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الجـامعــــــــة الإســــلاميــة - غـزة شئون الـبحث العلمي والاراسات العليا


# Methods in The Treatment of Singular Lagrangian 


by

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أنا الموقع أدناه مقدم الرسالة التي تحمل العنوان:

# Methods in The Treatment of Singular Lagrangian 

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## نتيجة المكم على أطروحة هاجستير

بناءً على موافقة عمـادة البحث العلمي والدراسات العليا بالجامعة الإسـلامية بغزة على تشكيل لجنة الحكم
 الفبزيـــاء وموضوعها:
طرق في معالجات اللاجر انـج الآحادي

## Methods in The Treatments of Singular Lagrangian

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To My Parents, wife, daughter"Tala" Familly,,, to

Those who tought me


#### Abstract

In this thesis, the singular Lagrangian is studied by three approaches. The Hamiltonian formalism is treated using both Dirac's method and Güler method (HamiltonJacobi method ). The third approach is to treat the singular Lagrangian as field (or continuous) system.

In Dirac's method, one introduces a primary constraints to the first - class constraints which have vanishing poisson brackets. The equation of motion are obtained as total derivatives in terms of poisson brackets.

In Hamilton-Jacobi formulation, which developed by Güler, the equations of motion are written as total differential equations in many variables. These equations must satisfy the integrability conditions.

The third approach is the treating of the singular Lagrangian as field (or continuous) system, We mixed both Lagrangian formulation and Hamilton-Jacobi method to obtain a solvable partial differential equation of second order. The solution of these equations are obtained easily. These solution satisfied the equations of motion.

In these three approach, the equation of motion are built for several physical models and integrability conditions of these equations of motion are discussed. A comparison between the results of these approaches is done and it is shown that the results are the same.


## ملخص

تعتمد هذه الاطروحة علي دراسة اللاجر انج الاحادي من خلال ثلاث طرق. في البداية تم معالجة صيغة الهاميلنون باستخدام طريقتين دير الك وجولر ، اما الطريقة الثالثه فتتمثلّل بمعاملة اللاجر انج الأحادي لنظام متصل ( مجال ) .
في طريقة دير اك ، نستخدم أحد القيود الاساسية بحيث يتم تلاشي( أقواس بوسون)، ويتم الحصول علي معادلات الحركة من المشتقات الكليه لإفواس بوسون.

في طريقة هاميلتون ـ جاكوبي، التي طور ها العالم جولر، تكتب معادلات الحركة كمعادلات تفاضلية بمتغيرات متعددة و التي يجب ان تحقق شروط التكامل. في طريقة المجال المتصل، تم دمج طريقتى اللاجر انج وطريقة هاميلتون جاكوبي للحصول علي معادلات تفاضلية جزيئة من الدرجة الثانية، ويتم حل هذه المعادلات للحصول علي معادلات حركة.

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## Chapter 1

## Introduction

## Chapter 1

## Introduction

### 1.1 Historical Background

The Hamiltonian formulation of singular systems is usually made through the formalism developed by Dirac (1;2). In this formalism, the constraints caused by the singularity of Hess matrix $\quad \frac{\partial^{2} L}{\partial \dot{q}_{i} \partial \dot{q}_{j}}, \quad i, j=1, \ldots, n$, are added to the canonical Hamiltonian, and then the consistency conditions are worked out, being possible to eliminate some degrees of freedom of the system. Dirac also showed that the gauge freedom is caused by the presence of first class constraints. This formalism has a wide range of applications in field theory and it is still the main tool for the analysis of singular systems (3; 5). Despite the success of Dirac's method, it is always interesting to apply different formalisms to the analysis of singular systems.

The study of another formalisms for singular lagrangians systems may provide new tools to investigate these systems. In classical dynamics, different formalisms (Lagrangian, Hamiltonian, Hamilton-Jacobi) provide different approaches to the problems, each formalism has advantages and disadvantages in the study of some features of the systems and being equivalent among themselves. In the same way,
different formalisms provide different views of the features of singular systems, which justify the interest in their study.

Also in this thesis, we generalize the Hamilton-Jacobi formalism that was developed by Güler $(6 ; 7)$. This approach based on Carathéodory's equivalent Lagrangian method (8) to write down the Hamilton-Jacobi equations for the system and make use of its singularity to write the equations of motion as total differential equations in many variables. The advantage of the Hamilton-Jacobi formalism is that we have no difference between the first and the second constraints and we do not need gauge-fixing term because the gauge variables are separated in the processes of constructing an integrable system of total differential equation.

In this work we will investigate several models using Dirac and Hamilton-Jacobi (Güler) approaches. Further more, we will treate these models using Lagrangian formalism as field systems. In the following three sections we will give brief review of these formalisms.

### 1.2 Singular Systems

The singular Lagrangian system represents a special case of a more general dynamics called constrained system (2). The dynamics of the physical system is encoded by the Lagrangian, a function of positions and velocities of all degrees of freedoms which comprise the system . The singular Lagrangian can be achieved by two formulations, the Lagrangian and the Hamiltonian formulations.

This section serves as an initiation to the concept of singularities in the Lagrange formalism. We will introduce some basic notions such as constraints arising due to the singularities and the definition of the canonical momenta. We will start our discussion of Singular Lagrangian systems with the principle of least action. Any
physical system can be described by a function $L$ depending on the positions and velocities (9)

$$
\begin{equation*}
L=L\left(q_{i}(t), \dot{q}_{i}(t)\right), \quad i=1, \ldots, n \tag{1.2.1}
\end{equation*}
$$

We assume, for the sake of simplicity, that this Lagrange function exhibits no explicit time dependence. The abbreviations $q(t)$ and $\dot{q}(t)$ stand for the set of all positions $q_{i}(t)=q_{i}(t)$ and velocities $\dot{q}_{i}(t)=\dot{q}_{i}(t)$, respectively, with $i=1, \ldots, n$. The system motion proceeds in a way that the action integral

$$
\begin{equation*}
A=\int_{t_{1}}^{t_{2}} d \tau L\left(q_{i}(t), \dot{q}_{i}(t)\right) \tag{1.2.2}
\end{equation*}
$$

becomes stationary under infinitesimal variations $\delta q_{i}(t)$. Assuming that the end points are fixed during the variation, i.e. $\delta q_{i}\left(t_{1}\right)=\delta q_{i}\left(t_{2}\right)=0$, yields the equations of motion for the classical path, which is called Euler-Lagrange equation

$$
\begin{equation*}
\frac{\partial L}{\partial q_{i}}-\frac{d}{d t} \frac{\partial L}{\partial \dot{q}_{i}}=0 \tag{1.2.3}
\end{equation*}
$$

Executing the total time derivative gives

$$
\begin{equation*}
\ddot{q}_{j} \frac{\partial^{2} L}{\partial \dot{q}_{i} \partial \dot{q}_{j}}=\frac{\partial L}{\partial q_{i}}-\dot{q}_{j} \frac{\partial^{2} L}{\partial q_{i} \partial q_{j}}, \tag{1.2.4}
\end{equation*}
$$

In this form we recognize that the accelerations $\ddot{q}_{i}$ can be uniquely expressed by the position and the velocities $\dot{q}_{i}$ if and only if the Hess matrix

$$
\begin{equation*}
w_{i j}=\frac{\partial^{2} L}{\partial \dot{q}_{i} \partial \dot{q}_{j}}, \quad i, j=1, \ldots, n \tag{1.2.5}
\end{equation*}
$$

is invertible. In other words its determinant must not vanish.

$$
\begin{equation*}
\operatorname{det} w_{i j} \neq 0 \tag{1.2.6}
\end{equation*}
$$

Since we are interested in the Hamiltonian formulation, we have to perform a Legendre transformation from the velocities to the momenta. The latter are defined as

$$
\begin{equation*}
p_{i}=\frac{\partial L}{\partial \dot{q}_{i}} \tag{1.2.7}
\end{equation*}
$$

In the case that the determinant vanishes, the Lagrangian (1.2.1) is singular and some of the accelerations are not determined by the velocities and positions. This means that some of the variables are not independent from each other. The singularity of the Hessian is equivalent to the noninvertibility of (1.2.5 ). As a consequence, in a singular system we are not able to display the velocities as functions of the momenta and the positions. This gives rise to the existence of relations between the positions and momenta

$$
\begin{equation*}
\phi_{m}\left(p_{i}, q_{i}\right)=0, \tag{1.2.8}
\end{equation*}
$$

these relations are called primary constraints in Dirac's approach. They follow directly from the structure of the Lagrangian and the definition of the momenta (1.2.7). The interesting point is that these functions are real restrictions on the phase space.

### 1.3 Dirac's Method

The standard methods of classical mechanics can't be applied directly to the singular Lagrangian theories. However, the basic idea of the classical treatment and the quantization of such systems were presented along time by Dirac (1; 2). And is now widely used in investigating the theoretical models in a contemporary elementary particle physics and applied in high energy physics, especially in the gauge theories(5). This is because the first-class constraints are generators of gauge transformation which lead to the gauge freedom (14) Let us consider a system which is described by the Lagrangian (1.2.1) is singular if the rank of the Hess matrix $A_{i j}=\frac{\partial^{2} L}{\partial \dot{q}_{i} \partial \dot{q}_{j}}, \quad i, j=1, \ldots, n$, is $r=n-m, m<n$.otherwise, the Lagrangian will be regular The singular system characterized by the fact that all velocities $\dot{q}_{i}$ are not uniquely determined in terms of the coordinates and momenta only. In other
words, not all momenta are independent, and there must exist a certain set of relations among them of the form (1.2.8) The generalized momenta corresponding to the generalized coordinates $q_{i}$ are defined as

$$
\begin{array}{cc}
p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}, & a=1, \ldots, n-r, \\
p_{\mu}=\frac{\partial L}{\partial \dot{q}_{\mu}}, & \mu=n-r+1, \ldots, n, \tag{1.3.2}
\end{array}
$$

where $\dot{q}_{i}$ stands for the total derivative with respect to t . The Relations (1.3.2) enable us to write the primary constraint as $(1 ; 2)$

$$
\begin{equation*}
H_{\mu}^{\prime}=P_{\mu}+H_{\mu}=0 \tag{1.3.3}
\end{equation*}
$$

In this formulation the total Hamiltonian is defined as

$$
\begin{equation*}
H_{T}=H_{0}+\lambda_{\mu} H_{\mu}^{\prime} \tag{1.3.4}
\end{equation*}
$$

where the canonical Hamiltonian $H_{0}$ is defined as

$$
\begin{equation*}
H_{0}=p_{i} \dot{q}_{i}-L, \quad i=1, \ldots, n \tag{1.3.5}
\end{equation*}
$$

and $\lambda_{\mu}$ are arbitrary functions. (Throughout this thesis, we use Einstein's summation rule which means that the repeating of indices indicate to summation ). The equations of motion are obtained in term of Poisson brackets as

$$
\begin{align*}
& \dot{q}_{i}=\left\{q_{i}, H_{T}\right\}=\left\{q_{i}, H_{\circ}\right\}+\lambda_{\mu}\left\{q_{i}, H_{\mu}^{\prime}\right\},  \tag{1.3.6}\\
& \dot{p}_{i}=\left\{p_{i}, H_{T}\right\}=\left\{p_{i}, H_{\circ}\right\}+\lambda_{\mu}\left\{p_{i}, H_{\mu}^{\prime}\right\} . \tag{1.3.7}
\end{align*}
$$

The consistency conditions, which means that the total time derivative of the primary constrains should be identically zero are given as

$$
\begin{equation*}
H_{\mu}^{\prime}=\left\{H_{\mu}^{\prime}, H_{T}\right\}=\left\{H_{\mu}^{\prime}, H_{\circ}\right\}+\lambda_{\mu}\left\{H_{\mu}^{\prime}, H_{\nu}^{\prime}\right\} \approx 0 \tag{1.3.8}
\end{equation*}
$$

where $\mu, \nu=1, \ldots, r$. Equations (1.3.6, 1.3.7, 1.3.8 ) may be identically satisfied for the singular system with primary constraints. These equations may be reduced to 0 $=0$, where it is identically satisfied as a result of primary constraints, else they will be lead to new conditions which are called secondary constraints. Repeating this procedure as many times as needed, one arrives at a final set of constraints or/and specifies some of $\lambda_{\mu}$. Such constraints are classified into two types, a) Firstclass constraints which have vanishing Poisson brackets with all other constraints. b) Second-class constraints which have nonvanishing Poisson brackets. The secondclass constraints could be used to eliminate conjugated pairs of the $p$ 's and $q$ 's from the theory by expressing them as functions of the remaining $p{ }^{\prime} s$ and $q ' s$. The total Hamiltonian for the remaining variable is then the canonical Hamiltonian plus the primary constraints $H_{\mu}^{\prime}$ of the first type as in Eq. (1.3.4), where $H_{\mu}^{\prime}$ are all the independent remaining first-class constraints.

### 1.4 Hamilton-Jacobi Approach (Güler Method)

The aim is to obtain a valid and consistent Hamilton-Jacobi theory of singular systems. The main point of the method is to define the equivalent Lagrangian (variational principle) and then pass to the phase space. This formulation leads us to a set of Hamilton-Jacobi partial differential equation (6), (7) and (8).

### 1.4.1 Construction of Phase Space

The starting point of the Hamilton - Jacobi method is to consider the Lagrangian $L=L\left(q_{i}, \dot{q}_{i}, t\right)$ with the Hess matrix (1.2.5) of rank $(n-r), r<n$. Then we can
solve (1.3.1) for $\dot{q_{a}}$ in term of $q_{i}, \dot{x_{\mu}}, p_{a}$ and t as

$$
\begin{equation*}
\dot{q_{a}}=\dot{q}_{a}\left(q_{i}, \dot{x_{\mu}}, p_{b} ; t\right) . \tag{1.4.1}
\end{equation*}
$$

Substituting (1.3.1) into (1.3.2), we get

$$
\begin{equation*}
p_{\mu}=\frac{\partial L}{\partial \dot{q}_{\mu}}=-H_{\mu}\left(q_{i}, \dot{x_{\mu}}, p_{a} ; t\right) \tag{1.4.2}
\end{equation*}
$$

Relations (1.4.2 ) indicate the fact that the generalized momenta $p_{\mu}$ are not all independent which is a natural result of the singular nature of the Lagrangian. Although, it seems that $H_{\mu}$ are functions of $\dot{x_{\mu}}$, it is a task to show that they do not depend on it explicitly. The fundamental equations of the equivalent Lagrangian method read as

$$
\begin{equation*}
p_{0}=\frac{\partial S}{\partial t}=-H_{0}\left(q_{i}, \dot{x_{\mu}}, p_{a} ; t\right) ; \quad p_{a}=\frac{\partial S}{\partial q_{\mu}}, \quad p_{\mu}=\frac{\partial S}{\partial q_{\mu}} \equiv-H_{\mu} \tag{1.4.3}
\end{equation*}
$$

where the function $S \equiv S\left(q_{i}, t\right)$ is the action. The Hamiltonian $H_{0}$ reads as

$$
\begin{equation*}
H_{0}=p_{i} \dot{q}_{i}+\left.p_{\mu} \dot{x_{\mu}}\right|_{p \nu=-H \nu}-L\left(t, q_{i}, \dot{x_{\nu}} \dot{q}_{a}\right), \quad \quad \mu, \nu=n-r+1, \ldots, n . \tag{1.4.4}
\end{equation*}
$$

Like the functions $H_{\mu}$, the Hamiltonian $H_{0}$ is also not an explicit function of $\dot{x_{\mu}}$. Therefore, the function $S \equiv S\left(q_{i}, t\right)$ should satisfy the following set of Hamilton Jacobi partial differential equation (HJPDEs ) which is expressed as

$$
\begin{align*}
& H_{0}^{\prime}\left(t, x_{\mu}, q_{a}, p_{i}=\frac{\partial S}{\partial q_{i}}, p_{0}=\frac{\partial S}{\partial t}\right)=0  \tag{1.4.5}\\
& H_{\mu}^{\prime}\left(t, x_{\mu}, q_{a}, p_{i}=\frac{\partial S}{\partial q_{i}}, p_{0}=\frac{\partial S}{\partial t}\right)=0 \tag{1.4.6}
\end{align*}
$$

where

$$
\begin{equation*}
H_{0}^{\prime}=p_{0}+H_{\circ}, \quad \quad H_{\mu}^{\prime}=p_{\mu}+H_{\mu} \tag{1.4.7}
\end{equation*}
$$

Equations (1.4.5) and (1.4.6) may be expressed in a compact form as

$$
\begin{array}{r}
H_{\alpha}^{\prime}\left(t_{\beta}, q_{a}, p_{i}=\frac{\partial S}{\partial q_{i}}, p_{0}=\frac{\partial S}{\partial t}\right)=0,  \tag{1.4.8}\\
\alpha, \beta=0, n-r+1, \ldots, n, \quad a=1, \ldots, n-r,
\end{array}
$$

where

$$
\begin{equation*}
H_{\alpha}^{\prime}=p_{\alpha}+H_{\alpha} . \tag{1.4.9}
\end{equation*}
$$

The equations of motion are written as total differential equations in many variables $t_{\beta}$ as follows (7)

$$
\begin{array}{cc}
d q_{i}=\frac{\partial H_{\alpha}^{\prime}}{\partial p_{i}} d t_{\alpha}, & i=0,1, \ldots n, \\
d p_{a}=-\frac{\partial H_{\alpha}^{\prime}}{\partial q_{a}} d t_{\alpha}, & a=1, \ldots n-r, \\
d p_{\mu}=-\frac{\partial H_{\alpha}^{\prime}}{\partial q_{\mu}} d t_{\alpha}, & \alpha=0, n-r+1, \ldots, n . \tag{1.4.12}
\end{array}
$$

We define

$$
\begin{equation*}
Z=S\left(t_{\alpha}, q_{a}\right) \tag{1.4.13}
\end{equation*}
$$

and making use of Eq.(1.4.8) and definitions of generalized momenta (1.4.10, 1.4.11, 1.4.12, 1.4.13) we obtain

$$
\begin{equation*}
d Z=\frac{\partial S}{\partial t_{\alpha}} d t_{\alpha}+\frac{\partial S}{\partial q_{a}} d t_{a}=\left(-H_{\alpha} d t_{\alpha}+p_{a} d q_{a}\right)=\left(-H_{\alpha}+p_{a} \frac{\partial H_{\alpha}^{\prime}}{\partial p_{a}}\right) d t_{\alpha} \tag{1.4.14}
\end{equation*}
$$

Equations (1.4.10-1.4.12) and (1.4.14) are called the total differential equations for the characteristics. If these equations form a completely integrable set, the simultaneous solutions of them determine the function $S\left(t_{\alpha}, q_{a}\right)$ uniquely by the prescribed initial conditions. The set of Equations (1.26-1.28) is integrable if and only if the variations of $H_{0}^{\prime}$ and $H_{\mu}^{\prime} \quad$ vanish identically $(6 ; 17 ; 18)$ that is

$$
\begin{gather*}
d H_{0}^{\prime}=0  \tag{1.4.15}\\
d H_{\mu}^{\prime}=0, \quad \mu=n-r+1, \ldots, n . \tag{1.4.16}
\end{gather*}
$$

If condition (1.4.15) and (1.4.16) are not satisfied identically, one considers them as new constraints and again testes the integrability conditions. Hence, the canonical formulation leads to obtain the set of canonical phase space coordinates $q_{\alpha}$ and $p_{a}$
as functions of $t_{a}$, besides the canonical action integral is obtained in terms of the canonical coordinates. The Hamiltonians $H_{\alpha}$ are considered as the infinitesimal generators of canonical transformations given by parameters $t_{\alpha}$ respectively (6), (7) and (8).

### 1.5 Mixture of Lagrangian and Hamiltonian Formulation of Constrained System

### 1.5.1 Singular Lagrangian as Field System

Singular Lagrangian as field system has been studied in Ref [9]. As a natural extension of the Hamiltonian formulation we would like to study the Lagrangian approach of a constrained system. The usual way to pass from the Hamiltonian to the Lagrangian approach is to use Eqs. (1.4.10-1.4.12) Since there are additional constraints, $\mathrm{Eq}(1.4 .7)$ given in the phase space, they should also appear as constraints in the configuration space. As we have stated before, Eqs. (1.4.10-1.4.12) and Eq.(1.4.7) allow us to treat the system as a continuous or field system. Thus, we propose that the Euler-Lagrange equations of a constrained system are in the form (field system)

$$
\begin{equation*}
\frac{\partial}{\partial x_{\mu}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{\mu} q_{a}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{a}}=0 \tag{1.5.1}
\end{equation*}
$$

with constraints

$$
\begin{equation*}
d G_{\mu}=-\frac{\partial L \prime}{\partial x_{\mu}} d t \tag{1.5.2}
\end{equation*}
$$

where

$$
\begin{equation*}
L^{\prime}\left(x_{\mu}, \partial \mu q_{a}, \dot{x_{\nu}}, q_{a}\right) \equiv L\left[q_{a}, x_{\mu}, \dot{q_{a}}=\left(\partial \mu q_{a}\right) \dot{x}\right], \quad \dot{x_{\nu}}=\frac{d x_{\nu}}{d t} \tag{1.5.3}
\end{equation*}
$$

and

$$
\begin{equation*}
G_{\mu}=H_{\mu}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \tag{1.5.4}
\end{equation*}
$$

where $H_{\mu}$ is obtained form the Hamilton-Jacobi formalism (1.4.7).
One should notic that variations of constraints should be considered in order to have a consistent theory. Many physical models has been investigated by HamiltonJacobi approach,(9) -(16). The validity of this methods need further physical applications.

The thesis is arranged as follows: in chapter two models of Singular Lagrangians are studied using Dirac's method. In chapter three also the same models are investigated using Hamilton-Jacobi (Güler) method. The treatment of singular Lagrangian systems as field (continuous) systems is discussed in chapter four. Chapter five is devoted to the conclusion of the our study.

## Chapter 2

## On singular Lagrangian and Dirac's Method

## Chapter 2

## On singular Lagrangian and Dirac's Method

In this chapter, we study some singular Lagrangians from the classical mechanics of particles and apply Dirac's method for building the equations of motion. We will construct the total Hamiltonian $H_{T}$ of the systems and obtain the equations of motion. The consistency conditions will be discussed.

### 2.1 The First singular Lagrangian

The First singular Lagrangian is given as (20)

$$
\begin{equation*}
L=\frac{m}{2}\left(\dot{q}_{1}^{2}+\dot{q}_{2}^{2}+l^{2} \dot{q}_{3}^{2}+2 l \dot{q}_{1} \dot{q}_{3} \cos q_{3}+2 l \dot{q}_{2} \dot{q}_{3} \sin q_{3}\right)+V\left(q_{1}, q_{2}, q_{3}\right) \tag{2.1.1}
\end{equation*}
$$

The Lagrangian (2.1.1) is singular since the Hess matrix (1.2.5) is of rank two. The generalized momenta,(1.3.1) read as

$$
\begin{align*}
p_{1} & =m \dot{q}_{1}+m l \dot{q}_{3} \cos q_{3},  \tag{2.1.2}\\
p_{2} & =m \dot{q}_{2}+m l \dot{q}_{3} \sin q_{3}, \tag{2.1.3}
\end{align*}
$$

$$
\begin{equation*}
p_{3}=m l^{2} \dot{q}_{3}+m l\left(\dot{q}_{1} \cos q_{3}+\dot{q}_{2} \sin q_{3}\right) \tag{2.1.4}
\end{equation*}
$$

Multiplying (2.1.2) by $\cos q_{3}$ and (2.1.3) by $\sin q_{3}$ and then subtracting the sum of the result from (2.1.5), one gets the primary constraints (1.2.8) according to Dirac as

$$
\begin{equation*}
\phi_{1}=p_{3}-l p_{1} \cos q_{3}-l p_{2} \sin q_{3}=0 \tag{2.1.5}
\end{equation*}
$$

Now, let us rewrite (2.1.2) and (2.1.3) as

$$
\begin{align*}
& p_{1}-m \dot{q}_{1}=m l \dot{q_{3}} \cos q_{3},  \tag{2.1.6}\\
& p_{2}-m \dot{q}_{2}=m l \dot{q}_{3} \sin q_{3} . \tag{2.1.7}
\end{align*}
$$

From Eqs (2.1.6) and (2.1.7) one gets

$$
\begin{equation*}
p_{1} \dot{q}_{1}+p_{2} \dot{q}_{2}=\frac{p_{1}^{2}+p_{2}^{2}}{2 m}+\frac{m}{2}\left(\dot{q}_{1}^{2}+{\dot{q_{2}}}^{2}\right)-m l^{2} \dot{q}_{3}^{2} . \tag{2.1.8}
\end{equation*}
$$

The usual Hamiltonian (1.3.5) is

$$
\begin{equation*}
H_{0}=p_{1} \dot{q_{1}}+p_{2} \dot{q_{2}}+p_{3} \dot{q_{3}}-L \tag{2.1.9}
\end{equation*}
$$

Using (2.1.2) , (2.1.3) and (2.1.4), (2.1.9) takes the form

$$
\begin{equation*}
H_{0}=\frac{p_{1}^{2}+p_{2}^{2}}{2 m}-V, \tag{2.1.10}
\end{equation*}
$$

and using (1.3.4) , the total Hamiltonian is

$$
\begin{equation*}
H_{T}=\frac{p_{1}^{2}+p_{2}^{2}}{2 m}-V+\nu\left(p_{3}-l p_{1} \cos q_{3}-l p_{2} \sin q_{3}\right) \tag{2.1.11}
\end{equation*}
$$

Now, the consistency condition reads as

$$
\begin{equation*}
\dot{\phi}_{1}=\left[\phi_{1}, H_{T}\right]=V_{, 3}-l \cos q_{3} V_{, 1}-l \sin q_{3} V_{, 2} \tag{2.1.12}
\end{equation*}
$$

where,

$$
\begin{equation*}
V_{, i}=\frac{\partial V}{\partial q_{i}}, \quad i=1,2,3 \tag{2.1.13}
\end{equation*}
$$

Relation (2.1.12) leads to the secondary constraint which is a relation between the coordinates

$$
\begin{equation*}
\phi_{2}=V_{, 3}-l \cos q_{3} V_{, 1}-l \sin q_{3} V_{, 2}, \tag{2.1.14}
\end{equation*}
$$

so one can write the secondary constraint in the form

$$
\begin{equation*}
\phi_{2}=q_{3}-F\left(q_{1}, q_{2}\right) . \tag{2.1.15}
\end{equation*}
$$

Now let us evaluate $\dot{\phi}_{2}$

$$
\begin{equation*}
\dot{\phi}_{2}=\left[\phi_{2}, H_{T}\right]=F_{, 1} \frac{P_{1}}{m}+F_{, 2} \frac{P_{2}}{m}-\nu\left(1+l \cos q_{3} F_{, 1}+l \sin q_{3} F_{, 2}\right) \equiv 0 \tag{2.1.16}
\end{equation*}
$$

from which we get the multiplier $\nu$ as

$$
\begin{equation*}
\nu=\frac{p_{2} F_{, 1}+p_{2} F_{, 2}}{m\left(1+l \cos q_{3} F_{, 1}+l \sin q_{3} F_{, 2}\right)} . \tag{2.1.17}
\end{equation*}
$$

The equations of motion (1.3.6) and (1.3.7) read as

$$
\begin{gather*}
\dot{q_{1}}=\frac{p_{1}}{m}-\nu l \cos q_{3}  \tag{2.1.18}\\
\dot{q_{2}}=\frac{p_{2}}{m}-\nu l \sin q_{3}  \tag{2.1.19}\\
\dot{q_{3}}=\nu  \tag{2.1.20}\\
\dot{p_{1}}=\frac{\partial V}{\partial q_{1}}  \tag{2.1.21}\\
\dot{p_{2}}=\frac{\partial V}{\partial q_{2}}  \tag{2.1.22}\\
\dot{p_{3}}=\frac{\partial V}{\partial q_{3}}-\nu\left(l p_{1} \sin q_{3}-l p_{2} \cos q_{3}\right) . \tag{2.1.23}
\end{gather*}
$$

The set of equations (2.1.18-2.1.23) with (2.1.17) represent a consistent set of ordinary differential equations.

### 2.2 Mittelstaedt's Lagrangian

The second model is Mittelstaedt's Lagrangian model (20), which is given as

$$
\begin{equation*}
L=\frac{1}{2 m}\left(\dot{q}_{1}+\dot{q}_{2}\right)^{2}+\frac{1}{2 \mu} \dot{q}_{3}^{2}+V\left(q_{1}, q_{2}, q_{3}\right) . \tag{2.2.1}
\end{equation*}
$$

As the Hess matrix of the above Lagrangian is of rank two, this Lagrangian is Singular.

We start with obtaining the momenta (1.2.7), which are given as

$$
\begin{equation*}
p_{1}=p_{2}=\frac{1}{m}\left(\dot{q}_{1}+\dot{q}_{2}\right), \quad p_{3}=\frac{1}{\mu} \dot{q}_{3}, \tag{2.2.2}
\end{equation*}
$$

The primary constraint is then

$$
\begin{equation*}
\phi_{1}=p_{2}-p_{1}=0 \tag{2.2.3}
\end{equation*}
$$

The original Hamiltonian becomes

$$
\begin{equation*}
H_{0}=\frac{m}{2} p_{1}^{2}+\frac{\mu}{2} p_{3}^{2}-V . \tag{2.2.4}
\end{equation*}
$$

The total Hamiltonian is then

$$
\begin{equation*}
H_{T}=\frac{m}{2} p_{1}^{2}+\frac{\mu}{2} p_{3}^{2}-V+\nu\left(p_{2}-p_{1}\right) . \tag{2.2.5}
\end{equation*}
$$

The consistency condition $\dot{\phi}_{1}=\left[\phi_{1}, H_{T}\right]$ leads to the constraint

$$
\begin{equation*}
\phi_{2}=\frac{\partial V}{\partial q_{1}}-\frac{\partial V}{\partial q_{2}}=0 . \tag{2.2.6}
\end{equation*}
$$

This is a relation between $q_{1}, q_{2}$ and $q_{3}$ which briefly, is written as

$$
\begin{equation*}
\phi_{2}=q_{2}-F\left(q_{1}, q_{3}\right)=0 . \tag{2.2.7}
\end{equation*}
$$

We then build the consistency condition $\dot{\phi}_{2}$ as

$$
\begin{equation*}
\dot{\phi}_{2}=\left[\phi_{2}, H_{T}\right], \tag{2.2.8}
\end{equation*}
$$

from which we find

$$
\begin{equation*}
\nu\left(1+F_{, 1}\right)-m P_{1} F_{, 1}-\mu P_{3} F_{, 3}=0, \quad\left(F_{, i}=\frac{\partial F}{\partial q_{i}}\right) \tag{2.2.9}
\end{equation*}
$$

From Eq (2.31) one allows fixing varible $\nu$ as

$$
\begin{equation*}
\nu=\frac{m p_{1} F_{, 1}+\mu p_{3} F_{, 3}}{1+F_{, 1}} . \tag{2.2.10}
\end{equation*}
$$

We can now write the canonical equations of motion (1.3.6) and (1.3.7) as

$$
\begin{align*}
& \dot{q}_{1}=m p_{1}-\nu, \quad \dot{q}_{2}=\nu, \quad \dot{q}_{3}=\mu p_{3},  \tag{2.2.11}\\
& \dot{p_{1}}=V_{, 1}, \quad \dot{p_{2}}=V_{, 1}, \quad \dot{p_{3}}=V_{, 3} . \tag{2.2.12}
\end{align*}
$$

Eqs. (2.2.11) and (2.2.12) with (2.2.10) represent a set of consistent differential equations.

### 2.3 Deriglazov Lagrangian

The third model is Deriglazov Lagrangian which is given as (20)

$$
\begin{equation*}
L=q_{2}^{2} \dot{q}_{1}^{2}+q_{1}^{2} \dot{q}_{2}^{2}+2 q_{1} q_{2} \dot{q}_{1} \dot{q}_{2}+V\left(q_{1}, q_{2}\right) . \tag{2.3.1}
\end{equation*}
$$

This Lagrangian is singular since the Hess matrix is of rank one. The momenta (1.3.1) read as

$$
\begin{align*}
& p_{1}=2 q_{2}^{2} \dot{q_{1}}+2 q_{1} q_{2} \dot{q_{2}},  \tag{2.3.2}\\
& p_{2}=2 q_{1}^{2} \dot{q_{2}}+2 q_{1} q_{2} \dot{q_{1}} . \tag{2.3.3}
\end{align*}
$$

Here the momenta $p_{1}$ and $p_{1}$ are not independent. The primary constraint is then

$$
\begin{equation*}
\dot{\phi}_{1}=q_{1} p_{1}-q_{2} p_{2}=0 . \tag{2.3.4}
\end{equation*}
$$

The original Hamiltonian takes the form

$$
\begin{equation*}
H_{0}=\frac{p_{1}^{2}}{4 q_{2}^{2}}-V\left(q_{1}, q_{2}\right) \tag{2.3.5}
\end{equation*}
$$

and the total Hamiltonian is then

$$
\begin{equation*}
H_{T}=\frac{p_{1}^{2}}{4 q_{2}^{2}}-V\left(q_{1}, q_{2}\right)+\nu\left(q_{1} P_{1}-q_{2} p_{2}\right) \tag{2.3.6}
\end{equation*}
$$

The consistency condition is

$$
\begin{equation*}
\dot{\phi}_{1}=\left[\phi_{1}, H_{T}\right]=q_{1} V_{, 1}-q_{2} V_{, 2} \equiv \phi_{2}=0, \tag{2.3.7}
\end{equation*}
$$

where $\phi_{2}$ is the secondary constraint.
Again $\dot{\phi}_{2}$ is

$$
\begin{equation*}
\dot{\phi}_{2}=\left[\phi_{2}, H_{T}\right]=0, \tag{2.3.8}
\end{equation*}
$$

from which we find

$$
\begin{equation*}
-F_{, 1} \frac{p_{1}}{2 q_{2}^{2}}-q_{1} F_{, 1} \nu-q_{2} \nu=0 \tag{2.3.9}
\end{equation*}
$$

From Eq.(2.3.9) one allows fixing variable $\nu$ as

$$
\begin{equation*}
\nu=-\frac{p_{1}}{2 F^{2}\left(F+q_{1} F_{, 1}\right)} F_{, 1}, \tag{2.3.10}
\end{equation*}
$$

with

$$
\begin{equation*}
q_{2}=F\left(q_{1}\right) . \tag{2.3.11}
\end{equation*}
$$

Therefore, the canonical equations of motion are

$$
\begin{equation*}
\dot{q}_{1}=\frac{p_{1}}{2 q_{2}^{2}}+q_{1} \nu \tag{2.3.12}
\end{equation*}
$$

Substituting Eq.(2.3.10) in (2.3.12) we get

$$
\begin{equation*}
\dot{q}_{1}=\frac{p_{1}}{2 F\left(F+q_{1} F_{, 1}\right)} . \tag{2.3.13}
\end{equation*}
$$

The other equation of motion is

$$
\begin{equation*}
\dot{p_{1}}=V_{, 1}+\frac{p_{1}^{2}}{2 F^{2}\left(F+q_{1} F_{, 1}\right)} F_{, 1} . \tag{2.3.14}
\end{equation*}
$$

One notices that Newton's equations of motion can be obtained if we let $V(x, y)=$ $x^{2}+y^{2} \quad$ and $\quad F(x)= \pm x$ as

$$
\begin{equation*}
2 F\left(F+q_{1} F_{, 1}\right) \ddot{q}_{1}+2 F\left(2 F_{, 1}+q_{1} F_{, 11}\right) \dot{q}_{1}^{2}-V_{, 1}=0 \tag{2.3.15}
\end{equation*}
$$

## Chapter 3

## Hamilton-Jacobi Method

## Chapter 3

## Hamilton-Jacobi Method

In this chapter we study some singular Lagrangian systems from the classical mechanics of particles and apply Hamilton-Jacobi Method to construct HamiltonJacobi Partial Differential Equations (HJPDE), and then we write the equations of motion.

### 3.1 Charged Particle Moving in a constant Magnetic Field

The motion of charged particle in a plane is described by the singular Lagrangian (21).

$$
\begin{equation*}
L=\frac{1}{2}\left(\dot{q}_{1}-q_{3} q_{2}\right)^{2}+\frac{1}{2}\left(\dot{q}_{2}+q_{3} q_{1}\right)^{2} . \tag{3.1.1}
\end{equation*}
$$

The rank of the Hess matrix (1.2.5) is two. Then the singularity of the Lagrangian enables us to write the generalized momenta (1.3.1) and (1.3.2) as

$$
\begin{align*}
& p_{1}=\dot{q}_{1}-q_{3} q_{2},  \tag{3.1.2}\\
& p_{2}=\dot{q_{2}}+q_{3} q_{1}, \tag{3.1.3}
\end{align*}
$$

$$
\begin{equation*}
p_{3}=0 \equiv-H_{3} . \tag{3.1.4}
\end{equation*}
$$

We solve (3.1.2) and (3.1.3) for $\dot{q}_{1}$ and $\dot{q}_{2}$ in terms of $p_{1}$ and $p_{2}$ as

$$
\begin{align*}
& \dot{q_{1}}=p_{1}+q_{3} q_{2} \equiv \omega_{1}  \tag{3.1.5}\\
& \dot{q_{2}}=p_{2}-q_{3} q_{1} \equiv \omega_{2} \tag{3.1.6}
\end{align*}
$$

The canonical Hamiltonian $H_{0}(1.3 .5)$ is then

$$
\begin{align*}
H_{0} & =p_{a} \dot{q}_{a}+p_{\mu} \dot{q}_{\mu}-\left.L\right|_{\dot{q}_{a} \equiv \omega_{a}} \\
& =\frac{1}{2} p_{1}^{2}+\frac{1}{2} p_{2}^{2}+p_{1} q_{3} q_{2}-p_{2} q_{3} q_{1} . \tag{3.1.7}
\end{align*}
$$

The set of (HJPDE) according to Eq.(1.4.7) is

$$
\begin{equation*}
H_{0}^{\prime}=p_{0}+\frac{1}{2} p_{1}^{2}+\frac{1}{2} p_{2}^{2}+p_{1} q_{3} q_{2}-p_{2} q_{3} q_{1}=0 \tag{3.1.8}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{3}^{\prime}=p_{3}+H_{3}=p_{3}=0 \tag{3.1.9}
\end{equation*}
$$

Relation (3.1.8) and (3.1.9) are the constraints that restrict the system. The total differential equations of motion (1.4.10), (1.4.11) and (1.4.12) are

$$
\begin{align*}
d q_{1} & =\frac{\partial H_{0}^{\prime}}{\partial p_{1}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{1}} d q_{3}  \tag{3.1.10}\\
d q_{2} & =\frac{\partial H_{0}^{\prime}}{\partial p_{2}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{2}} d q_{3}  \tag{3.1.11}\\
d q_{3} & =\frac{\partial H_{0}^{\prime}}{\partial p_{3}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{3}} d q_{3}  \tag{3.1.12}\\
d p_{1} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{1}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{1}} d q_{3}  \tag{3.1.13}\\
d p_{2} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{2}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{2}} d q_{3}  \tag{3.1.14}\\
d p_{3} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{3}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{3}} d q_{3} \tag{3.1.15}
\end{align*}
$$

Substituting Eqs.(3.1.8) and (3.1.9) in eqs.(3.1.10-3.1.15), we obtain the total differential equations of motion as

$$
\begin{gather*}
d q_{1}=\left(p_{1}+q_{3} q_{2}\right) d \tau  \tag{3.1.16}\\
d q_{2}=\left(p_{2}-q_{3} q_{1}\right) d \tau  \tag{3.1.17}\\
d q_{3}=d q_{3}  \tag{3.1.18}\\
d p_{1}=p_{2} q_{3} d \tau  \tag{3.1.19}\\
d p_{2}=-p_{1} q_{3} d \tau  \tag{3.1.20}\\
d p_{3}=-\left(p_{1} q_{2}-p_{2} q_{1}\right) d \tau \tag{3.1.21}
\end{gather*}
$$

To check whether the above set of equations is integrable or not, let us consider the total variations of $H_{0}^{\prime}$ and $H_{3}^{\prime}$. If fact

$$
\begin{gather*}
d H_{0}^{\prime}=0  \tag{3.1.22}\\
d H_{3}^{\prime}=d p_{3}=\left(-p_{1} q_{2}+p_{2} q_{1}\right) d \tau \tag{3.1.23}
\end{gather*}
$$

since $d H_{3}^{\prime}$ is not identically zero. we have a new constraint $H_{4}^{\prime}$,

$$
\begin{equation*}
H_{4}^{\prime}=\left(p_{1} q_{2}-p_{2} q_{1}\right) \equiv 0 . \tag{3.1.24}
\end{equation*}
$$

Thus for a valid theory, the total differential of $H_{4}^{\prime}$ is identically zero,

$$
\begin{equation*}
d H_{4}^{\prime}=p_{1} d q_{2}+q_{2} d p_{1}-p_{2} d q_{1}-q_{1} d p_{2}=0 \tag{3.1.25}
\end{equation*}
$$

so the system of Eqs.(3.1.16-3.1.21) together with Eq.(3.1.25) is integrable.

### 3.2 The second singular Lagrangian

As a second model, let us consider the singular Lagrangian,

$$
\begin{equation*}
L=\frac{m}{2}\left(\dot{q}_{1}^{2}+\dot{q}_{2}^{2}+l^{2} \dot{q}_{3}^{2}+2 l \dot{q}_{1} \dot{q}_{3} \cos q_{3}+2 l \dot{q}_{2} \dot{q}_{3} \sin q_{3}\right)+V\left(q_{1}, q_{2}, q_{3}\right) \tag{3.2.1}
\end{equation*}
$$

where $l$ and $m$ are constants,
The singularity of the Lagrangian enables us to write the generalized momenta in chapter two Eqs.(2.1.2),(2.1.3) and(2.1.4)

$$
\begin{gather*}
p_{1}=m \dot{q}_{1}+m l \dot{q_{3}} \cos q_{3},  \tag{3.2.2}\\
p_{2}=m \dot{q}_{2}+m l \dot{q}_{3} \sin q_{3},  \tag{3.2.3}\\
p_{3}=m l^{2} \dot{q}_{3}+m l\left(\dot{q_{1}} \cos q_{3}+\dot{q_{2}} \sin q_{3}\right) . \tag{3.2.4}
\end{gather*}
$$

Multiplying equation (3.2.2) by $l \cos q_{3}$ and (3.2.3) by $l \sin q_{3}$ one gets the constraint relation

$$
\begin{equation*}
H_{3}=p_{3}-l p_{1} \cos q_{3}-l p_{2} \sin q_{3}=0 \tag{3.2.5}
\end{equation*}
$$

The canonical Hamiltonian $H_{0}(1.3 .5)$ is then

$$
\begin{align*}
H_{0} & =p_{a} \dot{q}_{a}+p_{\mu} \dot{q}_{\mu}-\left.L\right|_{\dot{q}_{a} \equiv \omega_{a}} \\
& =\frac{p_{1}^{2}+p_{2}^{2}}{2 m}-V\left(q_{1}, q_{2}, q_{3}\right) . \tag{3.2.6}
\end{align*}
$$

The set of (HJPBE) according to $\mathrm{Eq}(1.4 .7)$ is

$$
\begin{gather*}
H_{0}^{\prime}=p_{0}+\frac{p_{1}^{2}+p_{2}^{2}}{2 m}-V\left(q_{1}, q_{2}, q_{3}\right)=0  \tag{3.2.7}\\
H_{3}^{\prime}=p_{3}-l p_{1} \cos q_{3}-l p_{2} \sin q_{3}=0 . \tag{3.2.8}
\end{gather*}
$$

The total differential equations of motion (1.4.10), (1.4.11) and (1.4.12) are

$$
\begin{align*}
d q_{1} & =\frac{\partial H_{0}^{\prime}}{\partial p_{1}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{1}} d q_{3}  \tag{3.2.9}\\
d q_{2} & =\frac{\partial H_{0}^{\prime}}{\partial p_{2}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{2}} d q_{3}  \tag{3.2.10}\\
d q_{3} & =\frac{\partial H_{0}^{\prime}}{\partial p_{3}} d \tau+\frac{\partial H_{3}^{\prime}}{\partial p_{3}} d q_{3}  \tag{3.2.11}\\
d p_{1} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{1}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{1}} d q_{3}  \tag{3.2.12}\\
d p_{2} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{2}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{2}} d q_{3},  \tag{3.2.13}\\
d p_{3} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{3}} d \tau-\frac{\partial H_{3}^{\prime}}{\partial q_{3}} d q_{3} . \tag{3.2.14}
\end{align*}
$$

Substituting Eqs.(3.2.7) and (3.2.8) in Eqs.(3.2.9-3.2.14), we obtain the total differential equations of motion as

$$
\begin{gather*}
d q_{1}=\frac{p_{1}}{m} d t-l \cos q_{3} d q_{3},  \tag{3.2.15}\\
d q_{2}=\frac{p_{2}}{m} d t-l \sin q_{3} d q_{3}  \tag{3.2.16}\\
d q_{3}=d q_{3}  \tag{3.2.17}\\
d p_{1}=V_{, 1} d t  \tag{3.2.18}\\
d p_{2}=V_{, 2} d t  \tag{3.2.19}\\
d p_{3}=V_{, 3} d t-\left(l p_{1} \sin q_{3}-l p_{2} \cos q_{3}\right) d q_{3} . \tag{3.2.20}
\end{gather*}
$$

From Eq. (3.2.17), one conduces that $q_{3}=$ constant. The set of equation of motion (3.2.15-3.2.20) are integrable if the variations of (3.2.7) and (3.2.8) are identically satisfied, that is

$$
\begin{equation*}
d H_{0}^{\prime}=\left(V_{, 1} l \cos q_{3}+V_{, 2} l \sin q_{3}-V_{, 3}\right) d q_{3} \tag{3.2.21}
\end{equation*}
$$

Similarly the variation of $H_{3}^{\prime}$ takes the form

$$
\begin{equation*}
d H_{3}^{\prime}=\left(V_{, 1} l \cos q_{3}+V_{, 2} l \sin q_{3}-V_{, 3}\right) d t . \tag{3.2.22}
\end{equation*}
$$

To be identically satisfied we should choose $V\left(q_{1}, q_{2}, q_{3}\right)$ such that

$$
\begin{equation*}
V_{, 3}=V_{, 1} l \cos q_{3}+V_{, 2} l \sin q_{3} \tag{3.2.23}
\end{equation*}
$$

### 3.3 The Mittelstaedt's Lagrangian

As a third model, let us consider the Mittelstaedt's singular Lagrangian (20)

$$
\begin{equation*}
L=\frac{1}{2 m}\left(\dot{q}_{1}+\dot{q}_{2}\right)^{2}+\frac{1}{2 \mu} \dot{q}_{3}^{2}+V\left(q_{1}, q_{2}, q_{3}\right), \tag{3.3.1}
\end{equation*}
$$

where $m$ and $\mu$ are constants.
The singularity of the Lagrangian enables us to write generalized momenta (1.3.1) and (1.3.2) as

$$
\begin{equation*}
p_{1}=p_{2}=\frac{1}{m} \dot{q}_{2}+\dot{q_{2}}, \tag{3.3.2}
\end{equation*}
$$

and

$$
\begin{equation*}
p_{3}=\frac{1}{\mu} \dot{q_{3}} . \tag{3.3.3}
\end{equation*}
$$

we solve (3.3.2) and (3.3.3) for $\dot{q}_{1}, \quad \dot{q}_{2}$ and $\dot{q}_{3}$ interns of $p_{1}, \quad p_{2}$ and $p_{3}$ as

$$
\begin{gather*}
\dot{q_{2}}+\dot{q_{2}}=m p_{1}-m p_{2},  \tag{3.3.4}\\
\dot{q_{3}}=\mu p_{3} . \tag{3.3.5}
\end{gather*}
$$

The constraint relation is

$$
\begin{equation*}
H_{2}=p_{1}-p_{2}=0 \tag{3.3.6}
\end{equation*}
$$

The canonical Hamiltonian $H_{0}$ takes the form

$$
\begin{align*}
H_{0} & =p_{a} \dot{q}_{a}+p_{\mu} \dot{q}_{\mu}-\left.L\right|_{q_{a} \equiv \omega_{a}} \\
& =\frac{m}{2} p_{1}^{2}+\frac{\mu}{2} p_{3}^{2}-V\left(q_{1}, q_{2}, q_{3}\right) . \tag{3.3.7}
\end{align*}
$$

The set of (HJPBE) according to $\mathrm{Eq}(1.4 .7)$ is

$$
\begin{equation*}
H_{0}^{\prime}=p_{0}+\frac{m}{2} p_{1}^{2}+\frac{\mu}{2} p_{3}^{2}-V\left(q_{1}, q_{2}, q_{3}\right)=0 \tag{3.3.8}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{2}^{\prime}=p_{1}-p_{2}=0 \tag{3.3.9}
\end{equation*}
$$

The total differential equations of motion (1.4.10), (1.4.11) and (1.4.12) read as

$$
\begin{align*}
d q_{1} & =\frac{\partial H_{0}^{\prime}}{\partial p_{1}} d \tau+\frac{\partial H_{2}^{\prime}}{\partial p_{1}} d q_{2}  \tag{3.3.10}\\
d q_{2} & =\frac{\partial H_{0}^{\prime}}{\partial p_{2}} d \tau+\frac{\partial H_{2}^{\prime}}{\partial p_{2}} d q_{2}  \tag{3.3.11}\\
d q_{3} & =\frac{\partial H_{0}^{\prime}}{\partial p_{3}} d \tau+\frac{\partial H_{2}^{\prime}}{\partial p_{3}} d q_{2}  \tag{3.3.12}\\
d p_{1} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{1}} d \tau-\frac{\partial H_{2}^{\prime}}{\partial q_{1}} d q_{2}  \tag{3.3.13}\\
d p_{2} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{2}} d \tau-\frac{\partial H_{2}^{\prime}}{\partial q_{2}} d q_{2}  \tag{3.3.14}\\
d p_{3} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{3}} d \tau-\frac{\partial H_{2}^{\prime}}{\partial q_{3}} d q_{2} \tag{3.3.15}
\end{align*}
$$

Substituting Eqs. (3.3.8) and (3.3.9) in Eqs.(3.3.10-3.3.15), we obtain the total differential equations of motion as

$$
\begin{gather*}
d q_{1}=m p_{1} d t-d q_{2}  \tag{3.3.16}\\
d q_{2}=d q_{2}  \tag{3.3.17}\\
d q_{3}=\mu p_{3} d t \tag{3.3.18}
\end{gather*}
$$

$$
\begin{align*}
& d p_{1}=V_{, 1} d t  \tag{3.3.19}\\
& d p_{2}=V_{, 2} d t  \tag{3.3.20}\\
& d p_{3}=V_{, 3} d t \tag{3.3.21}
\end{align*}
$$

The set of equation of motion (3.3.16-3.3.21) are integrable if only if the variations of (3.3.8) and (3.3.9) are identically satisfied. The variation of $H_{0}^{\prime}$ is

$$
\begin{equation*}
d H_{0}^{\prime}=\left(V_{, 1}-V_{, 2}\right) d q_{2} \tag{3.3.22}
\end{equation*}
$$

which is identically zero since $q_{2}$ is constant.

$$
\begin{align*}
d H^{\prime} & =d p_{2}-d p_{1}  \tag{3.3.23}\\
& =\left(V_{, 2}-V_{, 1}\right) d t
\end{align*}
$$

In order to obtain an integrable system $V_{, 1}$ must be equal to $V_{, 2}$

### 3.4 The Deriglazov Lagrangian

The last model is the Deriglazov singular Lagrangian (20)

$$
\begin{equation*}
L=q_{2}^{2} \dot{q}_{1}^{2}+q_{1}^{2} \dot{q}_{2}^{2}+2 q_{1} q_{2} \dot{q}_{1} \dot{q}_{2}+V\left(q_{1}, q_{2}\right) \tag{3.4.1}
\end{equation*}
$$

This Lagrangian is singular since the Hess matrix is of rank one, and the generalized momenta (1.3.1) and (1.3.2) read as

$$
\begin{align*}
& p_{1}=2 q_{2}^{2} \dot{q}_{1}+2 q_{1} q_{2} \dot{q_{2}},  \tag{3.4.2}\\
& p_{2}=2 q_{1}^{2} \dot{q_{2}}+2 q_{1} q_{2} \dot{q_{1}} . \tag{3.4.3}
\end{align*}
$$

Here $p_{1}$ and $p_{2}$ are dependent. Multiplying Eq. (3.4.2) in $q_{1}$ and Eq.(3.4.3) in $q_{2}$ and solving for $p_{1}$, we get becomes constraint equation are

$$
\begin{equation*}
p_{1}=\frac{q_{2} p_{2}}{q_{1}} . \tag{3.4.4}
\end{equation*}
$$

using (3.4.4), the canonical Hamiltonian $H_{0}$ (1.3.5) is then

$$
\begin{align*}
H_{0} & =p_{1} \dot{q}_{1}+p_{2} \dot{q_{2}}-L \\
& =\frac{p_{1}^{2}}{4 q_{2}^{2}}-V\left(q_{1}, q_{2}\right) \tag{3.4.5}
\end{align*}
$$

The Set of (HJPDE) according to Eq (1.4.7) is

$$
\begin{equation*}
H_{0}^{\prime}=p_{0}+\frac{p_{1}^{2}}{4 q_{2}^{2}}-V\left(q_{1}, q_{2}\right)=0 \tag{3.4.6}
\end{equation*}
$$

The above equation are the constraints restricting the system.
The total differential equations of motion (1.4.10), (1.4.11) and (1.4.12) are

$$
\begin{align*}
d q_{1} & =\frac{\partial H_{0}^{\prime}}{\partial p_{1}} d \tau  \tag{3.4.7}\\
d q_{2} & =\frac{\partial H_{0}^{\prime}}{\partial p_{2}} d \tau  \tag{3.4.8}\\
d p_{1} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{1}} d \tau  \tag{3.4.9}\\
d p_{2} & =-\frac{\partial H_{0}^{\prime}}{\partial q_{2}} d \tau \tag{3.4.10}
\end{align*}
$$

Substituting Eq. (3.4.7) in Eqs.(3.4.7-3.4.10), we obtain the total differential equations of motion as

$$
\begin{gather*}
d q_{1}=\frac{p_{1}}{2 q_{2}^{2}} d t  \tag{3.4.11}\\
d q_{2}=0  \tag{3.4.12}\\
d p_{1}=V_{, 1} d t  \tag{3.4.13}\\
d p_{2}=V_{, 2} d t \tag{3.4.14}
\end{gather*}
$$

The set of equation of motion (3.4.11-3.4.14) are integrable if the variations of (3.4.6) is identically satisfied, we notice that the variation

$$
\begin{equation*}
d H_{0}^{\prime}=\left(q_{2} V_{, 2}-q_{1} V_{, 1}\right) d q_{1} \equiv H_{0}^{\prime \prime} d q_{1} \tag{3.4.15}
\end{equation*}
$$

is identically satisfied for a choice $V\left(q_{1}, q_{2}\right)=q_{1} q_{2}$

## Chapter 4

## Singular Lagrangian as Field Systems

## Chapter 4

## Singular Lagrangian As Field

## Systems

The link between the treatment of singular Lagrangian as field system and the general Hamiltonian approach is studied. It is shown that the singular Lagrangian as field system are always in exact aggrement with the general approaches (9). The equations of motion in this treatment are second order partial differential equation

### 4.1 Preliminaries

The Euler-Lagrange equations for field system is given as

$$
\begin{equation*}
\frac{\partial}{\partial x_{\mu}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial \mu q_{a}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{a}}=0 \tag{4.1.1}
\end{equation*}
$$

and the constraints relation is defined as

$$
\begin{equation*}
d G_{\mu}=-\frac{\partial L \prime}{\partial x_{\mu}} d t \tag{4.1.2}
\end{equation*}
$$

where the modified Lagrangian $L^{\prime}$ is defined as

$$
\begin{equation*}
L^{\prime}\left(x_{\mu}, \partial_{\mu} q_{a}, \dot{x_{\nu}}, q_{a}\right) \equiv L\left(q_{a}, x_{\mu}, \dot{q_{a}}=\left(\partial \mu q_{a}\right) \dot{x_{\nu}}\right), \quad \dot{x_{\nu}}=\frac{d x_{\nu}}{d t} \tag{4.1.3}
\end{equation*}
$$

and the constraint relations are

$$
\begin{equation*}
G_{\mu}=H_{\mu}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \tag{4.1.4}
\end{equation*}
$$

### 4.2 Examples

In this section, the singular Lagrangians which are investigated in chapter two and chapter three will be studied. As field systems.(or continuous system).

### 4.2.1 Example One

As a first example we consider the singular Lagrangian

$$
\begin{equation*}
L=\frac{1}{2} \dot{q}_{1}^{2}-\frac{1}{4}\left(\dot{q}_{2}^{2}-2 \dot{q}_{2} \dot{q}_{3}+\dot{q}_{3}^{2}\right)+\left(q_{1}+q_{3}\right) \dot{q}_{2}-q_{1}-q_{2}-q_{3}^{2} . \tag{4.2.1}
\end{equation*}
$$

Since the rank of the Hess matrix is two, this system can be treated as a continuous system in the form

$$
\begin{equation*}
q_{1}=q_{1}\left(t, q_{2}\right), \quad q_{3}=q_{3}\left(t, q_{2}\right) \tag{4.2.2}
\end{equation*}
$$

Now, let us write $\dot{q}_{1}$ and $\dot{q}_{3}$ as

$$
\begin{equation*}
\dot{q}_{1}=\frac{d q_{1}}{d t}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}, \quad \dot{q}_{3}=\frac{d q_{3}}{d t}=\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2} . \tag{4.2.3}
\end{equation*}
$$

Substituting (4.2.3) into (4.2.1), we get the modified Lagrangian $L^{\prime}$ as

$$
\begin{gather*}
L^{\prime}=\frac{1}{2}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}\right)^{2}-\frac{1}{4} \dot{q}_{2}^{2}+\frac{1}{2} \dot{q}_{2}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)-\frac{1}{4}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)^{2} \\
+\left(q_{1}+q_{3}\right) \dot{q}_{2}-q_{1}-q_{2}-q_{3}^{2} \tag{4.2.4}
\end{gather*}
$$

The Euler-Lagrange equations (4.1.1)read as

$$
\begin{array}{rlrl} 
& & \frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{1}\right)}\right)+\frac{\partial}{\partial q_{2}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{2} q_{1}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{1}} & =0, \\
\text { and } \quad \frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{3}\right)}\right)+\frac{\partial}{\partial q_{2}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{2} q_{3}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{3}} & =0 . \tag{4.2.5}
\end{array}
$$

where $x_{0} \equiv t, \quad x_{2} \equiv q_{2}$.
More explicitly, the second order partial differential equations are

$$
\begin{array}{r}
\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}}-\dot{q}_{2}+1=0 \\
\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}-\ddot{q}_{2}+2 \dot{q}_{2}-4 q_{3}=0 \tag{4.2.7}
\end{array}
$$

Now, we have to check the validity of constraint (4.1.2). As they were defined in (9). The usual Hamiltonian and the constraint relation are

$$
\begin{equation*}
H_{0}=\frac{1}{2}\left(p_{1}^{2}-2 p_{3}^{2}\right)+q_{1}+q_{2}+q_{3}^{2} \tag{4.2.8}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{2}=p_{3}-q_{1}-q_{3} \tag{4.2.9}
\end{equation*}
$$

Hence,

$$
\begin{align*}
G_{0} & =H_{0}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \\
& =\frac{1}{2}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}\right)^{2}-\frac{1}{4} \dot{q}_{2}^{2}-\frac{1}{4}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)^{2}  \tag{4.2.10}\\
& +\frac{1}{2} \dot{q}_{2}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)^{2}+q_{1}+q_{2}+q_{3}^{2},
\end{align*}
$$

and

$$
\begin{align*}
G_{2} & =H_{2}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \\
& =-\frac{1}{2} \dot{q}_{2}+\frac{1}{2}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)+q_{1}+q_{3} \tag{4.2.11}
\end{align*}
$$

Now, we are ready to test whether (4.1.2) are satisfied or not. In fact (4.1.2) for $\mu=0$ is

$$
\begin{equation*}
d G_{0}=-\frac{\partial L^{\prime}}{\partial t} d t=0 \tag{4.2.12}
\end{equation*}
$$

Explicitly

$$
\begin{align*}
d G_{0} & =\left[\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}\right)\left(\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}}\right)-\frac{1}{2} \dot{q}_{2} \ddot{q}_{2}\right. \\
& -\frac{1}{2}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)\left(\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}^{2} \frac{\partial^{2} q_{3}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}\right)  \tag{4.2.13}\\
& +\frac{1}{2} \dot{q}_{2}\left(\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{3}}{\partial q_{2}^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}\right)+\frac{1}{2} \ddot{q}_{2}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right) \\
& \left.+\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2} \dot{q_{2}}+2 q_{3}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)\right] d t=0 .
\end{align*}
$$

Replacing the expressions in the parentheses from Egs.(4.2.6) and (4.2.7) one gets

$$
\begin{equation*}
d G_{0}=\left(\dot{q}_{2} F_{1}\right) d t=0 \tag{4.2.14}
\end{equation*}
$$

where

$$
\begin{equation*}
F_{1}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}+\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}-\dot{q}_{2}+2 q_{3}+1=0 \tag{4.2.15}
\end{equation*}
$$

Since $F_{1}$ is not identically zero, we consider it as a new constraint. Thus for a valid theory, variation of $F_{1}$ should be zero. Thus one gets

$$
\begin{equation*}
d F_{1}=F_{2} d t=0 \tag{4.2.16}
\end{equation*}
$$

where

$$
\begin{equation*}
F_{2}=\dot{q}_{2}-3-2\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)=0 \tag{4.2.17}
\end{equation*}
$$

Again, since $F_{2}$ is not identically zero, it is an additional constraint and its variation is

$$
\begin{equation*}
d F_{2}=\left[\ddot{q}_{2}-2\left(\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+\dot{q}_{2}^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}}\right)\right] d t=0 \tag{4.2.18}
\end{equation*}
$$

But due to $\mathrm{Eq}(4.2 .6)$, the expression in parentheses is $\dot{q}_{2}-1$. So (4.2.18) leads us to the following differential equation for $q_{2}$ :

$$
\begin{equation*}
d F_{2}=\ddot{q}_{2}-2 \dot{q}_{2}+2, \tag{4.2.19}
\end{equation*}
$$

which has the following solution :

$$
\begin{equation*}
q_{2}(t)=2 A e^{2 t}+t+c_{1} . \tag{4.2.20}
\end{equation*}
$$

Besides, (4.1.2) for $\mu=2$ is

$$
\begin{equation*}
d G_{2}=\frac{\partial L}{\partial q_{2}} d t=-d t \tag{4.2.21}
\end{equation*}
$$

Hence,

$$
\begin{align*}
d G_{2}=[ & -\frac{1}{2} \ddot{q}_{2}+\frac{1}{2}\left(\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{3}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}\right)  \tag{4.2.22}\\
& \left.+\dot{q_{2}} \frac{\partial q_{1}}{\partial q_{2}}+\dot{q_{2}} \frac{\partial q_{3}}{\partial q_{2}}+\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{3}}{\partial t}\right] d t=-d t .
\end{align*}
$$

Again the expression in the inner parentheses is replaced by $4 q_{3}+\ddot{q}_{2}-2 \dot{q}_{2}$, from (4.2.7). Then (4.2.22) becomes

$$
\begin{equation*}
d G_{2}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q_{2}}+\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}-\dot{q}_{2}+2 q_{3}+1=F_{1} d t \tag{4.2.23}
\end{equation*}
$$

However, the constraint (4.2.23) is the same as (4.2.15). Thus, it does not give an additional constraint. Now, our problem is reduced to solving partial differential equations (4.2.6) and (4.2.7) with independent constraints (4.2.15) and (4.2.17). Making use of these constraints, one gets

$$
\begin{equation*}
\frac{\partial^{2} q_{1}}{\partial^{2} q_{2}}=0 \tag{4.2.24}
\end{equation*}
$$

which may have a solution in the form

$$
\begin{equation*}
q_{1}\left(t, q_{2}\right)=K(t) q_{2}+T(t) \tag{4.2.25}
\end{equation*}
$$

where $K(t)$ and $T(t)$ are functions to be determined. Some simple calculations lead us to the expressions

$$
\begin{equation*}
K(t)=\text { constant }, \quad T(t)=A e^{2 t}-B t+D . \tag{4.2.26}
\end{equation*}
$$

Since $q_{2}$ is determined as a function of t , the expression (4.2.25) can be written as

$$
\begin{equation*}
q_{1}(t)=Q^{\prime} e^{2 t}-t+D^{\prime} \tag{4.2.27}
\end{equation*}
$$

Applying the same procedure to the variable $q_{3}$ we arrive at the differential equation

$$
\begin{equation*}
\frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}=0 \tag{4.2.28}
\end{equation*}
$$

which has the general solution

$$
\begin{equation*}
q_{3}\left(t, q_{2}\right)=C_{1}(t)+C_{2}\left(q_{2}\right) . \tag{4.2.29}
\end{equation*}
$$

However, further calculations give

$$
\begin{equation*}
C_{1}=A^{\prime \prime} e^{-2 t}+B^{\prime} C_{2}\left(q_{2}\right)=0 \tag{4.2.30}
\end{equation*}
$$

Thus, we have

$$
\begin{equation*}
q_{3}(t)=A e^{2 t}+B e^{-2 t}+C . \tag{4.2.31}
\end{equation*}
$$

Eqs (4.2.27), and (4.2.31) are the solution of the system in the phase space $q_{1}, q_{2}$ and $q_{3}$.

### 4.2.2 Example Two

As a second example we consider the singular Lagrangian

$$
\begin{equation*}
L=\frac{1}{2} a_{i j}\left(t, q_{k}\right) \dot{q}_{i} \dot{q}_{j}+b_{i}\left(t, q_{k}\right) \dot{q}_{i}-c\left(t, q_{k}\right), \quad i, j, k=1, \ldots, 6, \tag{4.2.32}
\end{equation*}
$$

where $a_{i j}$ is a $6 \times 6$ symmetric matrix of rank 2 , with matrix elements

$$
\begin{equation*}
a_{11}=a_{22}=1, \tag{4.2.33}
\end{equation*}
$$

$$
\begin{gather*}
a_{12}=a_{21}=2,  \tag{4.2.34}\\
a_{1 \mu}=a_{\mu 1}=\alpha_{\mu}+2 \alpha_{\mu}^{\prime},  \tag{4.2.35}\\
a_{2 \mu}=a_{\mu 2}=2 \alpha_{\mu}+\alpha_{\mu}^{\prime},  \tag{4.2.36}\\
a_{\mu \nu}=a_{\nu \mu}=\alpha_{\mu} \alpha_{\nu}+2\left(\alpha_{\mu} \alpha_{\nu}^{\prime}+\alpha_{\mu}^{\prime} \alpha_{\nu}\right)+\alpha_{\nu}^{\prime} \alpha_{\mu}^{\prime}, \quad \mu, \nu=3,4,5,6 . \tag{4.2.37}
\end{gather*}
$$

Here $\alpha_{\mu}$ and $\quad \alpha_{\mu}^{\prime}$ are constants and the functions $b_{i}$ and $\quad c$ are

$$
\begin{gather*}
b_{1}=q_{2}+\alpha_{\mu}^{\prime} q_{\mu}  \tag{4.2.38}\\
b_{2}=q_{2}-q_{1}-\left(\alpha_{\mu}-\alpha^{\prime}\right) q_{\mu}  \tag{4.2.39}\\
b_{\mu}=\alpha_{\mu} b_{1}+\alpha_{\mu}^{\prime} b_{2}  \tag{4.2.40}\\
c=q_{1}-2 q_{2}+\left(\alpha_{\mu}-2 \alpha_{\mu}^{\prime}\right) q_{\mu} \tag{4.2.41}
\end{gather*}
$$

As in the previous example this system can be treated as a continuous system in the form

$$
\begin{equation*}
q_{1}=q_{1}\left(t, q_{\mu}\right), \quad \quad q_{2}=q_{2}\left(t, q_{\mu}\right) \tag{4.2.42}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
\dot{q}_{1}=\frac{d q_{1}}{d t}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{\mu}} \dot{q}_{\mu}, \quad \dot{q}_{2}=\frac{d q_{2}}{d t}=\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{\mu}} \dot{q}_{\mu} . \tag{4.2.43}
\end{equation*}
$$

Relation (4.2.43) can be replaced in (4.2.32) to obtain the following modified Lagrangian $L^{\prime}$ :

$$
\begin{align*}
L^{\prime}=\frac{1}{2} & \left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{\mu}} \dot{q}_{\mu}\right)^{2}+\frac{1}{2}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{\mu}} \dot{q}_{\mu}\right)^{2} \\
& +2\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{\mu}} \dot{q_{\mu}}\right)\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{\mu}} \dot{q}_{\mu}\right)+\left(\alpha_{\mu}+2 \alpha_{\mu}^{\prime}\right) \dot{q}_{\mu}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{\mu}} \dot{q}_{\mu}\right) \\
& +\left(2 \alpha_{\mu}+\alpha_{\mu}^{\prime}\right) \dot{q_{\mu}}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{\mu}} \dot{q_{\mu}}\right)+\left(q_{2}-q_{1}-\left(\alpha_{\mu}-\alpha_{\mu}^{\prime}\right) q_{\mu}\right)\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{\mu}} \dot{q}_{\mu}\right) \\
& +\frac{1}{2} a_{\mu \nu} \dot{q}_{\mu} \dot{q}_{\nu}+\left(q_{2}+\alpha_{\mu}^{\prime} q_{\mu}\right)\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{\mu}} \dot{q_{\mu}}\right)+\left(\alpha_{\mu}\left(q_{2}+\alpha_{\mu}^{\prime} q_{\mu}\right)\right. \\
& \left.+\alpha_{\mu}^{\prime}\left(q_{2}-q_{1}-\left(\alpha_{\mu}-\alpha_{\mu}^{\prime}\right) q_{\mu}\right)\right) \dot{q_{\mu}}-q_{1}+2 q_{2}-\left(\alpha_{\mu}-2 \alpha_{\mu}^{\prime}\right) q_{\mu} . \tag{4.2.44}
\end{align*}
$$

The Euler-Lagrange equations (4.1.1) read as

$$
\begin{align*}
\frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{1}\right)}\right)+\frac{\partial}{\partial q_{\mu}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{\mu} q_{1}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{1}} & =0 \\
\text { and } \quad \frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{2}\right)}\right)+\frac{\partial}{\partial q_{\mu}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{\mu} q_{2}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{2}} & =0 \tag{4.2.45}
\end{align*}
$$

where $x_{0} \equiv t, \quad x_{\mu} \equiv q_{\mu}, \quad \mu=3,4,5,6 .$,
More explicitly, the equations of motion (4.2.45) are

$$
\begin{array}{r}
\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{\mu} \frac{\partial^{2} q_{1}}{\partial t \partial q_{\mu}}+\dot{q}_{\mu} \frac{\partial^{2} q_{1}}{\partial q_{\mu} \partial q_{\nu}} \dot{q}_{\nu}+2 \frac{\partial^{2} q_{2}}{\partial t^{2}}+4 \dot{q}_{\mu} \frac{\partial^{2} q_{2}}{\partial t \partial q_{\mu}} \\
+2 \dot{q}_{\mu} \frac{\partial^{2} q_{2}}{\partial q_{\mu} \partial q_{\nu}} \dot{q}_{\nu}+2 \frac{\partial q_{2}}{\partial t}+\ddot{q_{\mu}} \frac{\partial q_{1}}{\partial q_{\mu}}+\left(2 \ddot{q}_{\mu}+2 \dot{q}_{\mu}\right) \frac{\partial q_{2}}{\partial q_{\mu}}  \tag{4.2.46}\\
+\left(\alpha_{\mu}+2 \alpha_{\mu}^{\prime}\right) \ddot{q}_{\mu}+2 \alpha_{\mu}^{\prime} \dot{q}_{\mu}+1=0 .
\end{array}
$$

and

$$
\begin{array}{r}
2 \frac{\partial^{2} q_{1}}{\partial t^{2}}+4 \dot{q}_{\mu} \frac{\partial^{2} q_{1}}{\partial t \partial q_{\mu}}+2 \dot{q}_{\mu} \frac{\partial^{2} q_{1}}{\partial q_{\mu} \partial q_{\nu}} \dot{q}_{\nu}+\frac{\partial^{2} q_{2}}{\partial t^{2}}+2 \dot{q}_{\mu} \frac{\partial^{2} q_{2}}{\partial t \partial q_{\mu}} \\
+\dot{q}_{\mu} \frac{\partial^{2} q_{2}}{\partial q_{\mu} \partial q_{\nu}} \dot{q}_{\nu}-2 \frac{\partial q_{1}}{\partial t}+\ddot{q}_{\mu} \frac{\partial q_{2}}{\partial q_{\mu}}+\left(2 \ddot{q}_{\mu}-2 \dot{q}_{\mu}\right) \frac{\partial q_{1}}{\partial q_{\mu}}  \tag{4.2.47}\\
\\
+\left(2 \alpha_{\mu}+2 \alpha_{\mu}^{\prime}\right) \ddot{q}_{\mu}-2 \alpha_{\mu} \dot{q}_{\mu}-2=0 .
\end{array}
$$

To solve the equations of motion, let us consider the following transformation: $q_{\mu} \longrightarrow \mu_{\mu}, \quad t \longrightarrow t$.

Hence, Eqs. (4.2.46) and (4.2.47) are reduced to the following canonical forms (9)

$$
\begin{equation*}
\frac{\partial^{2} q_{1}}{\partial t^{2}}=-2 \frac{\partial^{2} q_{2}}{\partial t^{2}}-2 \frac{\partial q_{2}}{\partial t}-1 \tag{4.2.48}
\end{equation*}
$$

and

$$
\begin{equation*}
2 \frac{\partial^{2} q_{1}}{\partial t^{2}}=-\frac{\partial^{2} q_{2}}{\partial t^{2}}+2 \frac{\partial q_{1}}{\partial t}+2 \tag{4.2.49}
\end{equation*}
$$

Integrating (4.2.48) we get

$$
\begin{equation*}
\frac{\partial q_{1}}{\partial t}=-2 \frac{\partial q_{2}}{\partial t}-2 q_{2}-t \tag{4.2.50}
\end{equation*}
$$

Substituting (4.2.48) and (4.2.50) in (4.2.49) we obtain

$$
\begin{equation*}
\frac{\partial^{2} q_{2}}{\partial t^{2}}-\frac{4}{3} q_{2}=-\frac{4}{3}+\frac{2}{3} t . \tag{4.2.51}
\end{equation*}
$$

The general solution of the homogeneous equation is

$$
\begin{equation*}
q_{2}(t)=A e^{2 t \sqrt{3}}+B e^{-2 t \sqrt{3}} . \tag{4.2.52}
\end{equation*}
$$

Choosing a particular solution as

$$
\begin{equation*}
q_{2}^{\text {part }}=-\frac{1}{2} t+1, \tag{4.2.53}
\end{equation*}
$$

and inserting it in (4.2.50) , we obtain

$$
\begin{equation*}
q_{2}(t)=A e^{2 t \sqrt{3}}+B e^{-2 t \sqrt{3}}-\frac{1}{2} t+1 \tag{4.2.54}
\end{equation*}
$$

Substituting (4.2.54) in (4.2.50) and integrating the result, $q_{1}$ is determined as

$$
\begin{equation*}
q_{1}(t)=K e^{2 t \sqrt{3}}+L e^{-2 t \sqrt{3}}-M t . \tag{4.2.55}
\end{equation*}
$$

### 4.2.3 Example Three

Let us consider the singular Lagrangian discussed in chapter two and chapter three

$$
\begin{equation*}
L=\frac{m}{2}\left(\dot{q}_{1}^{2}+{\dot{q_{2}}}^{2}+l^{2} \dot{q}_{3}^{2}+2 l \dot{q}_{1} \dot{q}_{3} \cos q_{3}+2 l \dot{q}_{2} \dot{q}_{3} \sin q_{3}\right)+V\left(q_{1}, q_{2}, q_{3}\right) \tag{4.2.56}
\end{equation*}
$$

where $l$ and $m$ are constants. Since the rank of the Hess matrix is two, this system can be be treated as a continuous system in the form

$$
\begin{equation*}
q_{1}=q_{1}\left(t, q_{3}\right), \quad q_{2}=q_{2}\left(t, q_{3}\right), \tag{4.2.57}
\end{equation*}
$$

Now, let us write $\dot{q}_{1}$ and $\dot{q}_{3}$ as

$$
\begin{equation*}
\dot{q}_{1}=\frac{d q_{1}}{d t}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q}_{3}, \quad \quad \dot{q}_{2}=\frac{d q_{2}}{d t}=\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3} . \tag{4.2.58}
\end{equation*}
$$

Relations (4.2.58) can be replaced in (4.2.56) to obtain the following "modified Lagrangian" $L^{\prime}$ as

$$
\begin{align*}
L^{\prime} & =\frac{m}{2}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q}_{3}\right)^{2}+\frac{m}{2}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3}\right)^{2}+\frac{m}{2} l^{2} \dot{q}_{3}{ }^{2} \\
& +m l \dot{q_{3}} \cos q_{3}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q_{3}}\right)+m l \dot{q_{3}} \sin q_{3}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3}\right)+V^{\prime}\left(q_{1}, q_{2}\right) \tag{4.2.59}
\end{align*}
$$

The Euler-Lagrange equations (4.1.1) read as

$$
\begin{align*}
\frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{1}\right)}\right)+\frac{\partial}{\partial q_{3}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{3} q_{1}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{1}} & =0,  \tag{4.2.60}\\
\text { and } \quad \frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{2}\right)}\right)+\frac{\partial}{\partial q_{3}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{3} q_{2}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{2}} & =0 .
\end{align*}
$$

where $x_{0} \equiv t, \quad x_{2} \equiv q_{2}$. More explicitly, the second order partial differential equations are

$$
\begin{align*}
m \frac{\partial^{2} q_{1}}{\partial t^{2}} & +2 m \dot{q}_{3} \frac{\partial^{2} q_{1}}{\partial t \partial q_{3}}+m \dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+m \ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{3}}  \tag{4.2.61}\\
& +m l \ddot{q}_{3} \cos q_{3}-2 m l \dot{q}_{3} \sin q_{3}-V_{, 1}=0,
\end{align*}
$$

and

$$
\begin{array}{r}
m \frac{\partial^{2} q_{2}}{\partial t^{2}}+2 m \dot{q}_{3} \frac{\partial^{2} q_{2}}{\partial t \partial q_{3}}+m \dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{2}}{\partial q_{2}{ }^{2}}+m \ddot{q}_{2} \frac{\partial q_{2}}{\partial q_{3}}  \tag{4.2.62}\\
\quad+m l \ddot{q}_{3} \sin q_{3}+2 m l \dot{q}_{3}{ }^{2} \cos q_{3}-V_{, 2}=0,
\end{array}
$$

The usual Hamiltonian $H_{0}$ and the constraint relation $H_{3}$ are given as

$$
\begin{equation*}
H_{0}^{\prime}=p_{0}+\frac{p_{1}^{2}+p_{2}^{2}}{2 m}-V=0, \tag{4.2.63}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{3}=-m l^{2} \dot{q}_{3}^{2}-m l\left(\dot{q}_{1} \cos q_{3}+\dot{q_{2}} \sin q_{3}\right) . \tag{4.2.64}
\end{equation*}
$$

Hence, $G_{0}$ and $G_{3}$ defined in (4.1.2) respecify are

$$
\begin{align*}
G_{0}= & H_{0}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \\
= & \frac{m}{2}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q}_{3}\right)^{2}+\frac{m}{2}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3}\right)^{2}+\frac{m}{2} l^{2} \dot{q}_{3}^{2}  \tag{4.2.65}\\
& +m l \dot{q}_{3} \cos q_{3}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q}_{3}\right)+m l \dot{q}_{3} \sin q_{3}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3}\right) \\
& \quad+V^{\prime}\left(q_{1}, q_{2}\right),
\end{align*}
$$

and

$$
\begin{align*}
G_{3} & =H_{3}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right)  \tag{4.2.66}\\
& =-m l^{2} \dot{q}_{3}-m l \cos q_{3}\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q_{3}}\right)-m l \cos q_{3}\left(\frac{\partial q_{2}}{\partial t}+\frac{\partial q_{2}}{\partial q_{3}} \dot{q}_{3}\right) .
\end{align*}
$$

Now we are ready to test whether (4.1.2) are satisfied or not. In Fact (4.1.2) for $\mu=0$ is

$$
\begin{equation*}
d G_{0}=-\frac{\partial L^{\prime}}{\partial t} d t=0 \tag{4.2.67}
\end{equation*}
$$

Explicitly,

$$
\begin{equation*}
m l \dot{q}_{1} \dot{q}_{3}{ }^{2} \sin q_{3}-m l \dot{q_{2}}{\dot{q_{3}}}^{2} \cos q_{3}+l \dot{q_{3}} \cos q_{3} V_{, 1}+l \dot{q_{3}} \sin q_{3} V_{, 2}=0 . \tag{4.2.68}
\end{equation*}
$$

Equation (4.1.2) for $\mu=3$ is

$$
\begin{equation*}
\frac{d G_{3}}{d t}=0 \tag{4.2.69}
\end{equation*}
$$

Explicitly,

$$
\begin{equation*}
m l \dot{q}_{1} \dot{q}_{3}^{2} \sin q_{3}-m l \dot{q}_{2} \dot{q}_{3}^{2} \cos q_{3}-l \dot{q_{3}} \cos q_{3} V_{, 1}-l \dot{q_{3}} \sin q_{3} V_{, 2}=0, \tag{4.2.70}
\end{equation*}
$$

Subtracting equation (4.2.68) and (4.2.70) we get

$$
\begin{equation*}
2 l \dot{q}_{3} \cos q_{3} V_{, 1}+2 l \dot{q}_{3} \sin q_{3} V_{, 2}=0 \tag{4.2.71}
\end{equation*}
$$

Choosing $V=q_{1} \cos q_{3}+q_{2} \sin q_{3}$ equation (4.2.71) becomes

$$
\begin{equation*}
2 l \dot{q}_{3}\left(\cos ^{2} q_{3}+\sin ^{2} q_{3}\right)=0 \tag{4.2.72}
\end{equation*}
$$

The solution of equation (4.2.72) is

$$
\begin{equation*}
q_{3}=\text { constant } \tag{4.2.73}
\end{equation*}
$$

Substituting Eq (4.2.73) in Eq (4.2.61), we get

$$
\begin{equation*}
m \frac{\partial^{2} q_{1}}{\partial^{2} t}-V_{, 1}=0 \tag{4.2.74}
\end{equation*}
$$

The solution of Eq. (4.2.74) is obtain as

$$
\begin{equation*}
q_{1}=\frac{t^{2}}{2 m}+A t+B \tag{4.2.75}
\end{equation*}
$$

where A and B are constants.
Substituting Eq. (4.2.75) in Eq. (4.2.62), becomes

$$
\begin{equation*}
q_{2}=\frac{t^{2}}{2 m}+C t+D \tag{4.2.76}
\end{equation*}
$$

where C and D are constants.

### 4.2.4 Example Four

As a last example let us consider the Mittelstaedt's Lagrangian model (20), which is given as

$$
\begin{equation*}
L=\frac{1}{2 m}\left(\dot{q}_{1}+\dot{q}_{2}\right)^{2}+\frac{1}{2 \mu} \dot{q}_{3}^{2}+V\left(q_{1}, q_{2}, q_{3}\right) \tag{4.2.77}
\end{equation*}
$$

This Lagrangian is singular since the Hess matrix is of rank two.
This system can be be treated as a continuous system in the form

$$
\begin{align*}
& q_{1}=q_{1}\left(t, q_{2}\right), \quad q_{3}=q_{3}\left(t, q_{2}\right),  \tag{4.2.78}\\
& \dot{q}_{1}=\frac{d q_{1}}{d t}=\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}, \quad \dot{q}_{3}=\frac{d q_{3}}{d t}=\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2} . \tag{4.2.79}
\end{align*}
$$

Relations (4.2.79) with can be replaced in (4.2.77) to obtain following the " modified Lagrangian" $L^{\prime}$ as:

$$
\begin{equation*}
L^{\prime}=\frac{1}{2 m}\left[\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{3}} \dot{q}_{3}\right)+\dot{q}_{2}\right]^{2}+\frac{1}{2 \mu}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)^{2}+V\left(q_{1}, q_{3}\right) \tag{4.2.80}
\end{equation*}
$$

The Euler-Lagrange equations (4.1)read as

$$
\begin{array}{rlrl}
\frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{1}\right)}\right)+\frac{\partial}{\partial q_{2}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{2} q_{1}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{1}} & =0,  \tag{4.2.81}\\
\text { and } \quad & \frac{\partial}{\partial t}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{0} q_{3}\right)}\right)+\frac{\partial}{\partial q_{2}}\left(\frac{\partial L^{\prime}}{\partial\left(\partial_{2} q_{3}\right)}\right)-\frac{\partial L^{\prime}}{\partial q_{3}} & =0 .
\end{array}
$$

where $x_{0} \equiv t, \quad x_{2} \equiv q_{2}$.
More explicitly, the second order partial differential equations are

$$
\begin{array}{r}
\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}}-\ddot{q}_{2}-m V_{1}=0 \\
\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}-\mu V_{, 1}=0 \tag{4.2.83}
\end{array}
$$

The quantities $H_{0}$ and $H_{3}$ are

$$
\begin{equation*}
H_{0}=\frac{m}{2} p_{1}^{2}+\frac{\mu}{2} p_{3}^{2}-V\left(q_{1}, q_{3}\right), \tag{4.2.84}
\end{equation*}
$$

$$
\begin{equation*}
H_{3}=p_{3}=\frac{1}{\mu} \dot{q_{3}} . \tag{4.2.85}
\end{equation*}
$$

Hence,

$$
\left.\left.\begin{array}{l}
G_{0}=H_{0}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \\
=\frac{1}{2 m}\left[\left(\frac{\partial q_{1}}{\partial t}+\frac{\partial q_{1}}{\partial q_{2}} \dot{q}_{2}\right)\right.
\end{array}\right)+\dot{q}_{2}\right]^{2}+\frac{1}{2 \mu}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)^{2}-V\left(q_{1}, q_{3}\right), ~ \begin{aligned}
G_{3} & =H_{3}\left(q_{a}, x_{\mu}, p_{a}=\frac{\partial L}{\partial \dot{q}_{a}}\right) \\
& =\frac{1}{\mu}\left(\frac{\partial q_{3}}{\partial t}+\frac{\partial q_{3}}{\partial q_{2}} \dot{q}_{2}\right)
\end{aligned}
$$

Now we are ready to test whether (4.1.2) are satisfied or not . In fact (4.1.2) for $\mu=0$ is

$$
\begin{gather*}
d G_{0}=-\frac{\partial L^{\prime}}{\partial t} d t=0  \tag{4.2.88}\\
d G_{0}=\left(\dot{q_{2}} \frac{\partial V}{\partial q_{1}}\right) d t=0 \tag{4.2.89}
\end{gather*}
$$

for $\mu=3$ is

$$
\begin{equation*}
d G_{3}=-\frac{\partial L^{\prime}}{\partial q_{3}} d t=\left(\frac{\partial V}{\partial q_{3}}\right) d t=0 \tag{4.2.90}
\end{equation*}
$$

The Eqs. (4.2.82) and (4.2.83) becomes,

$$
\begin{gather*}
\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+{\dot{q_{2}}}^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}}+\ddot{q}_{2}=0  \tag{4.2.91}\\
\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}{ }^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}}=0 \tag{4.2.92}
\end{gather*}
$$

Eqs. (4.2.91) and (4.2.92) are second order partial differential equations. From equations (4.2.79) the second derivatives of $q_{1}$ and $q_{3}$ are

$$
\begin{align*}
& \ddot{q}_{1}=\frac{\partial^{2} q_{1}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{1}}{\partial t \partial q_{2}}+\dot{q}_{2}{ }^{2} \frac{\partial^{2} q_{1}}{\partial q_{2}^{2}}+\ddot{q}_{2} \frac{\partial q_{1}}{\partial q_{2}},  \tag{4.2.93}\\
& \ddot{q}_{3}=\frac{\partial^{2} q_{3}}{\partial t^{2}}+2 \dot{q}_{2} \frac{\partial^{2} q_{3}}{\partial t \partial q_{2}}+\dot{q}_{2}^{2} \frac{\partial^{2} q_{3}}{\partial q_{2}^{2}}+\ddot{q}_{2} \frac{\partial q_{3}}{\partial q_{2}} . \tag{4.2.94}
\end{align*}
$$

Therefore Eq. (4.2.91) becomes

$$
\begin{equation*}
\ddot{q}_{1}+\ddot{q}_{2}=0 . \tag{4.2.95}
\end{equation*}
$$

The solution of Eq. (4.2.95) is given as

$$
\begin{equation*}
q_{1}(t)=-q_{2}+A t+B, \tag{4.2.96}
\end{equation*}
$$

where A and B are constants
The Eq. (4.2.92) becomes

$$
\begin{equation*}
\ddot{q}_{3}=0, \tag{4.2.97}
\end{equation*}
$$

with solution given as

$$
\begin{equation*}
q_{3}(t)=C t+D . \tag{4.2.98}
\end{equation*}
$$

where C and D are constants.

## Chapter 5

## Conclusion

## Chapter 5

## Conclusion

The Hamiltonian and Lagrangian formulations of singular Lagrangian systems are used to investigate some models of physical systems study to compare these techniques of these formulation.

In the Hamiltonian formulation both Dirac's method and Hamilton-Jacobi method (Güler's method) are used. In the Lagrangian formulation. The technique of treatment the singular Lagrangian as field (continuous) system was used. Besides, the Hamilton-Jacobi method is unified with Lagrangian formulation.

In Dirac's method, one introduces primary constraints to construct the total Hamiltonian, which consists of the primary constraints multiplied by the Lagrange multipliers added to canonical (usual) Hamiltonian.

The first - class constraints have vanishing poisson brackets. The equation of motion are obtained as total derivatives interms of poisson brackets.

In Hamilton-Jacobi formulation, which developed by Güler, the equations of motion are written as total differential equations in many variables. These equations must satisfy the integrability conditions. If the integrability conditions are not identically satisfied, then these will be continued until we obtain a complete
system. Three models of physical system are discussed using these two method. The result are in exact agreement. In Hamilton-Jacobi method, we did not need to introduce an unknown multipliers as in Dirac's method.

In chapter four, the same physical models are discussed as field (continuous) systems in Lagrangian formulation. We mixed both Lagrangian formulation and Hamilton-Jacobi method to obtain a solvable partial differential equations of second order. The Euler - Lagrange equations of motion for field system are used to obtain the equations of motion.

Simultaneous solution of Euler - Lagrange equations with the constraints equations gives us the solutions of the dynamical systems.

These constraints equations are obtained from Hamilton-Jacobi approach.
These solutions satisfied the equations of motion that obtained in both Dirac's method and Hamilton-Jacobi method.

In fact, this comparison study needs more applications in physical systems in classical mechanics and field theory.

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