.

| $E_{\nu} \backslash$ Evis | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ | $2.5-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $8.0-15.0$ | 1.9 | 1.1 | 0 | 0 | 0 | 0 | 0 | 3.0 |
| $15.0-20.0$ | 0.9 | 2.5 | 0.4 | 0 | 0 | 0 | 0 | 3.8 |
| $20.0-30.0$ | 1.7 | 4.7 | 3.3 | 2.3 | 0.2 | 0 | 0 | 12.2 |
| $30.0-50.0$ | 1.2 | 5.4 | 5.8 | 9.5 | 7.3 | 0.1 | 0.1 | 29.4 |
| $50.0-100.0$ | 0.7 | 3.9 | 4.1 | 9.9 | 19.0 | 12.4 | 0.3 | 50.2 |
| $100.0-300.0$ | 0.3 | 0.7 | 1.2 | 3.2 | 9.3 | 18.8 | 6.8 | 40.2 |
| $2.5-300.0$ | 6.7 | 18.3 | 14.8 | 24.9 | 35.8 | 31.3 | 7.2 | 139 |

Table 6.37: Raw signal and fully corrected data after normalization.

| Evis $(\mathrm{GeV})$ | $\bar{\nu}_{\mu}$-CC reconstructed | normalized signal | Raw signal | Tot-Bkg | $N_{M C}$ | $N_{\text {data }}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1168.5 | 1301.7 | 1010.7 | 31.3 | 1333.0 | 1042 |
| $8.0-15.0$ | 5598.0 | 6236.2 | 5079.5 | 85.5 | 6231.7 | 5165 |
| $15.0-20.0$ | 4016.0 | 4473.8 | 3969.9 | 69.1 | 4542.9 | 4039 |
| $20.0-30.0$ | 5038.1 | 5612.4 | 5774.7 | 116.3 | 5728.7 | 5891 |
| $30.0-50.0$ | 3988.4 | 4443.1 | 5094.8 | 167.2 | 4610.3 | 5262 |
| $50.0-100.0$ | 2652.2 | 2954.5 | 3877.8 | 146.2 | 3100.7 | 4024 |
| $100.0-300.0$ | 461.3 | 513.8 | 1022.4 | 33.6 | 547.4 | 1056 |
| $2.5-300.0$ | 22922.4 | 25535.5 | 25829.9 | 649.1 | 26184.6 | 26479 |

Table 6.38: Fully corrected signal get from efficiency matrix.

| Evis(GeV) | Raw signal | Efficiency | Full Corrected Signal |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 1010.7 | 0.738 | 1369.5 |
| $8.0-15.0$ | 5079.5 | 0.655 | 7755.0 |
| $15.0-20.0$ | 3969.9 | 0.625 | 6351.8 |
| $20.0-30.0$ | 5774.7 | 0.632 | 9137.2 |
| $30.0-50.0$ | 5094.8 | 0.680 | 7492.4 |
| $50.0-100.0$ | 3877.8 | 0.684 | 5669.3 |
| $100.0-300.0$ | 1022.4 | 0.644 | 1587.6 |
| $2.5-300.0$ | 25829.9 | 0.655 | 39435.0 |

## Coherent $\pi^{-}$in Anti-neutrino Mode

Table 6.39: Normalization of the MC events.

| $\bar{\nu}_{\mu}$ CC | 35000 |
| :--- | :--- |
| $\bar{\nu}_{\mu}$ CC QE | 782.0 |
| $\bar{\nu}_{\mu}$ CC RES | 1042.7 |
| $\bar{\nu}_{\mu}$ CC DIS | 33175.4 |
| $\bar{\nu}_{\mu}$ NC DIS | 12250 |
| $\bar{\nu}_{\mu}$ cohPi | 406.6 |
| $\bar{\nu}_{\mu}$ cohRho | 244.0 |
| $\nu_{\mu}$ CC | 13000 |
| $\nu_{\mu}$ QE generated | 290.5 |
| $\nu_{\mu}$ RES generated | 387.3 |
| $\nu_{\mu}$ CC DIS generated | 12322.3 |
| $\nu_{\mu}$ NC DIS generated | 4550 |

Total $\nu_{\mu}$-CC events with in Fiducial Volume is normalized to 13000 events; $\bar{\nu}_{\mu^{-}}$ CC events equals to 35000 events; NC/CC equals to 0.35 , non-Dis/Dis equals to 0.055 ;QE/RES equals to 0.75 ; Coh $-\pi$ /QE equals to $2 \times 0.26$; Coh- $\rho /$ Coh $-\pi$ equals to 0.6 .

Table 6.40: Summary of event reduction in different data and MC samples for the preselection cuts.

|  | $N_{\bar{\nu}_{\mu}}^{C C}$ | $N_{\nu_{\mu}}^{C C}$ | $N^{N C}$ | Coh $\rho^{-}$ | TotBkg | Coh $\pi^{-}$ | $N_{\text {data }}$ | Ratio |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Total | 45462.8 | 14987.6 | 21038.9 | 282.6 | 81771.9 | 658.4 | - | - |
| Pfermi | 45073.8 | 14859.9 | 20853.3 | 282.6 | 81069.6 | 658.4 | - | - |
| W2S | 39409.3 | 14157.7 | 18721.1 | 282.6 | 72570.7 | 658.4 | - | - |
| Gen-FV | 35012.3 | 13073.7 | 16851.4 | 250.9 | 65188.3 | 572.8 | - | - |
| $E_{\pi}>0.5($ Coh only $)$ | 35012.3 | 13073.7 | 16851.4 | 250.9 | 65188.3 | 407.2 | - | - |
| $2.5<$ Evis $<300$ | 35000.0 | 13000.0 | 16800 | 244.0 | 65044.0 | 406.6 | - | - |
| total | 37718.5 | 12846.4 | 13049.0 | 242.3 | 63856.3 | 543.7 | 189609 | 2.94 |
| pfermi | 37404.5 | 12740.5 | 12939.0 | 242.3 | 63326.3 | 543.7 | 189609 | 2.96 |
| W2 | 32812.3 | 12175.0 | 12862 | 242.3 | 58091.6 | 543.7 | 189609 | 3.24 |
| Rec-Fidu | 29244.2 | 11271.13 | 11651.2 | 214.8 | 52381.2 | 487.5 | 137286 | 2.60 |
| $E_{\pi}>0.5($ Coh only $)$ | 29244.2 | 11271.1 | 11651.1 | 196.4 | 52362.8 | 376.0 | 137286 | 2.60 |
| Ph2mu | 27325.2 | 10009.6 | 341.0 | 169.0 | 37844.8 | 365.3 | 48963 | 1.28 |
| Tube $/$ Veto | 27153.4 | 9975.1 | 339.1 | 168.1 | 37635.6 | 363.2 | 46123 | 1.21 |
| Ncand=2 | 4534.5 | 1035.5 | 17.2 | 136.6 | 5723.8 | 320.2 | 7065 | 1.17 |
| $\mu^{+} / E_{\mu}$ | 4529.3 | 2.8 | 10.0 | 135.1 | 4677.2 | 318.9 | 5519 | 1.10 |
| Had- | 2752.2 | 1.5 | 8.0 | 134.1 | 2895.8 | 316.8 | 3558 | 1.08 |
| $\theta_{\mu^{+}, \pi^{-}}<177.5$ | 2750.2 | 1.5 | 8.0 | 134.1 | 2893.8 | 316.6 | 3293 | 1.03 |
| Fit Matrix error | 2732.3 | 1.3 | 8.0 | 133.0 | 2874.6 | 314.5 | 3230 | 1.01 |
| $P_{\pi}>1.0$ | 1045.7 | 1.0 | 5.2 | 124.2 | 1176.2 | 264.8 | 1542 | 1.07 |
| $\theta_{\mu^{+}, \pi^{-}}<0.5$ | 740.4 | 0.8 | 5.0 | 123.6 | 869.9 | 254.9 | 1252 | 1.13 |
| $P_{T}^{m}<0.5$ | 327.3 | 0.2 | 0.6 | 113.0 | 441.1 | 246.8 | 789 | 1.14 |
| $\left(p_{\pi}-p_{\text {neu }}\right) /\left(p_{\pi}+p_{\text {neu }}\right)>0$ | 257.9 | 0.1 | 0.3 | 64.6 | 323.0 | 246.3 | 651 | 1.14 |
| nV0 cut | 250.0 | 0.1 | 0.3 | 56.0 | 306.5 | 246.1 | 630 | 1.14 |
| Nclu cut | 216.0 | 0.1 | 0.2 | 25.3 | 241.5 | 240.0 | 531 | 1.10 |
| mgg<0.5 | 216.0 | 0.1 | 0.2 | 25.3 | 241.5 | 240.0 | 531 | 1.10 |
| $19512<$ Run $<21270$ | 216.0 | 0.1 | 0.2 | 25.3 | 241.5 | 240.0 | 531 | 1.10 |

## Kinematic Analysis



Figure 6.42: $\quad P_{T}^{m}$ distribution of the anti-neutrino beam mode data (negative focusing data: FocN).

The distributions of missing transverse momentum $P_{T}^{m}$, the Bjorken variable Xbj , the outgoing meson forwardness $\zeta$ are shown from Figure 6.42 to Figure 6.44. These three variables were used during the neural network analysis.

Figure 6.47 shows the distribution of sensitivity (defined in Equation (6.7)) in anti-neutrino beam mode. From this figure, we could see that, in the anti-neutrino beam mode analysis, the best cut value for the signal is 0.8 , however, to be consistent with the event selection in the neutrino beam mode, we also chose 0.4 and 0.7 for background and signal region.

Similar to the analysis of neutrino beam mode, in this anti-neutrino beam mode analysis, some kinematic variable distributions after neural network are shown from Figure 6.48 to Figure 6.67. It is obvious that there is also a good agreement between Monte Carlo events and NOMAD data after the neural network analysis. Same to the


Figure 6.43: Xbj distribution of the anti-neutrino beam mode data (negative focusing data: FocN).


Figure 6.44: $\zeta$ distribution of the anti-neutrino beam mode data (negative focusing data: FocN).


Figure 6.45: The NN distribution comparison of background and signal.


Figure 6.46: The NN distribution comparison of Data and MC.


Figure 6.47: The distribution of sensitivity of the neural network in coherent $\pi^{-}$ analysis of anti-neutrino beam mode.
plots in neutrino beam mode, in these figures, "Dt" represents NOMAD data; "Total MC" represents total Monte Carlo events; "Tot Bkg" represents total background events; " $\bar{\nu} \mathrm{CC}$ " represents $\bar{\nu}_{\mu}$ charged current events; " $\nu \mathrm{CC}$ " represents $\nu_{\mu}$ charged current events; "NC" represents the combination of $\bar{\nu}_{\mu}$ and $\nu_{\mu}$ neutral current events; "Coh $\pi^{-"}$ represents the coherent $\pi^{-}$events; which is also the signal in this analysis; "Coh $\rho^{-"}$ represents the coherent $\rho^{-}$events.


Figure 6.48: The $Y_{b j}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.49: The $Y_{b j}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.50: The $X_{b j}$ distribution from different contributions, $\nu$ - $\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.51: The $X_{b j}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}$, $\bar{\nu}-\operatorname{Coh} \rho^{-}$and $\operatorname{Coh} \pi^{-}$in signal ( $>0.70$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.52: The $\zeta_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.53: The $\zeta_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in signal $(>0.7)$ region and the Comparison between Data points with error bars) and MC (histogram).


Figure 6.54: The $Q^{2}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.55: The $Q^{2}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.56: The $P_{T}^{m}$ distribution from different contributions, $\nu$ - $\mathrm{CC}, \bar{\nu}$ - $\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.57: The $P_{T}^{m}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}, \bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.58: The $t$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.59: The $t$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.60: The $t^{\prime}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ $\operatorname{Coh} \rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.61: The $t^{\prime}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.62: The $E_{\pi}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$ - CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.63: The $E_{\pi}$ distribution from different contributions, $\nu-\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}, \bar{\nu}-$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.64: The $\Phi_{h a d}^{P T}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.65: The $\Phi_{\text {had }}^{P T}$ distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$ Coh $\rho^{-}$and Coh $\pi^{-}$in signal ( $>0.7$ ) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.66: The angle $\theta$ between muon and pion distribution from different contributions, $\nu$-CC, $\bar{\nu}$-CC, NC, $\bar{\nu}$-Coh $\rho^{-}$and Coh $\pi^{-}$in background (control) region and the Comparison between Data (points with error bars) and MC (histogram).


Figure 6.67: The angle $\theta$ distribution from different contributions, $\nu$ - $\mathrm{CC}, \bar{\nu}-\mathrm{CC}, \mathrm{NC}$, $\bar{\nu}$ - $\operatorname{Coh} \rho^{-}$and $\operatorname{Coh} \pi^{-}$in signal ( $>0.7$ ) region and he Comparison between Data points with error bars and MC (histogram).

Table 6.41: The NN cut table (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu^{-}}$CC and $\nu_{\mu}$-CC events).

| NN | Tot-Bkg | BN | Norm-Bkg | Data | Raw Sig. | Eff. | Corr-Sig. (Err = Stat., BN) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.4 | 76.6 | 1.19 | 91.5 | 311 | 219 | 0.533 | $411.5 \pm 39.61 \pm 19.29$ |
| 0.41 | 74.4 | 1.2 | 89 | 306 | 217 | 0.531 | $408.7 \pm 39.37 \pm 18.82$ |
| 0.42 | 72.9 | 1.2 | 87.2 | 304 | 217 | 0.529 | $410 \pm 39.32 \pm 18.51$ |
| 0.43 | 71.2 | 1.2 | 85.1 | 301 | 216 | 0.527 | $409.9 \pm 39.19 \pm 18.14$ |
| 0.44 | 69.1 | 1.19 | 82.4 | 300 | 218 | 0.524 | $415.1 \pm 39.2 \pm 17.69$ |
| 0.45 | 68 | 1.19 | 81.1 | 300 | 219 | 0.522 | $419.6 \pm 39.29 \pm 17.5$ |
| 0.46 | 66.2 | 1.19 | 78.9 | 297 | 218 | 0.519 | $420.3 \pm 39.22 \pm 17.14$ |
| 0.47 | 64.9 | 1.19 | 77.4 | 294 | 217 | 0.516 | $419.9 \pm 39.18 \pm 16.9$ |
| 0.48 | 63.8 | 1.19 | 76.1 | 291 | 215 | 0.513 | $418.9 \pm 39.14 \pm 16.7$ |
| 0.49 | 63 | 1.19 | 75 | 290 | 215 | 0.51 | $421.8 \pm 39.26 \pm 16.59$ |
| 0.5 | 60.3 | 1.19 | 71.7 | 289 | 217 | 0.506 | $429.5 \pm 39.29 \pm 16$ |
| 0.51 | 59.2 | 1.19 | 70.3 | 286 | 216 | 0.503 | $429.2 \pm 39.3 \pm 15.81$ |
| 0.52 | 58.2 | 1.19 | 69.1 | 285 | 216 | 0.498 | $433.5 \pm 39.49 \pm 15.69$ |
| 0.53 | 56.6 | 1.19 | 67.2 | 281 | 214 | 0.494 | $433.2 \pm 39.48 \pm 15.4$ |
| 0.54 | 54.8 | 1.19 | 65.1 | 275 | 210 | 0.488 | $430 \pm 39.39 \pm 15.06$ |
| 0.55 | 52.7 | 1.19 | 62.9 | 265 | 202 | 0.483 | $418.8 \pm 39.05 \pm 14.67$ |
| 0.56 | 50.6 | 1.2 | 60.6 | 254 | 193 | 0.476 | $406 \pm 38.69 \pm 14.27$ |
| 0.57 | 47.9 | 1.2 | 57.4 | 245 | 188 | 0.469 | $399.7 \pm 38.44 \pm 13.71$ |
| 0.58 | 44.3 | 1.19 | 52.9 | 244 | 191 | 0.461 | $415 \pm 38.76 \pm 12.92$ |
| 0.59 | 42.1 | 1.19 | 50.3 | 238 | 188 | 0.453 | $414.5 \pm 38.77 \pm 12.48$ |
| 0.6 | 40.1 | 1.2 | 47.9 | 230 | 182 | 0.445 | $408.8 \pm 38.65 \pm 12.08$ |
| 0.61 | 38.4 | 1.2 | 46 | 224 | 178 | 0.439 | $405.7 \pm 38.62 \pm 11.75$ |
| 0.62 | 36.6 | 1.2 | 43.9 | 217 | 173 | 0.432 | $400.5 \pm 38.51 \pm 11.38$ |
| 0.63 | 35.2 | 1.2 | 42.2 | 211 | 169 | 0.425 | $397 \pm 38.52 \pm 11.12$ |
| 0.64 | 33.7 | 1.2 | 40.5 | 203 | 162 | 0.419 | $387.6 \pm 38.27 \pm 10.78$ |
| 0.65 | 32.5 | 1.21 | 39.2 | 196 | 157 | 0.412 | $380.3 \pm 38.21 \pm 10.58$ |
| 0.66 | 31 | 1.21 | 37.4 | 191 | 154 | 0.406 | $378.7 \pm 38.24 \pm 10.27$ |
| 0.67 | 29.5 | 1.2 | 35.5 | 189 | 154 | 0.399 | $385.1 \pm 38.54 \pm 9.929$ |
| 0.68 | 28.6 | 1.2 | 34.4 | 186 | 152 | 0.392 | $386.5 \pm 38.76 \pm 9.771$ |
| 0.69 | 27.2 | 1.2 | 32.8 | 183 | 150 | 0.385 | $389.7 \pm 39.01 \pm 9.492$ |
| 0.7 | 26.3 | 1.21 | 31.7 | 175 | 143 | 0.379 | $378.4 \pm 38.86 \pm 9.335$ |
| 0.71 | 25.2 | 1.21 | 30.4 | 172 | 142 | 0.372 | $380.9 \pm 39.13 \pm 9.112$ |
| 0.72 | 24.2 | 1.2 | 29.2 | 169 | 140 | 0.365 | $382.7 \pm 39.36 \pm 8.896$ |
| 0.73 | 22.9 | 1.2 | 27.5 | 167 | 139 | 0.358 | $389.1 \pm 39.71 \pm 8.574$ |
| 0.74 | 21.3 | 1.2 | 25.7 | 161 | 135 | 0.351 | $385.3 \pm 39.68 \pm 8.158$ |
| 0.75 | 20.2 | 1.2 | 24.3 | 157 | 133 | 0.344 | $385.7 \pm 39.87 \pm 7.873$ |
| 0.76 | 18.9 | 1.2 | 22.8 | 153 | 130 | 0.336 | $387.3 \pm 40.14 \pm 7.56$ |
| 0.77 | 18 | 1.2 | 21.6 | 150 | 128 | 0.328 | $391.7 \pm 40.66 \pm 7.373$ |
| 0.78 | 16.6 | 1.2 | 19.9 | 146 | 126 | 0.319 | $394.6 \pm 40.98 \pm 6.979$ |
| 0.79 | 15.7 | 1.2 | 18.8 | 143 | 124 | 0.311 | $399.8 \pm 41.59 \pm 6.793$ |
| 0.8 | 14.8 | 1.2 | 17.8 | 139 | 121 | 0.301 | $402.3 \pm 42.16 \pm 6.611$ |
| 0.81 | 13.4 | 1.2 | 16.1 | 131 | 115 | 0.292 | $393.7 \pm 42.16 \pm 6.188$ |
| 0.82 | 12.8 | 1.2 | 15.4 | 126 | 111 | 0.282 | $392.2 \pm 42.73 \pm 6.099$ |
| 0.83 | 11.6 | 1.2 | 14 | 122 | 108 | 0.271 | $399 \pm 43.6 \pm 5.773$ |
| 0.84 | 10.9 | 1.2 | 13 | 117 | 104 | 0.259 | $401.9 \pm 44.61 \pm 5.653$ |
| 0.85 | 10.3 | 1.2 | 12.3 | 110 | 97.7 | 0.247 | $395.4 \pm 45.28 \pm 5.577$ |
| 0.86 | 9.44 | 1.2 | 11.3 | 104 | 92.7 | 0.234 | $395.3 \pm 46.32 \pm 5.409$ |
| 0.87 | 8.21 | 1.2 | 9.83 | 98 | 88.2 | 0.22 | $401.1 \pm 47.71 \pm 5.012$ |
| 0.88 | 7.34 | 1.2 | 8.83 | 88 | 79.2 | 0.203 | $389.5 \pm 48.9 \pm 4.853$ |
| 0.89 | 6.8 | 1.2 | 8.16 | 82 | 73.8 | 0.186 | $396.5 \pm 51.47 \pm 4.904$ |

Figure 6.68 shows a coherent $\pi^{-}$event picture after all the kinematic cut and neural network output cut of NOMAD events.


Figure 6.68: Coherent $\pi^{-}$event picture originated by $\bar{\nu}_{\mu}$ contamination in the anti-neutrino mode.

Table 6.42: Kinematic information of a coherent $\pi^{-}$event corresponding to Figure 6.68 survived from preselection and Neural Network.

| Neutrino Beam Mode |  |  |  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Run | Event | NN(lh) | XVR | YVR | ZVR |  |  |
| 20896 | 22826 | 0.621 | -32 | 107 | 223 |  |  |
| $P_{T}^{m}$ | $\Phi_{L H}$ | $\Phi_{m P t H}$ | $\Theta_{\mu \pi}$ | t | $\zeta$ | Xbj | Ybj |
| 0.204 | 167 | 71.9 | 0.124 | 0.0332 | 0.0363 | 0.0602 | 0.25 |
| Ncand | Nprim | Nsecond | Nvzero | nhitmuon | nhithad |  |  |
| 2 | 2 | 0 | 0 | 47 | 15 |  |  |
| particle | E | $\|\mathrm{P}\|$ | Px | Py | Pz | $\Theta$ | $\Phi$ |
| Muon | 27.3 | 27.3 | 0.397 | -0.784 | 27.3 | 1.84 | -63.2 |
| Pion | 8.59 | 8.59 | -0.201 | 0.763 | 8.56 | 5.27 | -75.3 |

## Determination of the Coherent $\pi^{-}$Cross-section

Table 6.43: BN and SN table in 7 Bins, calculated from variable BN depends on the $\operatorname{Evis}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis $(\mathrm{GeV})$ | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0.4263 | 0.37 | 0.2307 | $9.442 \pm 6.266$ |
| $8.0-15.0$ | 1.056 | 0.7615 | 0.2013 | $13.680 \pm 2.762$ |
| $15.0-20.0$ | 1.289 | 0.8975 | 0.3096 | $11.550 \pm 2.674$ |
| $20.0-30.0$ | 1.373 | 0.926 | 0.3261 | $8.325 \pm 1.939$ |
| $30.0-50.0$ | 1.333 | 1.337 | 0.4512 | $8.190 \pm 2.017$ |
| $50.0-100.0$ | 2.131 | 1.742 | 1.102 | $6.144 \pm 2.400$ |
| $100.0-300.0$ | 3.164 | 2.606 | 4.47 | $4.100 \pm 4.915$ |
| $2.5-8.0$ | 1.202 | 0.9585 | 0.1343 | $9.883 \pm 1.018$ |

Table 6.44: BN , and $\delta_{B N}$ table, using a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis $(\mathrm{GeV})$ | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1.202 | 0.37 | 0.1343 | $5.186 \pm 7.655$ |
| $8.0-15.0$ | 1.202 | 0.7615 | 0.1343 | $13.240 \pm 2.787$ |
| $15.0-20.0$ | 1.202 | 0.8975 | 0.1343 | $11.710 \pm 2.596$ |
| $20.0-30.0$ | 1.202 | 0.926 | 0.1343 | $8.584 \pm 1.841$ |
| $30.0-50.0$ | 1.202 | 1.337 | 0.1343 | $8.343 \pm 1.925$ |
| $50.0-100.0$ | 1.202 | 1.742 | 0.1343 | $6.903 \pm 2.016$ |
| $100.0-300.0$ | 1.202 | 2.606 | 0.1343 | $5.160 \pm 3.576$ |
| $2.5-8.0$ | 1.202 | 0.9585 | 0.1343 | $9.883 \pm 1.018$ |

Table 6.45: Signal in signal region, and Generated signal information calculated from variable BN.

| Evis $(\mathrm{GeV})$ | Sig-S | Sig-Gen | Efficiency |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | 8.664 | 34.980 | 0.248 |
| $8.0-15.0$ | 50.945 | 139.361 | 0.366 |
| $15.0-20.0$ | 32.407 | 81.768 | 0.396 |
| $20.0-30.0$ | 34.583 | 82.141 | 0.421 |
| $30.0-50.0$ | 20.115 | 45.861 | 0.438 |
| $50.0-100.0$ | 8.947 | 19.996 | 0.447 |
| $100.0-300.0$ | 1.083 | 2.498 | 0.434 |
| $2.5-300.0$ | 156.7 | 406.6 | 0.385 |



Figure 6.69: The distribution of BN as a function of visible energy(Evis) in 7 bins (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

Table 6.46: Norm-bkg, Corr-sig as a function of Evis in 7 bins calculated from variable BN.

| Evis | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1.86 | 0.426 | 0.794 | 4 | 3.21 | 0.248 | $12.94 \pm 8.411 \pm 1.735$ |
| $8.0-15.0$ | 8.71 | 1.06 | 9.21 | 48 | 38.8 | 0.366 | $106.1 \pm 20.87 \pm 4.798$ |
| $15.0-20.0$ | 4.59 | 1.29 | 5.91 | 35 | 29.1 | 0.396 | $73.39 \pm 16.61 \pm 3.586$ |
| $20.0-30.0$ | 5.81 | 1.37 | 7.98 | 40 | 32 | 0.421 | $76.06 \pm 17.13 \pm 4.499$ |
| $30.0-50.0$ | 3.83 | 1.33 | 5.1 | 32 | 26.9 | 0.438 | $61.36 \pm 14.59 \pm 3.941$ |
| $50.0-100.0$ | 2.07 | 2.13 | 4.42 | 20 | 15.6 | 0.447 | $34.83 \pm 12.61 \pm 5.103$ |
| $100.0-300.0$ | 0.372 | 3.16 | 1.18 | 4 | 2.82 | 0.434 | $6.509 \pm 6.795 \pm 3.835$ |
| $2.5-300.0$ | 27.2 | 1.2 | 32.8 | 183 | 150 | 0.385 | $389.7 \pm 39.01 \pm 9.492$ |

Table 6.47: Corrected signal (Corr-Sig) as a function of Evis in 7 bins calculated from variable $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis $(\mathrm{GeV})$ | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}($ Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 12.94 | 1370.7 | $9.442 \pm 6.266$ |
| $8-15$ | 11.84 | 106.1 | 7755.0 | $13.680 \pm 2.762$ |
| $15-20$ | 17.40 | 73.39 | 6352.6 | $11.550 \pm 2.674$ |
| $20-30$ | 24.60 | 76.06 | 9136.4 | $8.325 \pm 1.939$ |
| $30-50$ | 38.47 | 61.36 | 7492.4 | $8.190 \pm 2.017$ |
| $50-100$ | 71.54 | 34.83 | 5668.7 | $6.144 \pm 2.400$ |
| $100-300$ | 142.70 | 6.509 | 1587.6 | $4.100 \pm 4.915$ |
| $2.5-300$ | 25.00 | 389.7 | 39435.6 | $9.883 \pm 1.018$ |

Table 6.48: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins calculated from variable $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| $E_{\nu}$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | $13.162 \pm 8.164$ | 1370.7 | $9.602 \pm 5.956$ |
| $8-15$ | 11.84 | $106.167 \pm 21.991$ | 7755.0 | $13.690 \pm 2.836$ |
| $15-20$ | 17.40 | $74.131 \pm 17.080$ | 6352.6 | $11.669 \pm 2.689$ |
| $20-30$ | 24.60 | $76.719 \pm 17.907$ | 9136.4 | $8.397 \pm 1.960$ |
| $30-50$ | 38.47 | $60.443 \pm 15.147$ | 7492.4 | $8.067 \pm 2.022$ |
| $50-100$ | 71.54 | $34.634 \pm 13.077$ | 5668.7 | $6.110 \pm 2.307$ |
| $100-300$ | 142.70 | $5.943 \pm 6.872$ | 1587.6 | $3.743 \pm 4.329$ |
| $2.5-300$ | 25.00 | $389.7 \pm 40.15$ | 39435.6 | $9.883 \pm 1.018$ |

Figure 6.70 and Figure 6.71 show the distributions of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R} \times \mathrm{E}$ distributions in both linear scale and $\log$ scale, where R is calculated from variable BN.

Table 6.49 and Table 6.50 show the results calculated using a fixed BN of antineutrino beam mode data (Negative focusing data: FocN).


Figure 6.70: $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ distribution in both linear scale (top) and log scale (bottom), calculated from variable BN which depends on the Evis (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu^{-}}$-CC events).


Figure 6.71: $\mathrm{R} \times<\mathrm{E}>$ distribution in both linear scale (top) and log scale (bottom), calculated from variable BN which depends on the Evis (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

Table 6.49: Norm-bkg, Corr-sig as a function of Evis in 7 bins, calculated from a fixed BN.

| Evis | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8$ | 1.86 | 1.2 | 2.24 | 4 | 1.76 | 0.248 | $7.108 \pm 10.44 \pm 1.01$ |
| $8-15$ | 8.71 | 1.2 | 10.5 | 48 | 37.5 | 0.366 | $102.6 \pm 21.37 \pm 3.202$ |
| $15-20$ | 4.59 | 1.2 | 5.52 | 35 | 29.5 | 0.396 | $74.39 \pm 16.42 \pm 1.555$ |
| $20-30$ | 5.81 | 1.2 | 6.98 | 40 | 33 | 0.421 | $78.42 \pm 16.72 \pm 1.853$ |
| $30-50$ | 3.83 | 1.2 | 4.6 | 32 | 27.4 | 0.438 | $62.51 \pm 14.38 \pm 1.173$ |
| $50-100$ | 2.07 | 1.2 | 2.49 | 20 | 17.5 | 0.447 | $39.13 \pm 11.41 \pm 0.6221$ |
| $100-300$ | 0.372 | 1.2 | 0.447 | 4 | 3.55 | 0.434 | $8.192 \pm 5.676 \pm 0.1152$ |
| $2.5-300$ | 27.2 | 1.2 | 32.8 | 183 | 150 | 0.385 | $389.7 \pm 39.01 \pm 9.492$ |

Table 6.50: Corrected signal (Corr-sig) as a function of Evis in 7 bins, calculated from a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| Evis(GeV) | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 7.108 | 1370.7 | $5.186 \pm 7.655$ |
| $8-15$ | 11.84 | 102.6 | 7755.0 | $13.240 \pm 2.787$ |
| $15-20$ | 17.40 | 74.39 | 6352.6 | $11.710 \pm 2.596$ |
| $20-30$ | 24.60 | 78.42 | 9136.4 | $8.584 \pm 1.841$ |
| $30-50$ | 38.47 | 62.51 | 7492.4 | $8.343 \pm 1.925$ |
| $50-100$ | 71.54 | 39.13 | 5668.7 | $6.903 \pm 2.016$ |
| $100-300$ | 142.70 | 8.192 | 1587.6 | $5.160 \pm 3.576$ |
| $2.5-300$ | 25.00 | 389.7 | 39435.6 | $9.883 \pm 1.018$ |

Table 6.51: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins, calculated from a fixed $\mathrm{BN}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$.

| $E_{\nu}$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | $7.745 \pm 9.921$ | 1370.7 | $5.650 \pm 7.238$ |
| $8-15$ | 11.84 | $102.432 \pm 22.304$ | 7755.0 | $13.209 \pm 2.876$ |
| $15-20$ | 17.40 | $75.069 \pm 16.589$ | 6352.6 | $11.817 \pm 2.611$ |
| $20-30$ | 24.60 | $79.013 \pm 17.020$ | 9136.4 | $8.648 \pm 1.863$ |
| $30-50$ | 38.47 | $61.795 \pm 14.368$ | 7492.4 | $8.248 \pm 1.918$ |
| $50-100$ | 71.54 | $38.911 \pm 11.736$ | 5668.7 | $6.864 \pm 2.070$ |
| $100-300$ | 142.70 | $7.436 \pm 5.009$ | 1587.6 | $4.684 \pm 3.155$ |
| $2.5-300$ | 25.00 | $389.7 \pm 40.15$ | 39435.6 | $9.883 \pm 1.018$ |

Figure 6.72 and Figure 6.73 show the distributions of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R} \times \mathrm{E}$ distributions while $R$ is calculated from a fixed BN .

### 6.6 Comparison of Neutrino Mode and Anti-neutrino Mode

With the result of R in Neutrino Beam Mode and Anti-neutrino Beam Mode, we could calculate the average value of $\langle\mathrm{R}\rangle$. We use $\left\langle R^{-}\right\rangle$to denote this average value, "-" means this is the R of coherent $\pi^{-}$. To explain how to calculate $<R^{-}>$, let's take a simple example: for the variables $A \pm \delta A$ and $B \pm \delta B$, the average value of C and $\delta \mathrm{C}$ are calculated as:


Figure 6.72: $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ distribution in both linear scale (top) and log scale (bottom), calculated from a fixed BN (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu^{-}}$CC events).


Figure 6.73: $\mathrm{R} \times<\mathrm{E}>$ distribution in both linear scale (top) and log scale (bottom), calculated from a fixed BN (the beam(flux) reweight is applied to all the $\bar{\nu}_{\mu}$-CC events).

$$
\begin{array}{r}
C=\frac{\frac{A}{\delta A^{2}}+\frac{B}{\delta B^{2}}}{\frac{1}{\delta A^{2}}+\frac{1}{\delta B^{2}}} \\
\delta C=\sqrt{\frac{1}{\frac{1}{\delta A^{2}}+\frac{1}{\delta B^{2}}}} \tag{6.11}
\end{array}
$$

Using Equation (6.10) and Equation (6.11), the averaged value of $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$in neutrino beam mode (positive focusing data: FocP) and anti-neutrino beam mode (negative focusing data: FocN) are calculated and shown in Table 6.52.

Table 6.52: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of Evis using variable BN.)

| Evis(GeV $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocN $)$ | $<\mathrm{R}>$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $14.520 \pm 5.224$ | $9.442 \pm 6.266$ | $12.438 \pm 4.012$ |
| $8.0-15.0$ | $16.490 \pm 3.333$ | $13.68 \pm 2.762$ | $14.824 \pm 2.127$ |
| $15.0-20.0$ | $7.851 \pm 3.217$ | $11.55 \pm 2.674$ | $10.039 \pm 2.056$ |
| $20.0-30.0$ | $12.760 \pm 2.608$ | $8.325 \pm 1.939$ | $9.904 \pm 1.556$ |
| $30.0-50.0$ | $7.362 \pm 1.795$ | $8.19 \pm 2.017$ | $7.728 \pm 1.341$ |
| $50.0-100.0$ | $4.094 \pm 1.668$ | $6.144 \pm 2.400$ | $4.762 \pm 1.370$ |
| $100.0-300.0$ | $3.894 \pm 2.341$ | $4.100 \pm 4.915$ | $3.932 \pm 2.114$ |
| $2.5-300.0$ | $9.845 \pm 1.047$ | $9.883 \pm 1.018$ | $9.865 \pm 0.730$ |

Table 6.53: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of $E_{\nu}$ using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocN $)$ | $<\mathrm{R}\rangle$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $14.385 \pm 5.112$ | $9.602 \pm 5.956$ | $12.356 \pm 3.879$ |
| $8.0-15.0$ | $16.426 \pm 3.432$ | $13.690 \pm 2.836$ | $14.800 \pm 2.186$ |
| $15.0-20.0$ | $8.578 \pm 3.221$ | $11.669 \pm 2.689$ | $10.400 \pm 2.064$ |
| $20.0-30.0$ | $12.393 \pm 2.587$ | $8.397 \pm 1.960$ | $9.854 \pm 1.562$ |
| $30.0-50.0$ | $7.432 \pm 1.834$ | $8.067 \pm 2.022$ | $7.719 \pm 1.358$ |
| $50.0-100.0$ | $4.178 \pm 1.699$ | $6.110 \pm 2.307$ | $4.857 \pm 1.368$ |
| $100.0-300.0$ | $3.597 \pm 2.137$ | $3.743 \pm 4.329$ | $3.626 \pm 1.916$ |
| $2.5-300.0$ | $9.845 \pm 1.047$ | $9.883 \pm 1.018$ | $9.865 \pm 0.730$ |

Table 6.52 and Table 6.53 show the comparison of $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$in Neutrino Beam Mode and Anti-neutrino Beam Mode in 7 bins before and after smearing matrix


Figure 6.74: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}:$ FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of $E_{\nu}$ using variable BN.


Figure 6.75: Distribution of the sum of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}: \mathrm{FocP}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}: \mathrm{FocN}$ as a function of $E_{\nu}$ using variable BN.
correction. From these two tables, it is obvious that in each bins, the $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ are similar in this two beam modes. There are overlap in each bins. In the whole bin from 2.5 to $300 \mathrm{GeV}, \mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ are very close to each other in this two beam modes.

### 6.7 Comparison of Coherent $\pi^{-}$and Coherent $\pi^{+}$

With the average value of $\mathrm{R}\left(\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}\right)$, let's compare the average value with the result of coherent $\pi^{+}$which got by Xinchun Tian. In Table 6.54, it gives out that the Table 6.54: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of Evis using variable BN.

| Evis $(\mathrm{GeV})$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right) / \mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $12.438 \pm 4.012$ | $7.06 \pm 1.02$ | $1.762 \pm 0.623$ |
| $8.0-15.0$ | $14.824 \pm 2.127$ | $7.23 \pm 0.37$ | $2.053 \pm 0.313$ |
| $15.0-20.0$ | $10.039 \pm 2.056$ | $6.43 \pm 0.36$ | $1.561 \pm 0.331$ |
| $20.0-30.0$ | $9.904 \pm 1.556$ | $5.39 \pm 0.27$ | $1.837 \pm 0.303$ |
| $30.0-50.0$ | $7.728 \pm 1.341$ | $4.52 \pm 0.24$ | $1.710 \pm 0.310$ |
| $50.0-100.0$ | $4.762 \pm 1.370$ | $3.08 \pm 0.23$ | $1.546 \pm 0.460$ |
| $100.0-300.0$ | $3.932 \pm 2.114$ | $2.06 \pm 0.31$ | $1.909 \pm 1.066$ |
| $2.5-300.0$ | $9.865 \pm 0.730$ | $4.86 \pm 0.12$ | $2.030 \pm 0.158$ |

ratio between $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ is $2.030 \pm 0.158$, which is consistent to the theoratical prediction, because in NOMAD beam, the ratio between the number of $\nu_{\mu}$ and $\bar{\nu}_{\nu}$ is 2:1.

Table 6.55 shows the comparison between $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ after smearing matrix correction.


Figure 6.76: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ using variable BN.


Figure 6.77: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ using variable BN.

Table 6.55: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right) / \mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $12.356 \pm 3.879$ | $5.99 \pm 1.01$ | $2.063 \pm 0.735$ |
| $8.0-15.0$ | $14.800 \pm 2.186$ | $6.93 \pm 0.37$ | $2.136 \pm 0.335$ |
| $15.0-20.0$ | $10.400 \pm 2.064$ | $6.39 \pm 0.35$ | $1.628 \pm 0.335$ |
| $20.0-30.0$ | $9.854 \pm 1.562$ | $5.56 \pm 0.26$ | $1.772 \pm 0.293$ |
| $30.0-50.0$ | $7.719 \pm 1.358$ | $4.61 \pm 0.23$ | $1.674 \pm 0.306$ |
| $50.0-100.0$ | $4.857 \pm 1.368$ | $3.16 \pm 0.22$ | $1.537 \pm 0.446$ |
| $100.0-300.0$ | $3.626 \pm 1.916$ | $2.25 \pm 0.28$ | $1.612 \pm 0.875$ |
| $2.5-300.0$ | $9.865 \pm 0.730$ | $4.86 \pm 0.12$ | $2.030 \pm 0.158$ |

### 6.8 Systematic Uncertainties

## Background Subtraction Procedure

In the procedure of the normalization of background, there are two methods have been used. One is to normalize the background bin by bin with variable normalization factor. The other one is to normalize the background by the normalization factor of the whole bin from 2.5 to 300 GeV . From this subsection, we would give out the result of analysis using a single background normalization factor. The difference between these two ways of normalization of background is considered as a systematic uncertainty which is shown in Table 6.61. In Table 6.61, $\delta_{R}$ equals to the R calculated from a fixed BN minus the R calculated from variable BN . $\frac{\delta_{R}}{R}$ equals $\delta_{R}$ divide by the $R$ calculated from a fixed BN.

Table 6.56 shows the comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ in neutrino beam mode and antineutrino beam mode calculated from a single background normalization.

The averaged value of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ in neutrino beam mode and anti-neutrino beam mode are calculated by the Equation (6.10) and Equation (6.11), and are shown in Table 6.57.


Figure 6.78: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}:$ FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN calculated from a fixed BN.


Figure 6.79: Distribution of the sum of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}:$ FocN calculated from a fixed BN.


Figure 6.80: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ calculated from a fixed BN .


Figure 6.81: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ calculated from a fixed BN .

Table 6.56: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN calculated from a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}:\right.$FocN $)$ | $<\mathrm{R}>$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $13.656 \pm 5.143$ | $5.650 \pm 7.238$ | $10.97 \pm 4.192$ |
| $8.0-15.0$ | $16.321 \pm 3.420$ | $13.209 \pm 2.876$ | $14.498 \pm 2.201$ |
| $15.0-20.0$ | $9.136 \pm 2.991$ | $11.817 \pm 2.611$ | $10.658 \pm 1.967$ |
| $20.0-30.0$ | $12.254 \pm 2.562$ | $8.648 \pm 1.863$ | $9.895 \pm 1.507$ |
| $30.0-50.0$ | $7.252 \pm 1.806$ | $8.248 \pm 1.918$ | $7.720 \pm 1.315$ |
| $50.0-100.0$ | $4.778 \pm 1.451$ | $6.864 \pm 2.070$ | $5.465 \pm 1.188$ |
| $100.0-300.0$ | $3.373 \pm 2.087$ | $4.684 \pm 3.155$ | $3.772 \pm 1.741$ |
| $2.5-300$ | $9.845 \pm 1.047$ | $9.883 \pm 1.018$ | $9.865 \pm 0.730$ |

Table 6.57: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ calculated from a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right) / \mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $10.970 \pm 4.192$ | $5.809 \pm 1.01$ | $1.888 \pm 0.793$ |
| $8.0-15.0$ | $14.498 \pm 2.201$ | $7.002 \pm 0.37$ | $2.071 \pm 0.333$ |
| $15.0-20.0$ | $10.658 \pm 1.967$ | $6.405 \pm 0.35$ | $1.664 \pm 0.320$ |
| $20.0-30.0$ | $9.895 \pm 1.507$ | $5.599 \pm 0.26$ | $1.767 \pm 0.281$ |
| $30.0-50.0$ | $7.720 \pm 1.315$ | $4.582 \pm 0.23$ | $1.685 \pm 0.299$ |
| $50.0-100.0$ | $5.465 \pm 1.188$ | $3.156 \pm 0.22$ | $1.732 \pm 0.395$ |
| $100.0-300.0$ | $3.772 \pm 1.741$ | $2.208 \pm 0.28$ | $1.708 \pm 0.817$ |
| $2.5-300.0$ | $9.865 \pm 0.730$ | $4.86 \pm 0.12$ | $2.030 \pm 0.158$ |



Figure 6.82: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ calculated from a fixed BN to $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ calculated from variable BN.

Table 6.58: Comparison between different background subtractions.

| $E_{\nu}(\mathrm{GeV})$ | FixBN $(\nu$-Mode $)$ | VarBN $(\nu$-Mode $)$ | FixBN $(\bar{\nu}$-Mode $)$ | VarBN $(\bar{\nu}$-Mode $)$ | $\langle\mathrm{R}\rangle(\nu$-Mode $)$ | $\langle\mathrm{R}\rangle(\bar{\nu}$-Mode $)$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8$ | $13.656 \pm 5.143$ | $14.385 \pm 5.112$ | $5.650 \pm 7.238$ | $9.652 \pm 5.956$ | $10.97 \pm 4.192$ | $12.356 \pm 3.879$ |
| $8-15$ | $16.321 \pm 3.420$ | $16.426 \pm 3.432$ | $13.209 \pm 2.876$ | $13.690 \pm 2.836$ | $14.498 \pm 2.201$ | $14.800 \pm 2.186$ |
| $15-20$ | $9.136 \pm 2.991$ | $8.578 \pm 3.221$ | $11.817 \pm 2.611$ | $11.669 \pm 2.689$ | $10.658 \pm 1.967$ | $10.400 \pm 2.064$ |
| $20-30$ | $12.254 \pm 2.562$ | $12.393 \pm 2.587$ | $8.648 \pm 1.863$ | $8.397 \pm 1.960$ | $9.895 \pm 1.507$ | $9.854 \pm 1.562$ |
| $30-50$ | $7.252 \pm 1.806$ | $7.432 \pm 1.834$ | $8.248 \pm 1.918$ | $8.067 \pm 2.022$ | $7.720 \pm 1.315$ | $7.719 \pm 1.358$ |
| $50-100$ | $4.778 \pm 1.451$ | $4.178 \pm 1.699$ | $6.864 \pm 2.070$ | $6.110 \pm 2.307$ | $5.465 \pm 1.188$ | $4.857 \pm 1.386$ |
| $100-300$ | $3.373 \pm 2.087$ | $3.597 \pm 2.137$ | $4.684 \pm 3.155$ | $3.743 \pm 4.329$ | $3.772 \pm 1.741$ | $3.626 \pm 1.916$ |

Smearing matrix for Coh $\pi^{-}$:FocP
Table 6.59: Smearing matrix of $\operatorname{Coh} \pi^{-}$:FocP

| Evis $\backslash E_{\nu}$ | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 28.68 | 1.09 | 0.01 | 0 | 0 | 0 | 0 |
| $8.0-15.0$ | 1.06 | 56.72 | 1.45 | 0.02 | 0 | 0 | 0 |
| $15.0-20.0$ | 0 | 1.51 | 28.31 | 1.25 | 0 | 0 | 0 |
| $20.0-30.0$ | 0 | 0 | 1.55 | 37.55 | 1.34 | 0 | 0 |
| $30.0-50.0$ | 0 | 0 | 0.01 | 1.43 | 36.44 | 1.03 | 0 |
| $50.0-100.0$ | 0 | 0 | 0 | 0 | 1.11 | 20.96 | 0.23 |
| $100.0-300.0$ | 0 | 0 | 0 | 0 | 0 | 0.35 | 2.86 |

Smearing matrix for Coh $\pi^{-}$:FocN
Table 6.60: Smearing matrix of Coh $\pi^{-}$:FocN

| Evis $\backslash E_{\nu}$ | $2.5-8.0$ | $8.0-15.0$ | $15.0-20.0$ | $20.0-30.0$ | $30.0-50.0$ | $50.0-100.0$ | $100.0-300.0$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 12.98 | 0.93 | 0 | 0 | 0 | 0 | 0 |
| $8.0-15.0$ | 1.09 | 76.38 | 2.28 | 0.02 | 0 | 0 | 0 |
| $15.0-20.0$ | 0.01 | 2.33 | 45.38 | 2.48 | 0.01 | 0 | 0 |
| $20.0-30.0$ | 0 | 0.03 | 1.76 | 48.52 | 1.63 | 0 | 0 |
| $30.0-50.0$ | 0 | 0 | 0 | 1.13 | 28.01 | 0 | 0 |
| $50.0-100.0$ | 0 | 0 | 0 | 0 | 0.58 | 12.02 | 0.22 |
| $100.0-300.0$ | 0 | 0 | 0 | 0 | 0 | 0.12 | 1.37 |

Table 6.61: Systematic on Background Subtraction.

| $E_{\nu}(\mathrm{GeV})$ | $\delta_{R^{-}} \nu$ mode | $\frac{\delta_{R}-\nu \text { mode }}{R}$ | $\delta_{R^{-}} \bar{\nu}$ mode | $\frac{\delta_{R}-\bar{\nu} \text { mode }}{R}$ |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | -0.729 | -0.053 | -4.002 | -0.708 |
| $8-15$ | -0.105 | -0.006 | -0.481 | -0.036 |
| $15-20$ | 0.558 | 0.061 | 0.148 | 0.013 |
| $20-30$ | -0.139 | -0.011 | 0.251 | 0.029 |
| $30-50$ | -0.18 | -0.025 | 0.181 | 0.022 |
| $50-100$ | 0.6 | 0.126 | 0.754 | 0.109 |
| $100-300$ | -0.224 | -0.066 | 0.941 | 0.201 |

## Systematic from Final State Interaction

Another very important systematic uncertainty comes from the final state interaction, when the mesons are produced within the nucleus and interact with the nucleons. Coherent process is not sensitive to the final state interactions, because the pion in the final state does not emerge from the nucleons. However final state interactions are affecting the background processes as well as the $\bar{\nu}_{\mu}$ charged current events and $\nu_{\mu}$ charged current events. Therefore, we used NUANCE event generator to estimate such effect by turning on/off the final state interaction(FSI) effect. Table 6.62 shows the final state interaction error as a function of energy in 14 bins of coherent $\pi^{+}$in neutrino beam mode analysis.

## Use of Different Signal Models

In this analysis, the Rein-Sehgal(RS) Model [44] is used to simulate the Coherent $\pi^{-}$ interaction in NOMAD detector. The RS model used by both NUANCE and NEUGEN, describes the weak current only in terms of the pion field; The $Q^{2}$ dependence of the cross-section is assumed to have a dipole form. Other calculations reply on meson-dominance models [15] which include the dominant contributions from the $\rho$ and $a_{1}$ mesons, for example, Berger-Sehgal(BS) Model [15]. The Monte Carlo Coher-

Table 6.62: Final state interaction (FSI) error as a function of energy in 14 bins

| $\operatorname{Evis}(\mathrm{GeV})$ | FSI |
| :--- | :---: |
| $2.5-6.0$ | 6.5051 |
| $6.0-8.0$ | 10.5073 |
| $8.0-10.0$ | 6.4971 |
| $10.0-12.0$ | 4.6806 |
| $12.0-15.0$ | 3.9991 |
| $15.0-20.0$ | 4.5252 |
| $20.0-25.0$ | 2.4506 |
| $25.0-30.0$ | 2.3983 |
| $30.0-40.0$ | 1.7477 |
| $40.0-50.0$ | 1.6448 |
| $50.0-70.0$ | 0.7053 |
| $70.0-100.0$ | -2.8253 |
| $100.0-130.0$ | -7.0350 |
| $130.0-300.0$ | 17.7462 |
| $2.5-300.0$ | 3.7120 |

ent $\pi^{-}$events used above are simulated by RS Model. As a check, the BS Model is also used to simulate the coherent $\pi^{-}$interactions. In this section, we will give out the result of the analysis using the coherent $\pi^{-}$events simulated by the BS Model and compared it with the result of the RS Model. Because the momentum of the hadron in BS model is lower than in the RS model, the efficiency is lower. Thus, according to the theory, the corrected signal of BS Model should be more than the corrected signal in RS model. The difference between the result of BS and RS model is also considered as systematic uncertainty shown in Table 6.80. In Table 6.80, $\delta_{R}$ equals to the R calculated from analysis with RS model simulation minus the R calculated from analysis with BS model simulation. $\frac{\delta_{R}}{R}$ equals $\delta_{R}$ divide by the R calculated from analysis with RS model simulation.

Similar to the analysis with the coherent $\pi^{-}$events simulated using the ReinSehgal (RS) model, the coherent $\pi^{-}$events simulated using the Berger-Sehgal (BS) Model are used in the analysis of neutrino beam mode data and anti-neutrino beam mode data. In this section, the results of the analysis with the coherent $\pi^{-}$events
simulated using BS model are given out and compared with the result of analysis with the coherent $\pi^{-}$events simulated using the RS model.

## Neutrino Beam Mode Analysis:

Table 6.63 and Table 6.64 show the result of background and signal normalization factors (BN and SN) calculated from variable BN and fixed BN.

Table 6.63: BN and SN table in 7 Bins using variable BN depends on the Evis in 7 bins.

| Evis(GeV) | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 0.8302 | 0.5041 | 0.2787 | $14 \pm 6.215$ |
| $8-15$ | 0.7921 | 1.08 | 0.2023 | $22.09 \pm 4.389$ |
| $15-20$ | 1.354 | 0.7222 | 0.4074 | $11.2 \pm 4.198$ |
| $20-30$ | 0.8784 | 1.236 | 0.3057 | $15.14 \pm 3.199$ |
| $30-50$ | 0.8652 | 0.9873 | 0.2973 | $8.492 \pm 2.17$ |
| $50-100$ | 1.85 | 0.9401 | 0.7219 | $4.556 \pm 2.135$ |
| $100-300$ | 0.6892 | 1.756 | 1.561 | $3.639 \pm 2.803$ |
| $2.5-300$ | 0.9848 | 0.9612 | 0.1266 | $11.97 \pm 1.324$ |

Table 6.64: BN and SN table in 7 Bins using a fixed BN.

| Evis(GeV) | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 0.9848 | 0.5041 | 0.1266 | $13.1 \pm 6.253$ |
| $8-15$ | 0.9848 | 1.08 | 0.1266 | $21.22 \pm 4.443$ |
| $15-20$ | 0.9848 | 0.7222 | 0.1266 | $12.12 \pm 3.85$ |
| $20-30$ | 0.9848 | 1.236 | 0.1266 | $14.91 \pm 3.177$ |
| $30-50$ | 0.9848 | 0.9873 | 0.1266 | $8.265 \pm 2.145$ |
| $50-100$ | 0.9848 | 0.9401 | 0.1266 | $5.53 \pm 1.708$ |
| $100-300$ | 0.9848 | 1.756 | 0.1266 | $3.412 \pm 2.634$ |
| $2.5-300$ | 0.9848 | 0.9612 | 0.1266 | $11.97 \pm 1.324$ |

Table 6.65: Norm-bkg, Corr-sig as a function of Evis in 7 bins using variable BN.

| Evis(GeV) | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 3.72 | 0.83 | 3.09 | 12 | 8.91 | 0.153 | $58.37 \pm 25 \pm 6.791$ |
| $8.0-15.0$ | 7.06 | 0.792 | 5.59 | 40 | 34.4 | 0.275 | $125 \pm 24.3 \pm 5.192$ |
| $15.0-20.0$ | 2.93 | 1.35 | 3.96 | 17 | 13 | 0.32 | $40.68 \pm 14.79 \pm 3.724$ |
| $20.0-30.0$ | 4.17 | 0.878 | 3.66 | 33 | 29.3 | 0.33 | $89.03 \pm 18.41 \pm 3.867$ |
| $30.0-50.0$ | 5.05 | 0.865 | 4.37 | 27 | 22.6 | 0.349 | $64.75 \pm 15.98 \pm 4.299$ |
| $50.0-100.0$ | 3.05 | 1.85 | 5.65 | 18 | 12.4 | 0.346 | $35.7 \pm 15.47 \pm 6.365$ |
| $100.0-300.0$ | 0.554 | 0.689 | 0.382 | 3 | 2.62 | 0.278 | $9.434 \pm 6.563 \pm 3.117$ |
| $2.5-300.0$ | 26.5 | 0.985 | 26.1 | 150 | 124 | 0.275 | $450.6 \pm 48.34 \pm 12.23$ |

The total background (Tot-bkg), background normalization (BN), normalized background (Norm-bkg), Data, raw signal(Raw-sig), Efficiency (Eff), corrected signal (Corr-Sig) of the analysis using BS model are shown in Table 6.65. The total number corrected signal in the whole bin from 2.5 to 300 GeV is $450.6 \pm 49.86$, which is bigger than the total number of corrected signal in the result of RS model which is $370.6 \pm 38.26 \pm 9.509$.

Table 6.66: Corrected signal (Corr-Sig) as a function of Evis in 7 bins using variable BN.

| Evis(GeV) | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 58.37 | 4168.3 | $14 \pm 6.215$ |
| $8-15$ | 11.84 | 125 | 5660.2 | $22.09 \pm 4.389$ |
| $15-20$ | 17.40 | 40.68 | 3633.3 | $11.2 \pm 4.198$ |
| $20-30$ | 24.60 | 89.03 | 5881.0 | $15.14 \pm 3.199$ |
| $30-50$ | 38.47 | 64.75 | 7625.0 | $8.492 \pm 2.17$ |
| $50-100$ | 71.54 | 35.7 | 7835.2 | $4.556 \pm 2.135$ |
| $100-300$ | 142.70 | 9.434 | 2592.0 | $3.639 \pm 2.803$ |
| $2.5-300$ | 25.00 | 450.6 | 37645.6 | $11.97 \pm 1.324$ |

Table 6.67: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$-CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 58.31 | 4168.3 | $14 \pm 6.079$ |
| $8-15$ | 11.84 | 124.3 | 5660.2 | $21.96 \pm 4.523$ |
| $15-20$ | 17.40 | 42.76 | 3633.3 | $11.77 \pm 4.129$ |
| $20-30$ | 24.60 | 87.30 | 5881.0 | $14.84 \pm 3.210$ |
| $30-50$ | 38.47 | 65.32 | 7625.0 | $8.567 \pm 2.222$ |
| $50-100$ | 71.54 | 36.51 | 7835.2 | $4.660 \pm 2.186$ |
| $100-300$ | 142.70 | 8.558 | 2592.0 | $3.301 \pm 2.493$ |
| $2.5-300$ | 25.00 | 450.6 | 37645.6 | $11.97 \pm 1.324$ |

Table 6.68: Norm-bkg, Corr-sig as a function of Evis in 7 bins using a fixed BN(the the beam(flux) reweight is applied to both of the $\bar{\nu}_{\mu}$-CC events and $\nu_{\mu}$-CC events).

| Evis(GeV) | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 3.72 | 0.985 | 3.66 | 12 | 8.34 | 0.153 | $54.61 \pm 25.88 \pm 3.085$ |
| $8.0-15.0$ | 7.06 | 0.985 | 6.95 | 40 | 33 | 0.275 | $120.1 \pm 24.94 \pm 3.248$ |
| $15.0-20.0$ | 2.93 | 0.985 | 2.88 | 17 | 14.1 | 0.32 | $44.05 \pm 13.94 \pm 1.157$ |
| $20.0-30.0$ | 4.17 | 0.985 | 4.11 | 33 | 28.9 | 0.33 | $87.68 \pm 18.61 \pm 1.601$ |
| $30.0-50.0$ | 5.05 | 0.985 | 4.98 | 27 | 22 | 0.349 | $63.02 \pm 16.26 \pm 1.831$ |
| $50.0-100.0$ | 3.05 | 0.985 | 3.01 | 18 | 15 | 0.346 | $43.33 \pm 13.33 \pm 1.116$ |
| $100.0-300.0$ | 0.554 | 0.985 | 0.546 | 3 | 2.45 | 0.278 | $8.844 \pm 6.823 \pm 0.2527$ |
| $2.5-300.0$ | 26.5 | 0.985 | 26.1 | 150 | 124 | 0.275 | $450.6 \pm 48.34 \pm 12.23$ |

Table 6.68 shows the result of total background (Tot-bkg), background normalization (BN), normalized background (Norm-bkg), Data, raw signal (Raw-sig), Efficiency (Eff), and corrected signal (Corr-Sig) which are calculated using a single background normalization.

Table 6.69: Corrected signal (Corr-Sig) as a function of Evis in 7 bins using a fixed BN.

| Evis $(\mathrm{GeV})$ | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}($ Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 54.61 | 4168.3 | $13.1 \pm 6.253$ |
| $8-15$ | 11.84 | 120.1 | 5660.2 | $21.22 \pm 4.443$ |
| $15-20$ | 17.40 | 44.05 | 3633.3 | $12.12 \pm 3.85$ |
| $20-30$ | 24.60 | 87.68 | 5881.0 | $14.91 \pm 3.177$ |
| $30-50$ | 38.47 | 63.02 | 7625.0 | $8.265 \pm 2.145$ |
| $50-100$ | 71.54 | 43.33 | 7835.2 | $5.53 \pm 1.708$ |
| $100-300$ | 142.70 | 8.844 | 2592.0 | $3.412 \pm 2.634$ |
| $2.5-300$ | 25.00 | 450.6 | 37645.6 | $11.97 \pm 1.324$ |

Table 6.70: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins using a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$-CC | $\mathrm{R} \times 10^{-3}($ Stat. $)$ |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 54.60 | 4168.3 | $13.10 \pm 6.116$ |
| $8-15$ | 11.84 | 119.53 | 5660.2 | $21.12 \pm 4.565$ |
| $15-20$ | 17.40 | 45.64 | 3633.3 | $12.562 \pm 3.816$ |
| $20-30$ | 24.60 | 86.13 | 5881.0 | $14.645 \pm 3.178$ |
| $30-50$ | 38.47 | 64.05 | 7625.0 | $8.4 \pm 2.177$ |
| $50-100$ | 71.54 | 43.53 | 7835.2 | $5.556 \pm 1.778$ |
| $100-300$ | 142.70 | 8.14 | 2592.0 | $3.14 \pm 2.332$ |
| $2.5-300$ | 25.00 | 450.6 | 37645.6 | $11.97 \pm 1.324$ |

## Anti-neutrino Beam Mode Analysis:

Table 6.71: BN and SN table in 7 Bins using variable BN depends on the Evis in 7 bins.

| Evis $(\mathrm{GeV})$ | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 0.3439 | 0.2867 | 0.1952 | $8.538 \pm 6.822$ |
| $8.0-15.0$ | 1.055 | 0.8238 | 0.1975 | $16.42 \pm 3.478$ |
| $15.0-20.0$ | 1.24 | 0.9688 | 0.2968 | $13.77 \pm 3.141$ |
| $20.0-30.0$ | 1.283 | 0.9928 | 0.3057 | $10.12 \pm 2.333$ |
| $30.0-50.0$ | 1.479 | 1.35 | 0.4713 | $9.046 \pm 2.372$ |
| $50.0-100.0$ | 2.006 | 1.759 | 1.039 | $6.767 \pm 2.685$ |
| $100.0-300.0$ | 1.727 | 2.658 | 3.158 | $4.309 \pm 4.224$ |
| $2.5-300.0$ | 1.17 | 1.009 | 0.1294 | $11.6 \pm 1.226$ |

Table 6.72: BN and SN table in 7 Bins using a fixed BN.

| Evis(GeV) | BN | SN | $\delta_{B N}$ | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 1.17 | 0.2867 | 0.1294 | $2.363 \pm 8.942$ |
| $8.0-15.0$ | 1.17 | 0.8238 | 0.1294 | $15.97 \pm 3.494$ |
| $15.0-20.0$ | 1.17 | 0.9688 | 0.1294 | $13.91 \pm 3.064$ |
| $20.0-30.0$ | 1.17 | 0.9928 | 0.1294 | $10.33 \pm 2.236$ |
| $30.0-50.0$ | 1.17 | 1.35 | 0.1294 | $9.474 \pm 2.199$ |
| $50.0-100.0$ | 1.17 | 1.759 | 0.1294 | $7.597 \pm 2.24$ |
| $100.0-300.0$ | 1.17 | 2.658 | 0.1294 | $4.655 \pm 3.52$ |
| $2.5-300.0$ | 1.17 | 1.009 | 0.1294 | $11.6 \pm 1.226$ |

Table 6.73: Norm-bkg, Corr-sig as a function of Evis in 7 bins using variable BN.

| Evis(GeV) | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 2.02 | 0.344 | 0.694 | 3 | 2.31 | 0.197 | $11.7 \pm 9.351$ |
| $8.0-15.0$ | 8.64 | 1.05 | 9.12 | 45 | 35.9 | 0.282 | $127.4 \pm 26.97$ |
| $15.0-20.0$ | 4.33 | 1.24 | 5.37 | 34 | 28.6 | 0.327 | $87.46 \pm 19.95$ |
| $20.0-30.0$ | 5.91 | 1.28 | 7.58 | 39 | 31.4 | 0.34 | $92.43 \pm 21.32$ |
| $30.0-50.0$ | 3.87 | 1.48 | 5.73 | 31 | 25.3 | 0.373 | $67.77 \pm 17.77$ |
| $50.0-100.0$ | 2.15 | 2.01 | 4.32 | 19 | 14.7 | 0.383 | $38.36 \pm 15.22$ |
| $100.0-300.0$ | 0.346 | 1.73 | 0.597 | 3 | 2.4 | 0.351 | $6.841 \pm 6.706$ |
| $2.5-300.0$ | 27.3 | 1.17 | 31.9 | 174 | 142 | 0.31 | $457.6 \pm 48.36$ |

The total background (Tot-bkg), background normalization (BN), normalized background (Norm-bkg), Data, raw signal (Raw-sig), Efficiency (Eff), corrected signal (Corr-Sig) of the analysis using BS model are shown in Table 6.46. The total number corrected signal in the whole bin from 2.5 to 300 GeV is $457.6 \pm 48.36$, which is bigger than the total number of corrected signal in the result of RS model which is $389.7 \pm 39.01 \pm 9.492$.

Table 6.74: Corrected signal (Corr-sig) as a function of Evis in 7 bins using variable BN.

| Evis(GeV) | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 11.7 | 1370.7 | $8.538 \pm 6.822$ |
| $8-15$ | 11.84 | 127.4 | 7755.0 | $16.42 \pm 3.478$ |
| $15-20$ | 17.40 | 87.46 | 6352.6 | $13.77 \pm 3.141$ |
| $20-30$ | 24.60 | 92.43 | 9136.4 | $10.12 \pm 2.333$ |
| $30-50$ | 38.47 | 67.77 | 7492.4 | $9.046 \pm 2.372$ |
| $50-100$ | 71.54 | 38.36 | 5668.7 | $6.767 \pm 2.685$ |
| $100-300$ | 142.70 | 6.841 | 1587.6 | $4.309 \pm 4.224$ |
| $2.5-300$ | 25.00 | 457.6 | 39435.6 | $11.6 \pm 1.226$ |

Table 6.75: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $<\mathrm{E}>$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$-CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 12.2 | 1370.7 | $8.899 \pm 6.476$ |
| $8-15$ | 11.84 | 127.5 | 7755.0 | $16.44 \pm 3.565$ |
| $15-20$ | 17.40 | 87.20 | 6352.6 | $13.73 \pm 3.124$ |
| $20-30$ | 24.60 | 93.89 | 9136.4 | $10.28 \pm 2.383$ |
| $30-50$ | 38.47 | 66.77 | 7492.4 | $8.911 \pm 2.367$ |
| $50-100$ | 71.54 | 38.28 | 5668.7 | $6.753 \pm 2.768$ |
| $100-300$ | 142.70 | 6.079 | 1587.6 | $3.829 \pm 3.606$ |
| $2.5-300$ | 25.00 | 457.6 | 39435.6 | $11.6 \pm 1.226$ |

Table 6.76: Norm-bkg, Corr-sig as a function of Evis in 7 bins using a fixed BN.

| Evis $(\mathrm{GeV})$ | Tot-bkg | BN | Norm-bkg | Data | Raw-sig | Eff | Corr-Sig. (Err=Stat., BN) |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8.0$ | 2.02 | 1.17 | 2.36 | 3 | 0.638 | 0.197 | $3.239 \pm 12.26$ |
| $8.0-15.0$ | 8.64 | 1.17 | 10.1 | 45 | 34.9 | 0.282 | $123.8 \pm 27.09$ |
| $15.0-20.0$ | 4.33 | 1.17 | 5.06 | 34 | 28.9 | 0.327 | $88.38 \pm 19.47$ |
| $20.0-30.0$ | 5.91 | 1.17 | 6.91 | 39 | 32.1 | 0.34 | $94.4 \pm 20.43$ |
| $30.0-50.0$ | 3.87 | 1.17 | 4.53 | 31 | 26.5 | 0.373 | $70.98 \pm 16.48$ |
| $50.0-100.0$ | 2.15 | 1.17 | 2.52 | 19 | 16.5 | 0.383 | $43.07 \pm 12.7$ |
| $100.0-300.0$ | 0.346 | 1.17 | 0.405 | 3 | 2.6 | 0.351 | $7.39 \pm 5.588$ |
| $2.5-300.0$ | 27.3 | 1.17 | 31.9 | 174 | 142 | 0.31 | $457.6 \pm 48.36$ |

Table 6.77: Corrected signal (Corr-sig) as a function of Evis in 7 bins using a fixed BN.

| Evis(GeV) | $<\mathrm{E}>$ | Corr-Sig | $\bar{\nu}_{\mu}$ - CC | $\mathrm{R} \times 10^{-3}$ (Stat.) |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 3.239 | 1370.7 | $2.363 \pm 8.942$ |
| $8-15$ | 11.84 | 123.8 | 7755.0 | $15.97 \pm 3.494$ |
| $15-20$ | 17.40 | 88.38 | 6352.6 | $13.91 \pm 3.064$ |
| $20-30$ | 24.60 | 94.4 | 9136.4 | $10.33 \pm 2.236$ |
| $30-50$ | 38.47 | 70.98 | 7492.4 | $9.474 \pm 2.199$ |
| $50-100$ | 71.54 | 43.07 | 5668.7 | $7.597 \pm 2.24$ |
| $100-300$ | 142.70 | 7.39 | 1587.6 | $4.655 \pm 3.52$ |
| $2.5-300$ | 25.00 | 457.6 | 39435.6 | $11.6 \pm 1.226$ |

Table 6.78: Corrected signal (Corr-Sig-Enus) as a function of $E_{\nu}$ in 7 bins using a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $<\mathrm{E}\rangle$ | Corr-Sig-Enus | $\bar{\nu}_{\mu}$-CC | $\mathrm{R} \times 10^{-3}($ Stat. $)$ |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 6.25 | 4.407 | 1370.7 | $3.215 \pm 8.418$ |
| $8-15$ | 11.84 | 123.5 | 7755.0 | $15.93 \pm 3.609$ |
| $15-20$ | 17.40 | 88.03 | 6352.6 | $13.86 \pm 3.049$ |
| $20-30$ | 24.60 | 96.00 | 9136.4 | $10.51 \pm 2.281$ |
| $30-50$ | 38.47 | 70.04 | 7492.4 | $9.348 \pm 2.186$ |
| $50-100$ | 71.54 | 42.83 | 5668.7 | $7.556 \pm 2.314$ |
| $100-300$ | 142.70 | 6.584 | 1587.6 | $4.147 \pm 3.005$ |
| $2.5-300$ | 25.00 | 457.6 | 39435.6 | $11.6 \pm 1.226$ |

Table 6.79: Comparison between RS and BS signal model simulation.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{RS}(\nu$-mode $)$ | $\mathrm{BS}(\nu$-mode $)$ | $\mathrm{RS}(\bar{\nu}$-mode $)$ | $\mathrm{BS}(\bar{\nu}$-mode $)$ | $<\mathrm{R}\rangle(\mathrm{RS})$ | $<\mathrm{R}>(\mathrm{BS})$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| $2.5-8$ | $13.656 \pm 5.143$ | $13.100 \pm 6.116$ | $5.650 \pm 7.238$ | $3.215 \pm 8.418$ | $10.97 \pm 4.192$ | $9.685 \pm 4.948$ |
| $8-15$ | $16.321 \pm 3.420$ | $21.12 \pm 4.565$ | $13.209 \pm 2.876$ | $15.930 \pm 3.609$ | $14.498 \pm 2.201$ | $17.926 \pm 2.831$ |
| $15-20$ | $9.136 \pm 2.991$ | $12.562 \pm 3.816$ | $11.817 \pm 2.611$ | $13.86 \pm 3.049$ | $10.658 \pm 1.967$ | $13.354 \pm 2.382$ |
| $20-30$ | $12.254 \pm 2.562$ | $14.645 \pm 3.178$ | $8.648 \pm 1.863$ | $10.510 \pm 2.281$ | $9.895 \pm 1.507$ | $11.916 \pm 1.853$ |
| $30-50$ | $7.252 \pm 1.806$ | $8.400 \pm 2.177$ | $8.248 \pm 1.918$ | $9.348 \pm 2.183$ | $7.720 \pm 1.315$ | $8.872 \pm 1.543$ |
| $50-100$ | $4.778 \pm 1.451$ | $5.556 \pm 1.778$ | $6.864 \pm 2.070$ | $7.556 \pm 2.314$ | $5.465 \pm 1.188$ | $6.298 \pm 1.410$ |
| $100-300$ | $3.373 \pm 2.087$ | $3.140 \pm 2.332$ | $4.684 \pm 3.155$ | $4.147 \pm 3.005$ | $3.772 \pm 1.741$ | $3.519 \pm 1.842$ |
| $2.5-300$ | $9.845 \pm 1.047$ | $11.97 \pm 1.324$ | $9.883 \pm 1.018$ | $11.6 \pm 1.226$ | $9.865 \pm 0.730$ | $11.771 \pm 0.9$ |

Table 6.80: Systematic on signal modeling.

| $E_{\nu}(\mathrm{GeV})$ | $\delta_{R^{-}}-$mode | $\frac{\delta_{R}-\nu \text { mode }}{R}$ | $\delta_{R^{-}} \bar{\nu}$ mode | $\frac{\delta_{R}-\bar{\nu} \text { mode }}{R}$ |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-8$ | 0.556 | 0.041 | 2.435 | 0.431 |
| $8-15$ | -4.799 | -0.293 | -2.721 | -0.206 |
| $15-20$ | -3.426 | -0.375 | -2.043 | -0.173 |
| $20-30$ | -2.391 | -0.195 | -1.862 | -0.215 |
| $30-50$ | -1.148 | -0.158 | -1.1 | -0.133 |
| $50-100$ | -0.768 | -0.161 | -0.692 | -0.101 |
| $100-300$ | 0.233 | 0.069 | 0.537 | 0.115 |
| $2.5-300$ | -2.125 | -0.216 | -1.717 | -0.174 |



Figure 6.83: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ between BS and RS Model calculated from a fixed BN.

Table 6.81: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of Evis using variable BN.

| Evis(GeV) | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}:\right.$FocN $)$ | $<\mathrm{R}\rangle$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $14 \pm 6.215$ | $8.538 \pm 6.822$ | $11.523 \pm 4.594$ |
| $8.0-15.0$ | $22.09 \pm 4.389$ | $16.42 \pm 3.478$ | $18.607 \pm 2.726$ |
| $15.0-20.0$ | $11.2 \pm 4.198$ | $13.77 \pm 3.141$ | $12.848 \pm 2.515$ |
| $20.0-30.0$ | $15.14 \pm 3.199$ | $10.12 \pm 2.333$ | $11.863 \pm 1.885$ |
| $30.0-50.0$ | $8.492 \pm 2.17$ | $9.046 \pm 2.372$ | $8.744 \pm 1.601$ |
| $50.0-100.0$ | $4.556 \pm 2.135$ | $6.767 \pm 2.685$ | $5.412 \pm 1.671$ |
| $100.0-300.0$ | $3.639 \pm 2.803$ | $4.309 \pm 4.224$ | $3.844 \pm 2.336$ |
| $2.5-300.0$ | $11.97 \pm 1.324$ | $11.6 \pm 1.226$ | $11.771 \pm 0.900$ |

Table 6.82: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of $E_{\nu}$ using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocN $)$ | $<\mathrm{R}>$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $14 \pm 6.079$ | $8.899 \pm 6.476$ | $11.611 \pm 4.432$ |
| $8.0-15.0$ | $21.96 \pm 4.523$ | $16.44 \pm 3.565$ | $18.555 \pm 2.800$ |
| $15.0-20.0$ | $11.77 \pm 4.129$ | $13.73 \pm 3.124$ | $13.017 \pm 2.491$ |
| $20.0-30.0$ | $14.84 \pm 3.210$ | $10.28 \pm 2.383$ | $11.900 \pm 1.913$ |
| $30.0-50.0$ | $8.567 \pm 2.222$ | $8.911 \pm 2.367$ | $8.728 \pm 1.620$ |
| $50.0-100.0$ | $4.660 \pm 2.186$ | $6.753 \pm 2.768$ | $5.464 \pm 1.716$ |
| $100.0-300.0$ | $3.301 \pm 2.493$ | $3.829 \pm 3.606$ | $3.472 \pm 2.051$ |
| $2.5-300.0$ | $11.97 \pm 1.324$ | $11.6 \pm 1.226$ | $11.771 \pm 0.900$ |

Table 6.83: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of Evis calculated from a fixed BN.

| Evis(GeV) | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocN $)$ | $\langle\mathrm{R}\rangle$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $13.1 \pm 6.253$ | $2.363 \pm 8.942$ | $9.574 \pm 5.124$ |
| $8.0-15.0$ | $21.22 \pm 4.443$ | $15.97 \pm 3.494$ | $17.976 \pm 2.746$ |
| $15.0-20.0$ | $12.12 \pm 3.85$ | $13.91 \pm 3.064$ | $13.216 \pm 2.397$ |
| $20.0-30.0$ | $14.91 \pm 3.177$ | $10.33 \pm 2.236$ | $11.847 \pm 1.829$ |
| $30.0-50.0$ | $8.265 \pm 2.145$ | $9.474 \pm 2.199$ | $8.854 \pm 1.535$ |
| $50.0-100.0$ | $5.53 \pm 1.708$ | $7.597 \pm 2.24$ | $6.290 \pm 1.358$ |
| $100.0-300.0$ | $3.412 \pm 2.634$ | $4.655 \pm 3.52$ | $3.858 \pm 2.109$ |
| $2.5-300.0$ | $11.97 \pm 1.324$ | $11.6 \pm 1.226$ | $11.771 \pm 0.900$ |



Figure 6.84: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ using variable BN.


Figure 6.85: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ calculated from a fixed BN.


Figure 6.86: Ratio between $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ using variable BN .


Figure 6.87: Ratio between of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ calculated from a fixed BN.

Table 6.84: Comparison of $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocP and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ :FocN as a function of $E_{\nu}$ calculated from a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocP $)$ | $\mathrm{R}\left(\right.$ Coh $\pi^{-}:$FocN $)$ | $<\mathrm{R}\rangle$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $13.10 \pm 6.116$ | $3.215 \pm 8.418$ | $9.685 \pm 4.948$ |
| $8.0-15.0$ | $21.12 \pm 4.565$ | $15.93 \pm 3.609$ | $17.926 \pm 2.831$ |
| $15.0-20.0$ | $12.562 \pm 3.816$ | $13.86 \pm 3.049$ | $13.354 \pm 2.382$ |
| $20.0-30.0$ | $14.645 \pm 3.178$ | $10.51 \pm 2.281$ | $11.916 \pm 1.853$ |
| $30.0-50.0$ | $8.4 \pm 2.177$ | $9.348 \pm 2.186$ | $8.872 \pm 1.543$ |
| $50.0-100.0$ | $5.556 \pm 1.778$ | $7.556 \pm 2.314$ | $6.298 \pm 1.410$ |
| $100.0-300.0$ | $3.14 \pm 2.332$ | $4.147 \pm 3.005$ | $3.519 \pm 1.842$ |
| $2.5-300.0$ | $11.97 \pm 1.324$ | $11.6 \pm 1.226$ | $11.771 \pm 0.900$ |

Table 6.85: Ratio between $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ using variable BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ | $\mathrm{R}\left(\operatorname{Coh} \pi^{-}\right) / \mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $11.611 \pm 4.432$ | $7.82 \pm 1.230$ | $1.485 \pm 0.613$ |
| $8.0-15.0$ | $18.555 \pm 2.800$ | $8.33 \pm 0.446$ | $2.227 \pm 0.357$ |
| $15.0-20.0$ | $13.017 \pm 2.491$ | $7.21 \pm 0.423$ | $1.805 \pm 0.361$ |
| $20.0-30.0$ | $11.900 \pm 1.913$ | $5.65 \pm 0.303$ | $2.106 \pm 0.357$ |
| $30.0-50.0$ | $8.728 \pm 1.620$ | $4.64 \pm 0.272$ | $1.881 \pm 0.366$ |
| $50.0-100.0$ | $5.464 \pm 1.716$ | $3.32 \pm 0.256$ | $1.646 \pm 0.532$ |
| $100.0-300.0$ | $3.472 \pm 2.051$ | $1.76 \pm 0.342$ | $1.973 \pm 1.227$ |
| $2.5-300.0$ | $11.771 \pm 0.900$ | $5.24 \pm 0.134$ | $2.246 \pm 0.181$ |

Table 6.86: Ratio between $\mathrm{R}=\frac{\sigma\left(C o h \pi^{-}\right)}{\sigma\left(\bar{\nu}_{\mu} C C\right)}$ and $\mathrm{R}=\frac{\sigma\left(C o h \pi^{+}\right)}{\sigma\left(\nu_{\mu} C C\right)}$ as a function of $E_{\nu}$ calculated from a fixed BN.

| $E_{\nu}(\mathrm{GeV})$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ | $\mathrm{R}\left(\mathrm{Coh} \pi^{-}\right) / \mathrm{R}\left(\mathrm{Coh} \pi^{+}\right)$ |
| :--- | :---: | :---: | :---: |
| $2.5-8.0$ | $9.685 \pm 4.948$ | $7.82 \pm 1.230$ | $1.238 \pm 0.662$ |
| $8.0-15.0$ | $17.926 \pm 2.813$ | $8.33 \pm 0.446$ | $2.152 \pm 0.357$ |
| $15.0-20.0$ | $13.354 \pm 2.382$ | $7.21 \pm 0.423$ | $1.852 \pm 0.348$ |
| $20.0-30.0$ | $11.916 \pm 1.853$ | $5.69 \pm 0.298$ | $2.132 \pm 0.350$ |
| $30.0-50.0$ | $8.872 \pm 1.543$ | $4.64 \pm 0.268$ | $1.912 \pm 0.350$ |
| $50.0-100.0$ | $6.298 \pm 1.410$ | $3.24 \pm 0.255$ | $1.944 \pm 0.461$ |
| $100.0-300.0$ | $3.519 \pm 1.842$ | $1.86 \pm 0.297$ | $1.892 \pm 1.035$ |
| $2.5-300.0$ | $11.771 \pm 0.900$ | $5.24 \pm 0.134$ | $2.246 \pm 0.181$ |

Table 6.87: Variation of Selection Cuts.

| $\mathrm{E}(\mathrm{GeV})$ | $E_{\pi}-\sigma$ | $E_{\pi}+\sigma$ | $\theta_{\mu \pi}-\sigma$ | $\theta_{\mu \pi}+\sigma$ | $P_{T}^{m}-\sigma$ | $P_{T}^{m}+\sigma$ | $\|t\|-\sigma$ | $\|t\|+\sigma$ | $L H-\sigma$ | $L H+\sigma$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| $2.5-6.0$ | 11.6801 | 14.9215 | -0.0029 | 1.2152 | -0.6940 | 1.6924 | 3.1682 | 4.8313 | 7.9703 | 2.0819 |
| $6.0-8.0$ | -8.7922 | -13.2404 | 0.1911 | -2.2089 | -3.2015 | -0.1998 | -11.6918 | 7.8951 | -6.5333 | 3.0169 |
| $8.0-10.0$ | 2.9775 | -0.3931 | 0.6583 | -0.2867 | 1.3526 | 0.0433 | -6.7592 | 3.8665 | -0.2399 | 1.7187 |
| $10.0-12.0$ | 2.3871 | -3.4039 | -0.0359 | 0.6659 | -1.5546 | -0.1880 | -5.0496 | 9.1302 | -2.7272 | 1.0839 |
| $12.0-15.0$ | 0.1747 | -1.2485 | 2.1770 | 1.7885 | 0.7135 | 1.4232 | -3.2093 | 5.4010 | 0.8625 | 1.5466 |
| $15.0-20.0$ | 1.0206 | 0.7432 | 0.4749 | -0.3013 | -0.1227 | 0.0429 | -6.9257 | 3.4628 | -0.8809 | 0.8849 |
| $20.0-25.0$ | -0.5879 | 1.8357 | 1.1529 | 0.1288 | 0.9869 | 0.4498 | -2.6667 | 5.2470 | 0.5167 | 2.3634 |
| $25.0-30.0$ | 1.0098 | -0.3757 | 0.3817 | 0.9409 | 0.9814 | -0.1689 | -2.1256 | 5.3498 | 1.5725 | 1.5785 |
| $30.0-40.0$ | 0.8639 | -1.2536 | 0.4798 | -1.2473 | 0.1509 | -0.3995 | -6.4679 | 4.9091 | -0.2745 | 2.1057 |
| $40.0-50.0$ | -0.1103 | -1.1908 | 0.8992 | -1.5585 | -0.5193 | 0.8160 | -8.4046 | 5.0322 | 1.8468 | 2.2545 |
| $50.0-70.0$ | 0.5876 | 1.4756 | 0.6383 | -0.4734 | -0.7513 | -0.1182 | -9.1249 | 5.5142 | -2.7839 | 0.9751 |
| $70.0-100.0$ | -0.0111 | -1.6245 | 1.9125 | -0.3689 | 3.0378 | -1.6033 | -10.5984 | 7.0058 | -6.2656 | 2.6411 |
| $100.0-130.0$ | 3.1054 | -1.0962 | -0.4299 | -1.6977 | -0.2720 | -0.0633 | -21.2723 | -3.5256 | -2.0776 | 3.8175 |
| $130.0-300.0$ | -0.4959 | -3.8586 | -0.9748 | 6.6181 | -0.8720 | -1.7911 | 3.5501 | -0.8281 | 2.5231 | 6.4279 |
| $2.5-300.0$ | 0.6386 | -0.6010 | 0.7517 | -0.0230 | 0.2123 | 0.0910 | -5.8974 | 5.0962 | -0.6228 | 0.3345 |

## Variation of Selection Cuts

Table 6.87 shows the variation of the preslection variable cuts of coherent $\pi^{+}$in Neutrino Beam Mode Data analysis. For the variable $E_{\pi}, \theta_{\mu \pi}, P_{T}^{m}$, and the slope of t , change the parameters that we care about by $\pm \sigma$, and get the fully corrected signal as $\mathrm{N}^{\prime}$, assuming the nominal is N , the error is $\left(N^{\prime}-N\right) / N$.

## Variation of Neutrino and Anti-neutrino Flux

As the description in NOMAD flux, the neutrinos in the beam originate from the decay of mesons produced through four different mechanisms: proton-Be interactions in the target, proton interactions downstream of the target in material other than beryllium, re-interactions of particles in the target and interactions of particles downstream of the target. These four sources all contribute the systematic uncertainties which are described in paper [8].


Figure 6.88: Total energy-dependent uncertainties on the yields of each of the neutrino $\operatorname{species}\left(\nu_{\mu}\right.$ and $\left.\bar{\nu}_{\mu}\right)$.

Figure 6.88 shows the energy-dependent uncertainties of neutrino and anti-neutrino flux [8].

## Cross-section Uncertainties

Table 6.88: The error of cross-section of charged current resonance and coherent $\rho$ events in 14 bins.

| Evis $(\mathrm{GeV})$ | CCres- $\sigma$ | CCRes $+\sigma$ | CCCoh $\rho-\sigma$ | CCCoh $\rho-\sigma$ |
| :--- | :---: | :---: | :---: | :---: |
| $2.5-6.0$ | 7.3905 | -1.5665 | 1.4230 | -1.0336 |
| $6.0-8.0$ | -4.7827 | -2.9778 | -5.8231 | 1.2684 |
| $8.0-10.0$ | 0.6041 | -1.4013 | 1.3472 | 1.6682 |
| $10.0-12.0$ | -2.4972 | 0.8300 | -1.4289 | -0.1821 |
| $12.0-15.0$ | 0.9553 | 1.1081 | -0.6609 | 1.0299 |
| $15.0-20.0$ | -0.9298 | -0.5188 | -1.7648 | 0.6175 |
| $20.0-25.0$ | 0.4320 | 2.4922 | -1.5011 | 2.6735 |
| $25.0-30.0$ | 2.2773 | 1.2459 | -0.5308 | 3.9376 |
| $30.0-40.0$ | -0.0182 | -1.8955 | -2.3173 | 1.2494 |
| $40.0-50.0$ | 1.7604 | 2.7377 | -3.1267 | 3.4498 |
| $50.0-70.0$ | -2.9851 | -0.0075 | -2.4287 | 2.3243 |
| $70.0-100.0$ | -6.4030 | 2.6083 | -2.7508 | 0.7109 |
| $100.0-130.0$ | -2.2339 | -5.3554 | -6.4215 | 1.7497 |
| $130.0-300.0$ | 2.4863 | 9.4299 | -1.1441 | 8.1019 |
| $2.5-300.0$ | -0.3841 | 0.4714 | -1.6963 | 1.6761 |

Table 6.88 shows the uncertainties originated from the calculation of cross-sections of charged current resonance and coherent events. Similar to the calculation of the variation of selection cuts, the error of cross-section is also calculated by changing the parameter by $\pm \sigma$, and get the new fully corrected signal as $\mathrm{N}^{\prime}$, assuming the nominal is N , then the error is $\left(N^{\prime}-N\right) / N$. The $\sigma$ of charged current resonance is $7 \%$ which is from the NOMAD measurement [46]. The $\sigma$ of charged current coherent $\rho^{+}$is $8 \%$ which is also from the NOMAD measurement [45]. In this table, we only listed the charged current and resonance and coherent $\rho^{+}$processes error, compared to these two, the quasi-elastic is negligible, we use control region to reweight essentially DIS. Therefore only these two need to be considered.

Table 6.89: Summary of experimental measurements of coherent $\pi^{-}$production in $\bar{\nu}_{\mu}$ CC interactions.

| Experiment | Pub. <br> Year | $E_{\nu}(\mathrm{GeV})$ | $\sigma\left(\mathrm{Coh}-\pi^{-}\right)$ <br> $\left(10^{-40} \mathrm{~cm}^{2} /\right.$ nucleus $)$ | $\frac{\left.\sigma\left(\mathrm{Coh} \pi^{2}\right)\right)}{\sigma C C}$ <br> $\times 10^{-3}$ |
| :--- | :---: | :---: | :---: | :---: |
| BEBC | 1986 | $40(5-200)$ | $175 \pm 25$ | $9 \pm 1$ |
| SKAT | 1986 | $7(3-20)$ | $113 \pm 35$ | $18 \pm 5$ |
| FNAL 15'BC | 1989 | $70(40-300)$ | $270 \pm 110$ | $5.7 \pm 2.2$ |
| CHARM II | 1993 | 19.1 | $139 \pm 40(\mathrm{RS})$ <br> $132 \pm 32(\mathrm{BS})$ |  |
| NOMAD | 2015 | $16.5(2.5-300)$ |  | $9.865 \pm 0.730$ |

### 6.9 Comparison with Previous Measurements

In Table 6.89, there is a summary of experimental measurements including the result that I measured from NOMAD data. We could see that this is the first measurement of coherent $\pi^{-}$production after over 22 years and gives the best measurement to date.

## Chapter 7

## Coherent Rho and Absolute Flux Measurement

### 7.1 Neutrino Induced Coherent $\rho^{0} \& \rho^{+}$

The research on coherent $\rho$ is very important in understanding of coherent processes, because the vector current is assumed to be dominated by the $\rho$ meson $\left(J^{P}=1^{-}\right)$, whereas the axial current is dominated by the $a_{1}$ meson $\left(J^{P}=1^{+}\right)$. Different from the coherent $\rho^{+}$, much of information of the neutral current coherent $\rho^{0}$ events is difficult to measure directly, for example, the momentum transfer $Q^{2}$ and variables related to it. On the theoretical side, there is a certain connection between the coherent $\rho^{0}$ and coherent $\rho^{+}$or coherent $\rho^{-}$. In Chapter 3, the cross-section of coherent $\rho^{ \pm}$and coherent $\rho^{0}$ have been calculated.

### 7.2 Photo-production of Coherent $\rho^{0}$

Beside the neutrino induced coherent $\rho^{+}$process, photo-production of the coherent $\rho^{0}$ process can also be used to predict the information of neutrino induced coherent $\rho^{0}$, and has the following advantages compared to neutrino induced coherent $\rho^{+}$. First, the momentum transfer squared $Q^{2}$ and other related kinematic variables could be calculated; Second, different from the neutrinos, the incoming flux of electrons is easy to detect or measure as compared to neutrinos. In this section, the cross-section of photo-production of coherent $\rho^{0}$ is calculated using the Vector Dominance Model similar as the calculation of neutrino induced coherent $\rho$ events. The process is shown in Figure 7.1 diagrammatically.


Figure 7.1: Feynman diagram of photon induced coherent $\rho$ process.

Using Feynman Rules, we have

$$
\begin{align*}
\mathcal{M}= & -i e \bar{u}_{e f} \gamma^{\mu} u_{e i} \frac{i g_{\mu \nu}}{q^{2}}<0\left|j_{e m}^{\nu}\right| \rho_{j}^{0}>(-i e) \\
& \mathcal{A}_{j}\left(\rho^{0} \alpha \rightarrow \beta\right) \times \frac{i}{q^{2}-m_{\rho}^{2}} \\
= & -e^{2} \bar{u}_{e f} \gamma^{\mu} u_{e i} \frac{g_{\mu \nu}}{q^{2}}<0\left|j_{e m}^{\nu}\right| \rho_{j}^{0}>\mathcal{A}_{j}\left(\rho^{0} \alpha \rightarrow \beta\right) \times \frac{i}{q^{2}-m_{\rho}^{2}} \\
= & e^{2} \bar{u}_{e f} \gamma^{\mu} u_{e i} \frac{g_{\mu \nu}}{Q^{2}}<0\left|j_{e m}^{\nu}\right| \rho_{j}^{0}>\mathcal{A}_{j}\left(\rho^{0} \alpha \rightarrow \beta\right) \times \frac{i}{Q^{2}+m_{\rho}^{2}}, \tag{7.1}
\end{align*}
$$

where $j$ is polarization index number. For the decay constant part, we may write

$$
\begin{align*}
<0\left|j_{e m}^{\nu}\right| \rho_{j}^{0}> & =<0\left|\bar{u} \gamma^{\nu} u q_{u}+\bar{d} \gamma^{\nu} d q_{d}\right| \frac{1}{\sqrt{2}}(u \bar{u}-d \bar{d})> \\
& =\frac{1}{\sqrt{2}}\left(<0\left|\bar{u} \gamma^{\nu} u q_{u}\right| u \bar{u}>-<0\left|\bar{d} \gamma^{\nu} d q_{d}\right| d \bar{d}>\right) \\
& =\frac{1}{\sqrt{2}} f_{\rho} m_{\rho} \epsilon_{j}^{\nu} \tag{7.2}
\end{align*}
$$

where the flavor structure of $\rho^{0}$ has been used:

$$
\begin{equation*}
\rho^{0}=\frac{1}{\sqrt{2}}(u \bar{u}-d \bar{d}) . \tag{7.3}
\end{equation*}
$$

$q_{u, d}$ are electrical charges of $u, d$ quarks respectively. Eventually, we obtain

$$
\begin{align*}
\mathcal{M} & =e^{2} \bar{u}_{e f} \gamma^{\mu} u_{e i} \frac{g_{\mu \nu}}{Q^{2}} \frac{1}{\sqrt{2}} f_{\rho} m_{\rho} \epsilon_{j}^{\nu} \mathcal{A}_{j}\left(\rho^{0} \alpha \rightarrow \beta\right) \times \frac{i}{Q^{2}+m_{\rho}^{2}} \\
& =\frac{e^{2}}{\sqrt{2}} f_{\rho} m_{\rho} 2 \bar{u}_{e f} \gamma^{\mu} u_{e i} \epsilon_{j}^{\mu} \mathcal{A}_{j}\left(\rho^{0} \alpha \rightarrow \beta\right) \times \frac{i}{Q^{2}+m_{\rho}^{2}} \tag{7.4}
\end{align*}
$$

Compare this to the scattering amplitude with the neutrino induced coherent $\rho^{0}$, we might come to the conclusion that, the differences between the two come from three parts:

- Electromagnetic interaction vertex $\frac{e^{2}}{\sqrt{2}}=\frac{4 \pi \alpha}{\sqrt{2}}$; Weak interaction vertex $\frac{G_{F}}{2}(1-$ $2 \sin ^{2} \theta_{W}$ )
- Photon propagator $\sim \frac{1}{Q^{2}} ; Z^{0}$ propagator $\sim \frac{1}{Q^{2}+m_{Z}^{2}}$ in weak interaction $m_{Z}^{2} \gg Q^{2}$ then, we have $\frac{1}{Q^{2}+m_{Z}^{2}} \sim \frac{1}{m_{Z}^{2}}$, already included in $G_{F}$;
- The leptronic tensor:

For Electromagnetic interaction

$$
\begin{align*}
\operatorname{Tr}\left\{\bar{u}_{e f} \gamma^{\mu} u_{e i} \bar{u}_{e i} \gamma^{\alpha} u_{e f}\right\} & =\operatorname{Tr}\left\{u_{e f} \bar{u}_{e f} \gamma^{\mu} u_{e i} \bar{u}_{e i} \gamma^{\alpha}\right\} \\
& =\frac{1}{2} \operatorname{Tr}\left\{\left(P_{f}+m_{e}\right) \gamma^{\mu}\left(\not P_{i}+m_{e}\right) \gamma^{\alpha}\right\} \\
& =\frac{1}{2} \operatorname{Tr}\left\{P_{f} \gamma^{\mu} \not P_{i} \gamma^{\alpha}+m_{e}^{2} \gamma^{\mu} \gamma^{\alpha}\right\} . \tag{7.5}
\end{align*}
$$

Since $m_{e}^{2}$ is very small, after ignoring it, we obtain

$$
\begin{align*}
\operatorname{Tr}\left\{\bar{u}_{e f} \gamma^{\mu} u_{e i} \bar{u}_{e i} \gamma^{\alpha} u_{e f}\right\} & =\frac{1}{2} \operatorname{Tr}\left\{P_{f} \gamma^{\mu} \not P_{i} \gamma^{\alpha}\right\} \\
& =\frac{1}{2} \times 4\left(P_{f}^{\mu} P_{i}^{\alpha}-P_{f} \cdot P_{i} g^{\mu \alpha}+P_{f}^{\alpha} P_{i}^{\mu}\right) \\
& =2\left(P_{f}^{\mu} P_{i}^{\alpha}-P_{f} \cdot P_{i} g^{\mu \alpha}+P_{f}^{\alpha} P_{i}^{\mu}\right) \tag{7.6}
\end{align*}
$$

Combine all the three terms together, we have:

$$
\begin{align*}
\frac{d \sigma\left(e+N \rightarrow e+\rho^{0}+N\right)}{d \sigma\left(\nu+N \rightarrow \nu+\rho^{0}+N\right)} & =\frac{\left(\frac{4 \pi \alpha}{\sqrt{2}}\right)^{2}\left(\frac{1}{Q^{2}}\right)^{2} 2^{2}}{\frac{G_{F}}{2}\left(1-2 \sin ^{2} \theta_{W}\right)^{2} \times 8^{2}} \\
& =\frac{32 \pi^{2} \alpha^{2}}{Q^{4} \times 16 \times G_{F}^{2}\left(1-2 \sin ^{2} \theta_{W}\right)^{2}} \\
& =\frac{2 \pi^{2} \alpha^{2}}{G_{F}^{2}\left(1-2 \sin ^{2} \theta_{W}\right)^{2} \times Q^{4}} \tag{7.7}
\end{align*}
$$

### 7.3 Connection Between the Neutrino- \& Photo-production of Coh--Rho: Absolute Flux

From the ratio of the cross-sections of photon and neutrino induced coherent $\rho^{0}$ processes, we have

$$
\begin{align*}
\frac{d \sigma\left(e+N \rightarrow e+\rho^{0}+N\right)}{d Q^{2} d \nu d t}= & \frac{2 \pi^{2} \alpha^{2}}{G_{F}^{2}\left(1-2 \sin ^{2} \theta_{W}\right)^{2} Q^{4}} \\
& \times \frac{d \sigma\left(\nu+N \rightarrow \nu+\rho^{0}+N\right)}{d Q^{2} d \nu d t} \tag{7.8}
\end{align*}
$$

The factor $Q^{4}$ in the denominator implies the cross-section would be infinitely large when $Q^{2} \rightarrow 0$, and the integral over $Q^{2}$ would be divergent. Ignoring the lepton's mass, we would have $Q^{2} \simeq 4 E E^{\prime} \sin ^{2} \frac{\theta}{2}$, where E and E ' are incoming lepton and outgoing lepton energies. We can see that the infinity appears at $\theta=0$. This is a property of Coulomb Scattering. We could divide the integral into two parts:

$$
\begin{equation*}
\int_{0}^{a} \frac{d \sigma}{d Q^{2} d \nu d t}=\int_{0}^{\epsilon} \frac{d \sigma}{d Q^{2} d \nu d t}+\int_{\epsilon}^{a} \frac{d \sigma}{d Q^{2} d \nu d t} \tag{7.9}
\end{equation*}
$$

The first term is divergent, and the second term is finite. This means that the number of events with $Q^{2}$ between 0 and $\epsilon$ is too large compared to the events with $Q^{2}$ between $\epsilon$ and a. Practically, we could introduce a cut-off to avoid divergence.

It is easy to deduce the relation between the photon induced process and the neutrino induced coherent $\rho^{ \pm}$processes from Equation (3.117) and Equation (7.8):

$$
\begin{equation*}
\frac{d \sigma\left(e+N \rightarrow e+\rho^{0}+N\right)}{d Q^{2} d \nu d t}=\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}} \frac{d \sigma\left(\nu+N \rightarrow \mu+\rho^{+}+N\right)}{d Q^{2} d \nu d t} \tag{7.10}
\end{equation*}
$$

### 7.4 Simulation of Coherent $\rho^{+}$Production

The simulation of the coherent $\rho^{+}$events is based on the Neglib package which is used to simulate the Monte Carlo events of coherent $\pi^{+}$using NOMAD flux. The procedure includes:

- Set Input Variables: Instead of using the NOMAD flux, LBNF flux was used as the incoming flux. Since the design of HIRESMNU is based on the experience of NOMAD detector, the target used in the simulation of Coherent $\rho^{+}$is the same as in the Neglib package. Because the output of the StandAlone Code will be used to GENIE(Generates Events for Neutrino Interaction Experiments), all the codes are written in $\mathrm{C}++$. The mass of the coherent meson ( $\rho^{+}$) was generated randomly according to the distribution of relativistic breit wigner formula which is:

$$
\begin{equation*}
f(E)=\frac{k}{\left(E^{2}-M^{2}\right)^{2}+M^{2} \Gamma^{2}} \tag{7.11}
\end{equation*}
$$

where

$$
\begin{align*}
k & =\frac{2 \sqrt{2} M \Gamma \gamma}{\pi^{2} \sqrt{M^{2}+\gamma}} \\
\gamma & =\sqrt{M^{2}\left(M^{2}+\Gamma^{2}\right)} \tag{7.12}
\end{align*}
$$

Xbj and Ybj are generated randomly from 0 to 1 . The lepton energy(Elep) is calculated from the $E_{\nu}$ and $\mathrm{Ybj}\left(\operatorname{Elep}=E_{\nu} \times(1-\mathrm{Ybj})\right)$. The $|\mathrm{t}|$ is also calculated from a random number, which is

$$
\begin{equation*}
|t|=-\frac{1}{b} \times \ln (\text { Rand }) \tag{7.13}
\end{equation*}
$$

where the slope parameter $b$ is calculated according to Equation(9) in Rein and Sehgal's paper[44]. Rand represents a random number from 0 to 1. After setting the input values of $E_{\nu}, \mathrm{Xbj}, \mathrm{Ybj}$, mass of meson $m_{\rho}$, some other variables like $\nu, Q^{2}, W^{2}$ could be calculated from these three variables.

- Perform a Number of Kinematic boundaries: After generating the four input variables, we then apply some kinematic boundary cuts to select the physical events, including, maximum of Xbj , maximum of Ybj , minimum of Ybj , maximum of t , minimum of t and so on which are consistent to the values in Neglib package.
- Cross-Section Calculation: the cross-section of coherent $\rho^{+}$could be calculated from the variables generated and calculated for each event.


## - Reweight Distribution to Get Photo-production Coherent $\rho^{0}$ Events

 Using the result of Equation (7.7), reweighted all the variable distributions of Coherent $\rho^{+}$, we then obtain the distributions of all the corresponding variables in photo-production coherent $\rho^{+}$events.To get the $Q^{2}$ distribution of $e+\mathcal{A} \rightarrow e+\mathcal{A}+\rho^{0}$, we just need to reweight the corresponding distribution of neutrino induced coherent $\rho^{+}$process by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$. However, for other variables, we could not do this simply. What we can do is, suppose, we want to get the distribution of variable X: First, we need to create the 2-dimensional distribution of $X$ and $Q^{2}$, then reweight each bin by this factor; Second, integrate over $Q^{2}$, then we get the distribution of variable $X$.

Since the ratio of the cross-section between photon induced and neutrino induced coherent processes is a function of $Q^{2}$, it is reasonable to consider $Q^{2}$ first. To avoid the condition that $Q^{2}=0$, I applied a lower cut value of $Q^{2}$, which is $0.02 \mathrm{GeV}^{2}$ (we could also set this cut value from experimental experience).

The analysis was done with the $\mathrm{C}++$ standAlone code I wrote based on the Neglib package. Figure 7.2 is the distribution of $Q^{2}$ of neutrino induced coherent $\rho^{+}$events vs photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$. Similar plots for other important variables are shown from Figure 7.3 to Figure 7.7.


Figure 7.2: $\quad Q^{2}$ distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.


Figure 7.3: Xbj distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.


Figure 7.4: Ybj distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.


Figure 7.5: $\quad P_{T}^{m}$ distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.


Figure 7.6: $\quad \zeta_{\rho}$ distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.


Figure 7.7: t distribution of neutrino induced coherent $\rho^{+}$events generated from LBNE flux and photon induced coherent events obtained from $\nu$-induced coherent $\rho^{+}$ events reweighted by factor $\frac{\pi^{2} \alpha^{2}}{G_{F}^{2} Q^{4}}$.

## Chapter 8

## Summary and Future Work

Using the CVC, PCAC hypothesis, and Hadron Dominance Model, I calculated the the neutrino induced coherent $\pi$, and coherent $\rho$. I also calculated the cross-section of photon induce coherent $\rho^{0}$ process and the ratio between the photon induced coherent $\rho^{0}$ and neutrino induced coherent $\rho$ which gives a way to get additional information to constrain the neutrino fluxes. Beside the theoretical calculations, I measured the ratio between cross-section of coherent $\pi^{-}$and $\bar{\nu}_{\mu}$ CC interactions with the NOMAD data including neutrino beam mode data and anti-neutrino beam mode data, and compared it to the measurement of coherent $\pi^{+}$production. This is the best measurement of coherent $\pi^{-}$to date. Based upon the experience of analysis with NOMAD data, my final aim is to evaluate the sensitivity of ELBNF/DUNE project to coherent process. Then, in the second part, I used new C++ standAlone Code I wrote and simulated coherent $\rho^{+}$interactions. After that, I reweighted the distributions of some kinematic variables with the factor (ratio between cross-section of photon induced and neutrino induced coherent $\rho$ process). I calculated and get the corresponding distributions of photon induced coherent process.

In the future, I will integrate the $\mathrm{C}++$ simulation package I wrote into the GEINE neutrino event generator which is written in C++. GENIE is a comprehensive neutrino Monte Carlo generator supported and developed by an international collaboration of scientists [6]. The GENIE model is universal. It handles all neutrinos and nuclear targets, and all processes relevant from MeV to PeV energy scales. Some experiments(including T2K, No $\nu \mathrm{A}, \mathrm{MINER} \nu \mathrm{A}$, MicroBooNE, LAr1-ND, ELBNF and

IceCUBE) use GENIE whose predictions is a standard reference point for the entire community.

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## Appendix A

## CP VIOLATION

$$
\begin{equation*}
\nu^{C P}=\gamma^{0} \mathcal{C} \bar{\nu}^{T}=-\mathcal{C} \nu^{*} \tag{A.1}
\end{equation*}
$$

C: Particle $\rightleftharpoons$ Antiparticle P: Left-Handed $\rightleftharpoons$ Right-handed Neutrino fields under CP transformation is [17]:

$$
\begin{align*}
& \nu_{\alpha L}=\sum_{k} U_{\alpha k} \nu_{k L} \xrightarrow{C P} \nu_{\alpha L}^{C P}=\sum_{k} U_{\alpha k}^{*} \nu_{k l}^{C P}  \tag{A.2}\\
& \left|\nu_{\alpha}\right\rangle=\sum_{k} U_{\alpha k}^{*}\left|\nu_{k}\right\rangle \xrightarrow{C P}\left|\bar{\nu}_{\alpha}\right\rangle=\sum_{k} U_{\alpha k}\left|\bar{\nu}_{k}\right\rangle \tag{A.3}
\end{align*}
$$

Neutrinos $U \rightleftharpoons U *$ Antineutrinos

$$
\begin{align*}
& P_{\nu_{\alpha} \rightarrow \nu_{\beta}}(L, E)=\sum_{k}\left|U_{\alpha k}\right|^{2}\left|U_{\beta k}\right|^{2}+2 R e \sum_{k>j} U_{\alpha k}^{*} U_{\beta k} U_{\alpha j} U_{\beta j}^{*} \exp \left(-i \frac{\Delta m_{k j}^{2} L}{2 E}\right)  \tag{A.4}\\
& P_{\bar{\nu}_{\alpha} \rightarrow \bar{\nu}_{\beta}}(L, E)=\sum_{k}\left|U_{\alpha k}\right|^{2}\left|U_{\beta k}\right|^{2}+2 R e \sum_{k>j} U_{\alpha k} U_{\beta k}^{*} U_{\alpha j}^{*} U_{\beta j} \exp \left(-i \frac{\Delta m_{k j}^{2} L}{2 E}\right) \tag{A.5}
\end{align*}
$$

PMNS(Pontecorvo-Maki-Nakagawa-SaKata) Neutrino Mixing Matrix is a unitary matrix which cantains information on the mismatch of quantum states of leptons when they propagate freely and when they take part in the weak interactions. It is important in the understanding of neutrino oscillations. For three generations of
leptons, the matrix can be written as [17]:

$$
\left.\begin{array}{rl}
U= & R_{23} W_{13} R_{12} D(\lambda) \\
= & \left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & c_{23} & s_{23} \\
0 & -s_{23} & c_{23}
\end{array}\right)\left(\begin{array}{ccc}
c_{13} & 0 & s_{13} e^{-i \delta_{13}} \\
0 & 1 & 0 \\
-s_{13} e^{i \delta_{13}} & 0 & c_{13}
\end{array}\right)\left(\begin{array}{ccc}
c_{12} & s_{12} & 0 \\
-s_{12} & c_{13} & 0 \\
0 & 0 & 1
\end{array}\right) \\
& \times\left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & e^{i \lambda_{21}} & 0 \\
0 & 0 & e^{i \lambda_{31}}
\end{array}\right) \\
= & \left(\begin{array}{ll}
-s_{12} c_{13} \\
s_{23}-c_{12} s_{23} e^{0 \delta_{13} i} & c_{12} c_{23}-s_{12} s_{23} s_{13} e^{i \delta_{13}} \\
s_{12} c_{23} s_{13} e^{i \delta_{13}} & -c_{12} s_{23}-s_{12} c_{23} s_{13} e^{i \delta_{13}} \\
c_{23} c_{13}
\end{array}\right) \\
& \times\left(\begin{array}{cc}
1 & 0 \\
0 \\
0 & e^{i \lambda_{21}} \\
0 \\
0 & 0
\end{array}\right) e^{i \lambda_{31}} \tag{A.6}
\end{array}\right)
$$

PMNS Neutrino Mixing Matrix Then, $P_{\nu_{\mu} \rightarrow \nu_{e}}-P_{\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}}$ could be used as a direct test of CP symmetry,

$$
\begin{align*}
P_{\nu_{\mu} \rightarrow \nu_{e}}-P_{\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}}= & 2 R e \sum_{k>j}\left[U_{\alpha k}^{*} U_{\beta k} U_{\alpha j} U_{\beta j}^{*}-U_{\alpha k} U_{\beta k}^{*} U_{\alpha j}^{*} U_{\beta j}\right] \exp \left(-i \frac{\Delta m_{k j}^{2} L}{2 E}\right) \\
& 4 R e \sum_{k>j} i \Im\left[U_{\alpha k}^{*} U_{\beta k} U_{\alpha j} U_{\beta j}^{*}\right] \exp \left(-i \frac{\Delta m_{k j}^{2} L}{2 E}\right) \\
= & 4 \sum_{k>j}\left[U_{\alpha k}^{*} U_{\beta k} U_{\alpha j} U_{\beta j}^{*}\right] \sin \left(\frac{\Delta m_{k j}^{2} L}{2 E}\right) \tag{A.7}
\end{align*}
$$

where $J=\Im\left[U_{\alpha k}^{*} U_{\beta k} U_{\alpha k} U_{\beta j}^{*}\right]$ is is called a Jarlskog invariant, from the quark CKM
unitary triangle

$$
\begin{align*}
J= & c_{12} s_{12} c 23 s_{23} c_{13}^{2} s_{13} \sin \delta \\
= & \frac{1}{2} \sin 2 \theta_{12} \sin 2 \theta_{23} \sin 2 \theta_{13} \cos \theta 13 \sin \delta \\
= & 4 \sum_{k>j} \Im\left[U_{\alpha k}^{*} U_{\beta k} U_{\alpha j} U_{\beta j}^{*}\right] \sin \left(\frac{\Delta m_{k j}^{2} L}{2 E}\right) \\
= & \sin \delta \sin 2 \theta_{12} \sin 2 \theta_{23} \sin \theta_{13} \cos ^{2} \theta_{13} \sin \left(\frac{\Delta m_{31}^{2} L}{2 E}\right) \\
& -\sin \delta \sin 2 \theta_{12} \sin 2 \theta_{23} \sin \theta_{13} \cos ^{2} \theta_{13} \sin \left(\frac{\Delta m_{32}^{2} L}{2 E}\right) \\
& -\sin \delta \sin 2 \theta_{12} \sin 2 \theta_{23} \sin \theta_{13} \cos ^{2} \theta_{13} \sin \left(\frac{\Delta m_{21}^{2} L}{2 E}\right) \tag{A.8}
\end{align*}
$$

From Equation (A.8) we could know, that if any angle is zero, $\Delta P=0$; if any $\Delta m^{2}=0, \Delta P=0 ;$

$$
\begin{equation*}
A_{\alpha \beta}^{C P}=4 J \sum_{k>j} s_{\alpha \beta ; k j} \sin \left(\frac{\Delta m_{k j}^{2} L}{2 E}\right) \tag{A.9}
\end{equation*}
$$

- s for 31 is -1
- s for 32 is +1
- s for 21 is +1
where, we could see that for $e \rightarrow \mu, \mu \rightarrow \tau, \tau \rightarrow e, \mathrm{~s}$ is positive and the sign flips if flavors flip.

From the equation above, we could see that to determine the CP violation, first, all angles should be large; second, the detector should be placed at correct L(baseline) for neutrino beam E; third, $\Delta m_{21}^{2}$ is small compared to $\Delta m_{31}^{2}$. The normal and inverted hierarchy are

- Normal Hierarchy(NH): $\Delta m_{31}^{2}=\Delta m_{21}^{2}+\Delta m_{32}^{2}$,
- Inverted Hierarchy(IH): $\left|\Delta m_{31}^{2}\right|=\Delta m_{21}^{2}+\left|\Delta m_{32}^{2}\right|$.

